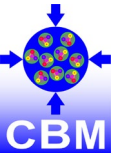


# The status of the Compressed Baryonic Matter experiment at FAIR

Andrej Kugler for the CBM Collaboration  
Nuclear Physics Institute Czech Academy of Sciences, Rez

Compact Stars in the QCD Phase Diagram, October 7-11, 2024, Kyoto

# Road map



What are we pursuing and why?



Physics motivation



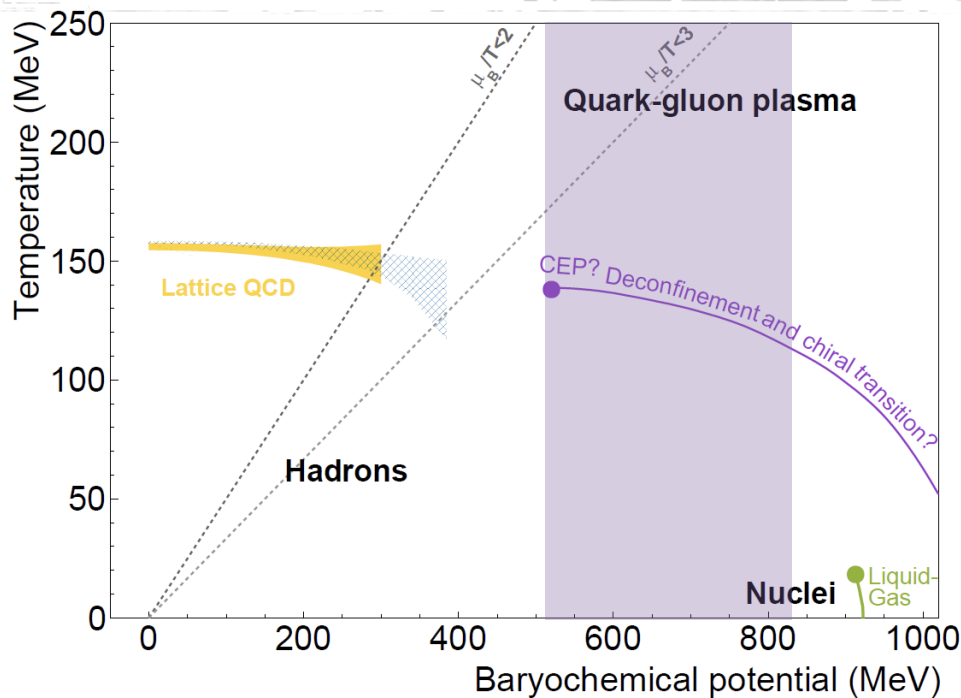
# QCD phase diagram

## Low $\mu_B$ , high $T$ :

- **Cross-over** transition from hadronic to quark matter - comprehensive studies of **QGP** properties
- No **critical point** anticipated for  $\mu_B/T < 3$  (LQCD)

## High $\mu_B$ , low $T$ :

- Unknown **phase structure** (first-order phase transition, critical point possible, mixed phases, new phases, ...)
- Properties of matter to determine
- Characteristics of hadrons
- Equation of State (**EoS**) to establish
- Neutron Star (**NS**)

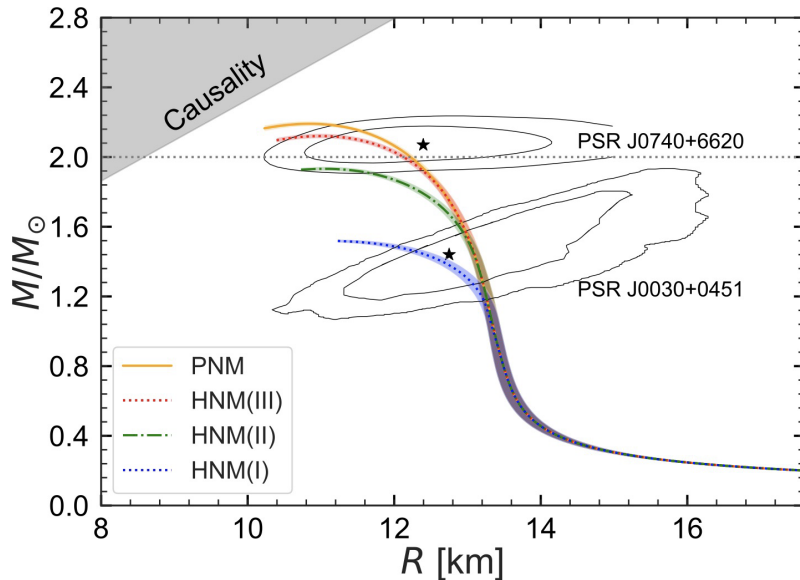
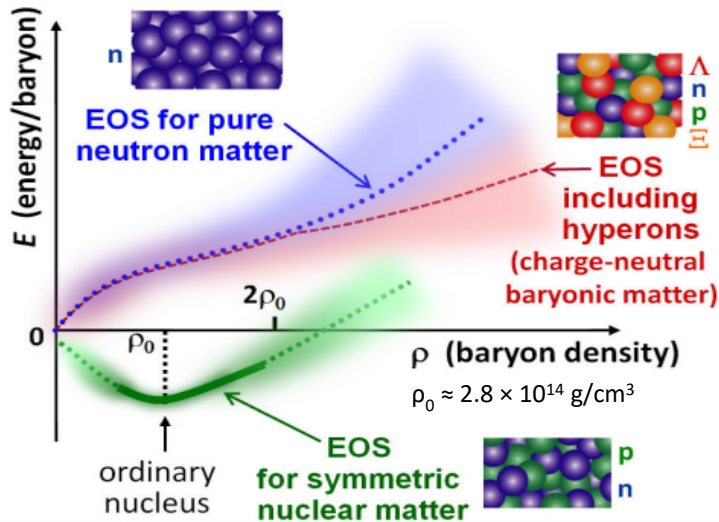


Bazavovet et al. [HotQCD], PLB 795 (2019) 15-21  
 Dinget al., [HotQCD], PRL 123 (2019) 6, 062002  
 Borsanyi et al., PRL 125(2020)5, 052001  
 Isserstedt et al. PRD 100 (2019) 074011  
 Gao, Pawłowski, PLB 820 (2021) 136584

# Neutron star (NS) puzzle

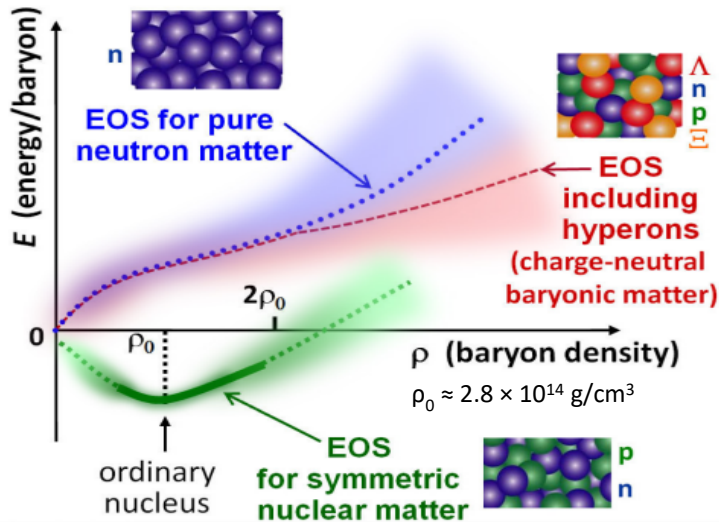
H. Tamura, JPS Conf. Proc., 011003 (2014)

„To establish the EoS applicable to the neutron star has been one of the most important subjects in nuclear physics for a long time but has not been achieved yet.” T. Hamura

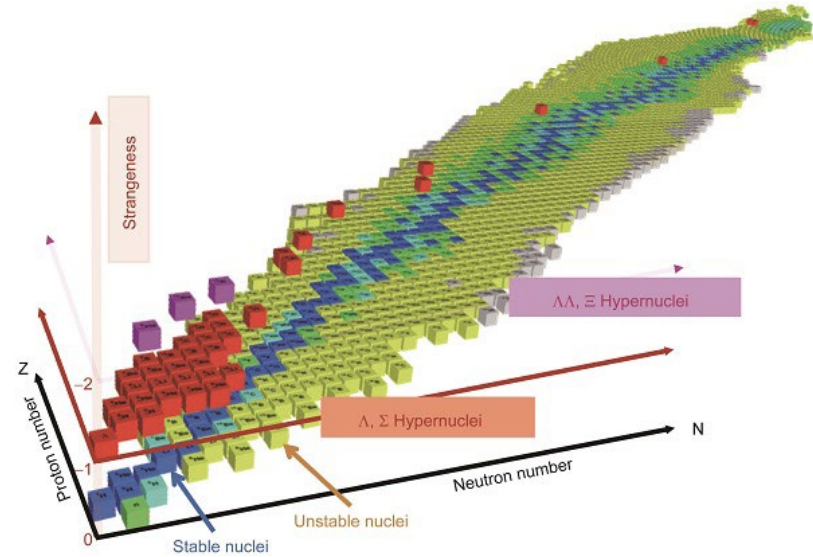


# Neutron star (NS) puzzle

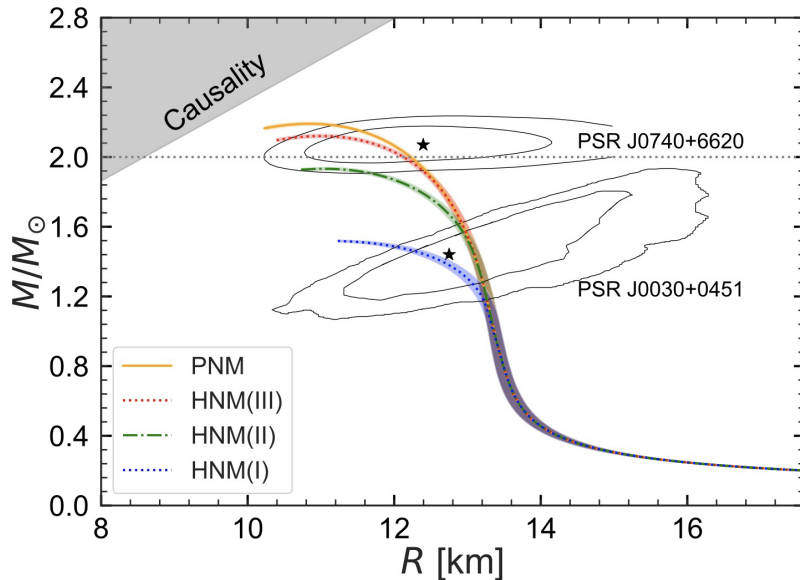
H. Tamura, JPS Conf. Proc., 011003 (2014)



„To establish the EoS applicable to the neutron star has been one of the most important subjects in nuclear physics for a long time but has not been achieved yet.” T. Hamura



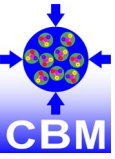
M. Kaneta, Department of Physics, Tohoku University, Japan



## Hypernuclei are pivotal for the EoS of the NS

- How do nuclei and hyper-nuclei form?
- What are their characteristics?
- How do nuclei (N) and hyperons (Y) interact?

# Road map



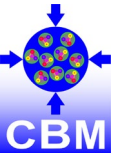
What are we pursuing and why?

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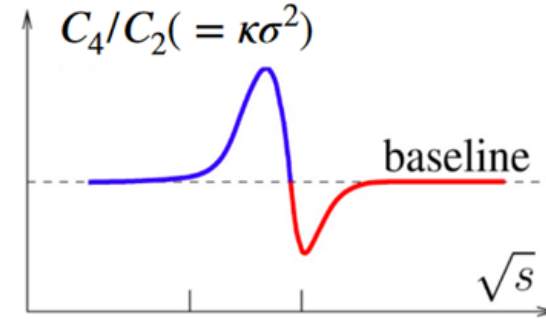
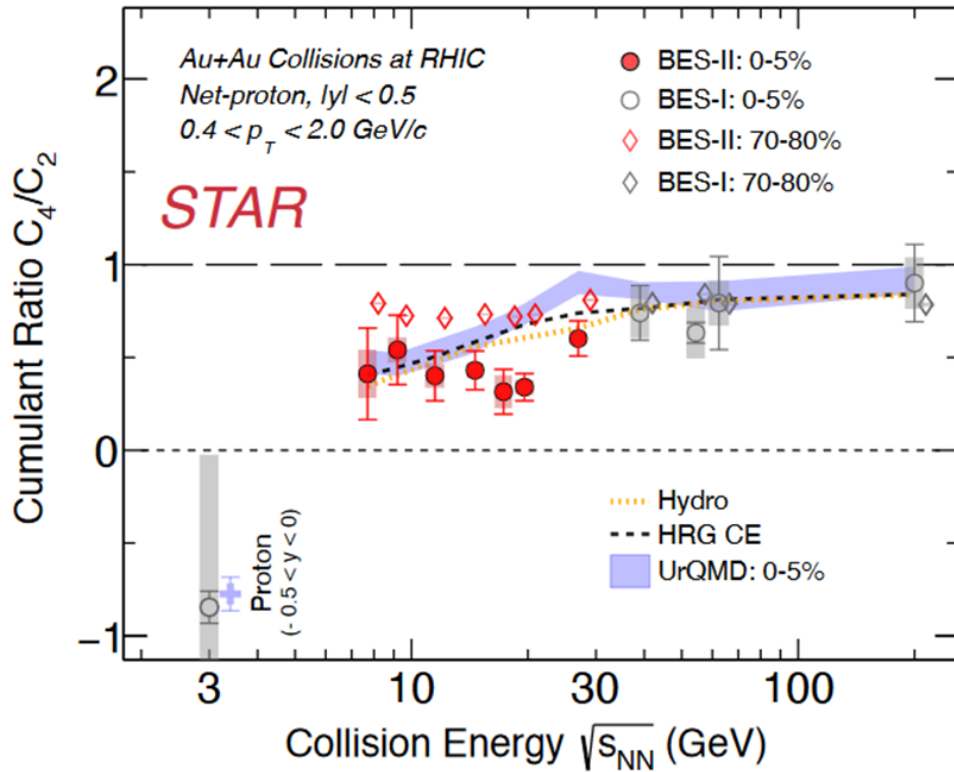
Where are we now?

Current status





# Search for signatures of Critical End Point: Fluctuations?



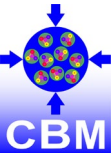
M. A. Stephanov, PRL 107 (2011) 052301

“..... Usual caveats apply: other (nontrivial) contributions to moments which do not behave singularly at the critical point can turn out to be relatively large. These include initial geometry fluctuations, jets, and other nonequilibrium effects. In addition, charge conservation effects may impose constraints on certain observables, such as total charge fluctuations. It is beyond the scope of this Letter to estimate these effects.....”

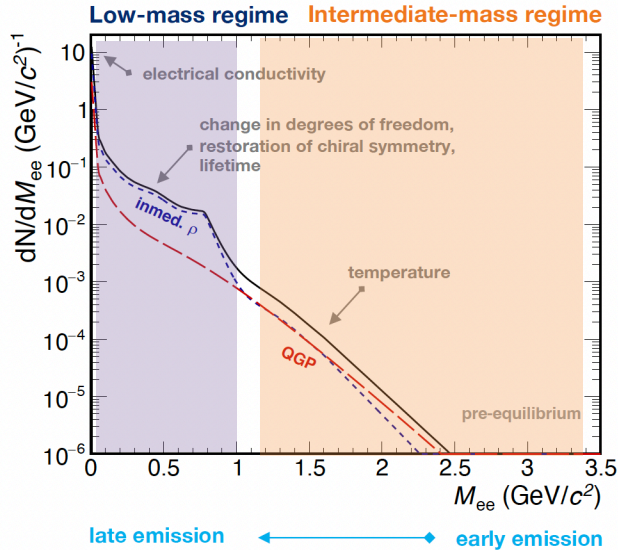
→ Correction of reaction volume fluctuations using mixed events or pion multiplicites, see [arXiv:2403.03598](https://arxiv.org/abs/2403.03598)  
 R. Holzmann, A. Rustamov, V. Koch, J. Stroth

A. Pandav for the STAR Collaboration at CPOD 2024

**Higher-order moments requires prominent statistics**  
**Detailed systematics studies indispensable**

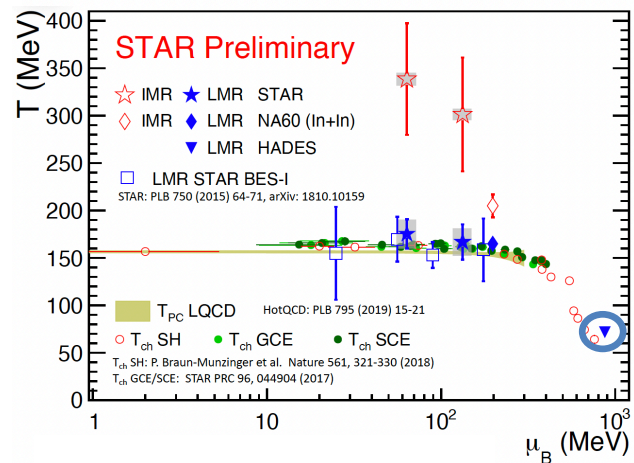


# E-M probes access the whole collision



Inscribes matter properties enabling estimation:

- degrees of freedom of the medium
- fireball's lifetime, temperature, acceleration, polarization
- transport properties
- restoration of chiral symmetry



Thermal dileptons in **LMR**:

- $T$  close to  $T_{ch}$  and  $T_{pc}$
- dominantly emitted around phase transition

Thermal dileptons in **IMR**:

- $T$  is higher than  $T_{pc}$
- Emitted from QGP phase

Effective size-signal:  $S_{eff} \sim R \frac{S}{B}$

$R$  - interaction rate

$S$  - signal

$B$  - combinatorial background

**Prominent interaction rate mandatory**

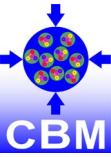
EPJC (2009) 59 607-623

Nature Physics 15, 1040-1045 (2019)

JPS Conf.Proc. 21 (2020) 010079

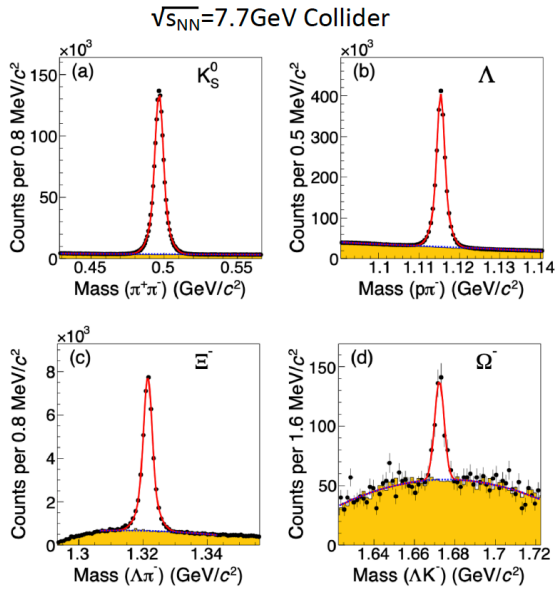


# Flow of strange particles

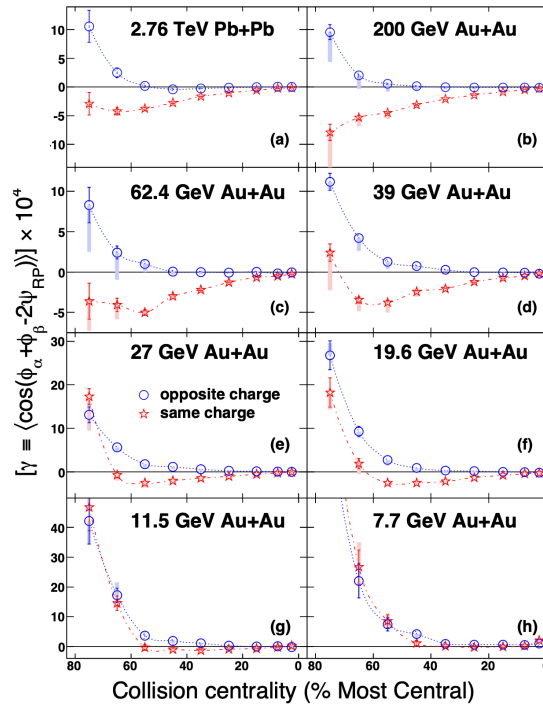


EoS investigations include vast number of measurements:

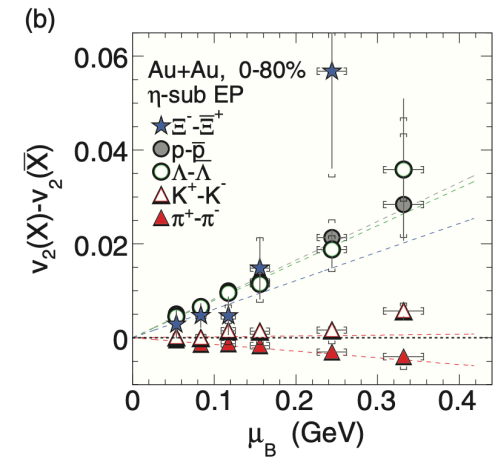
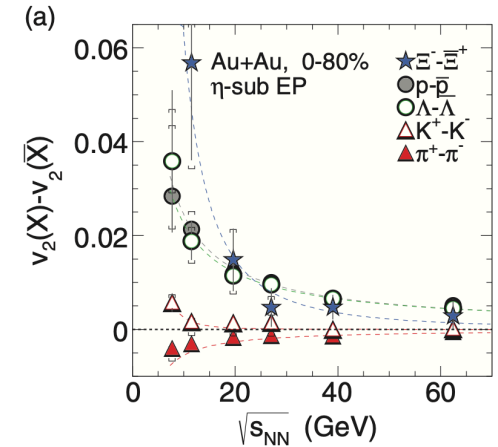
- Chemistry (strangeness, charm, hyper nuclei, ...)
- Collectivity
- Vorticity
- Fluctuations and correlations
- Interactions in the final states (NN, NY, YY, many-body, hyper-nuclei, ...)



PRC 102 (2020) 34909 (STAR)

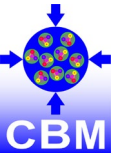


PRL 113 (2014) 52302



PRC 93 (2016) 14907 (STAR)

# Road map



What are we pursuing and why?

Physics motivation

Where are we now?

Current status

Who is involved?

Experimental facilities

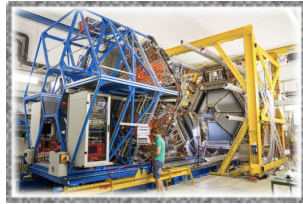


# High $\mu_B$ facilities

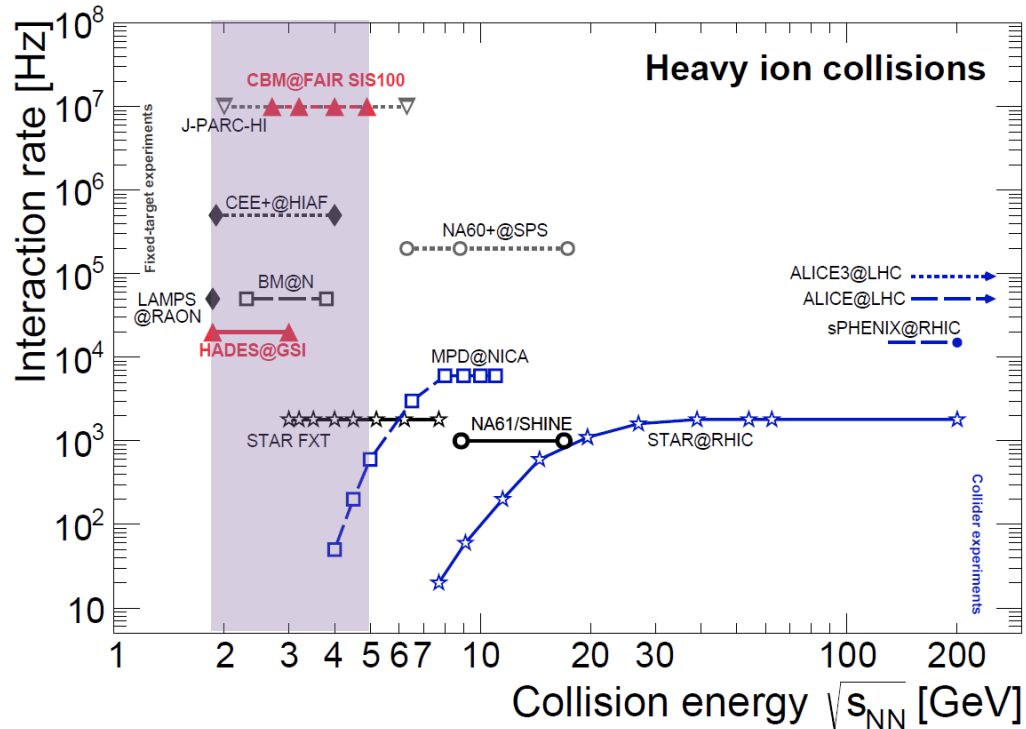
NA61/SHINE@SPS



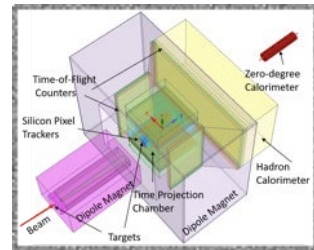
HADES@SIS18



STAR@RHIC



J-PARC-HI



T. Galatyuk, NPA 982 (2019), update 2024

[https://github.com/tgalatyuk/interaction\\_rate\\_facilities](https://github.com/tgalatyuk/interaction_rate_facilities)

CBM, EPJ A 53 2 (2017) 60

# High $\mu_B$ facilities

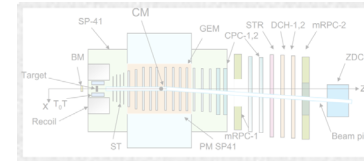
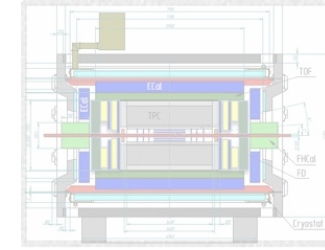
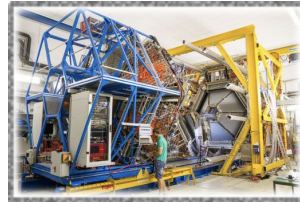
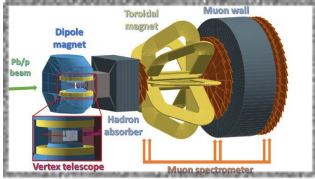
CBM / HADES@ SIS100 (>2028)

MPD, MB@N@NICA

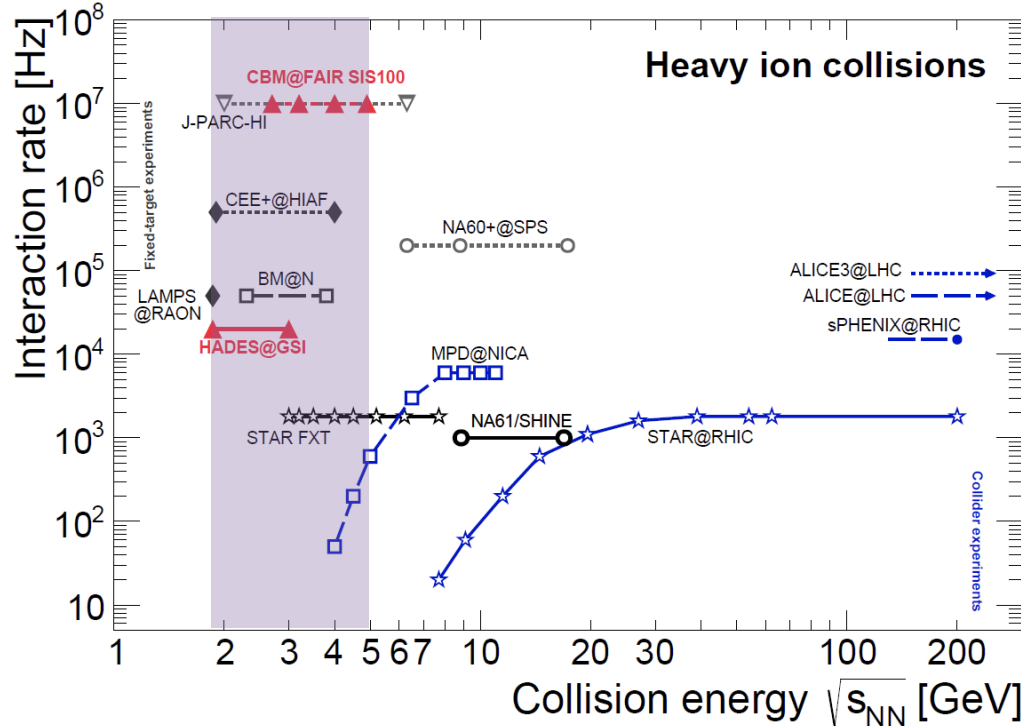
NA60@SPS(>2030)

NA61/SHINE@SPS

HADES@SIS18



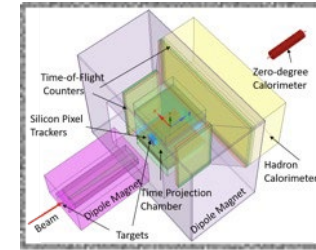
STAR@RHIC



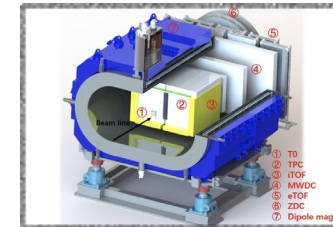
CBM / HADES:  
operations  
at  $\sqrt{s_{NN}} \sim 2 - 5$   
GeV

T. Galatyuk, NPA 982 (2019), update 2024  
[https://github.com/tgalatyuk/interaction\\_rate\\_facilities](https://github.com/tgalatyuk/interaction_rate_facilities)

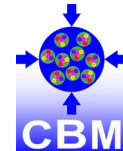
J-PARC-HI



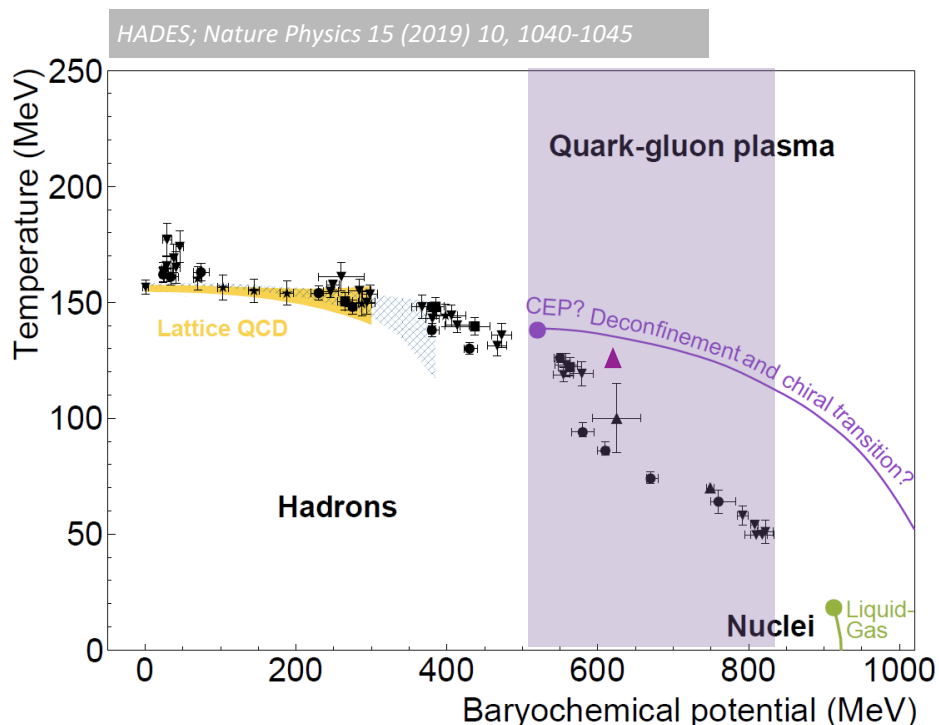
CEE@HIAF (>2027)



# Coverage of the QCD phase diagram



CBM / HADES @ SIS100 experimental exploration of the region  $\mu_B \sim 520 - 830 \text{ MeV}$



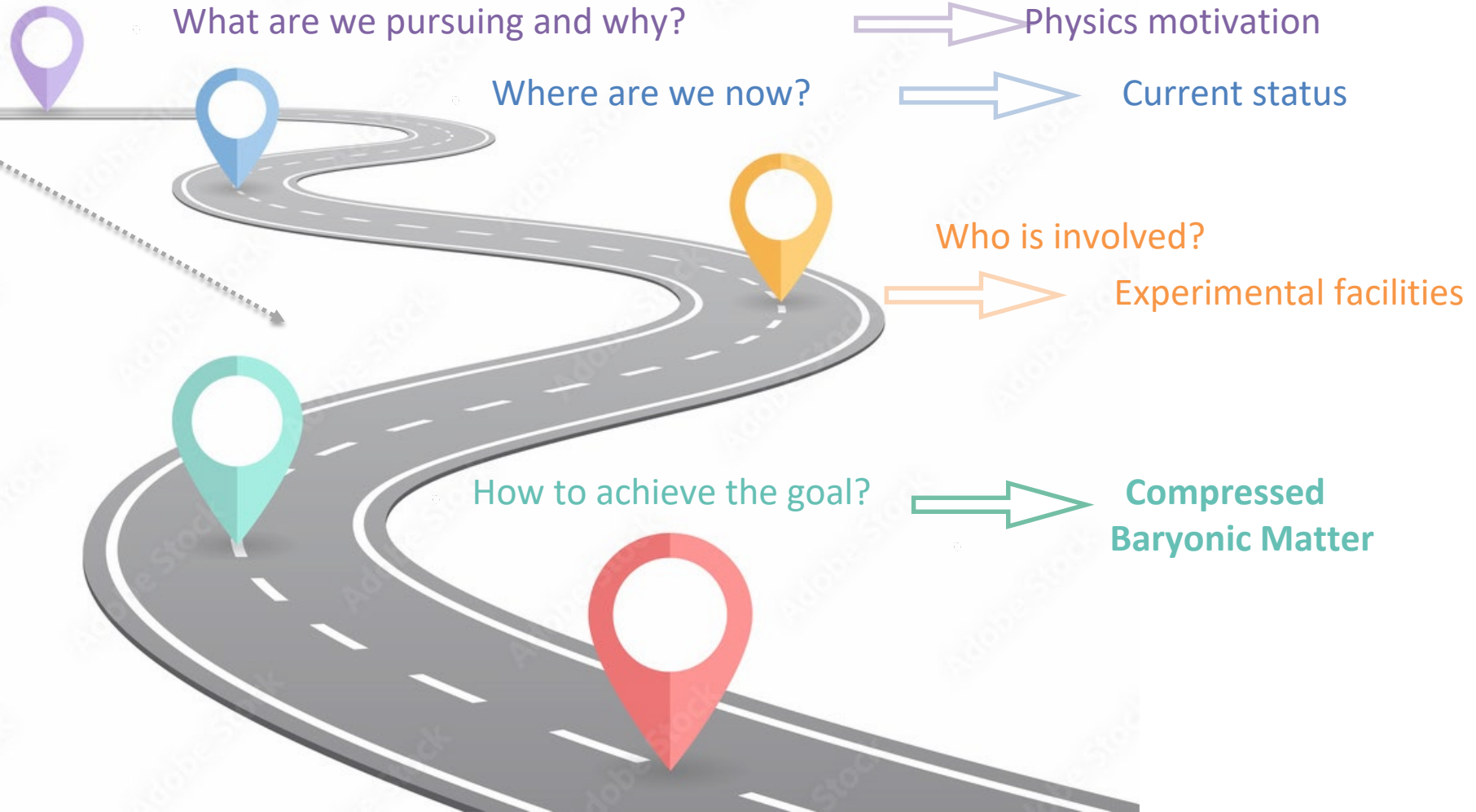
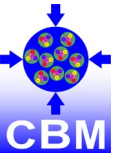
	$\sqrt{s_{NN}}$ (GeV)	$\mu_B$ (MeV)
<b>HADES@SIS18</b>	<b>2-2.5</b>	<b>830-760</b>
<b>CBM@SIS100</b>	<b>2.3-5.3</b>	<b>785-520</b>
NA61/SHINE@SPS	5.1-17.3	530-220
STAR-COLL@RHIC	7.7-200	400-22
STAR-FXT@RHIC	3-13.7	700-265

A. Andronic, P. Braun-Munzinger, K. Redlich and B. J. Stachel, *Nature* 561, no. 7723, 321 (2018)

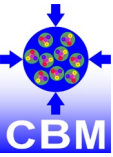
Bazavov et al. [*HotQCD*], *PLB* 795 (2019) 15-21  
 Ding et al., [*HotQCD*], *PRL* 123 (2019) 6, 062002  
 Borsanyi et al., *PRL* 125(2020)5, 052001  
 Isserstedt et al. *PRD* 100 (2019) 074011  
 Gao, Pawłowski, *PLB* 820 (2021) 136584

Fu et al., *PRD* 101 (2020), 054032  
 Gunkel, Fischer, *PRD* 104 (2021) 5, 054022

# Road map

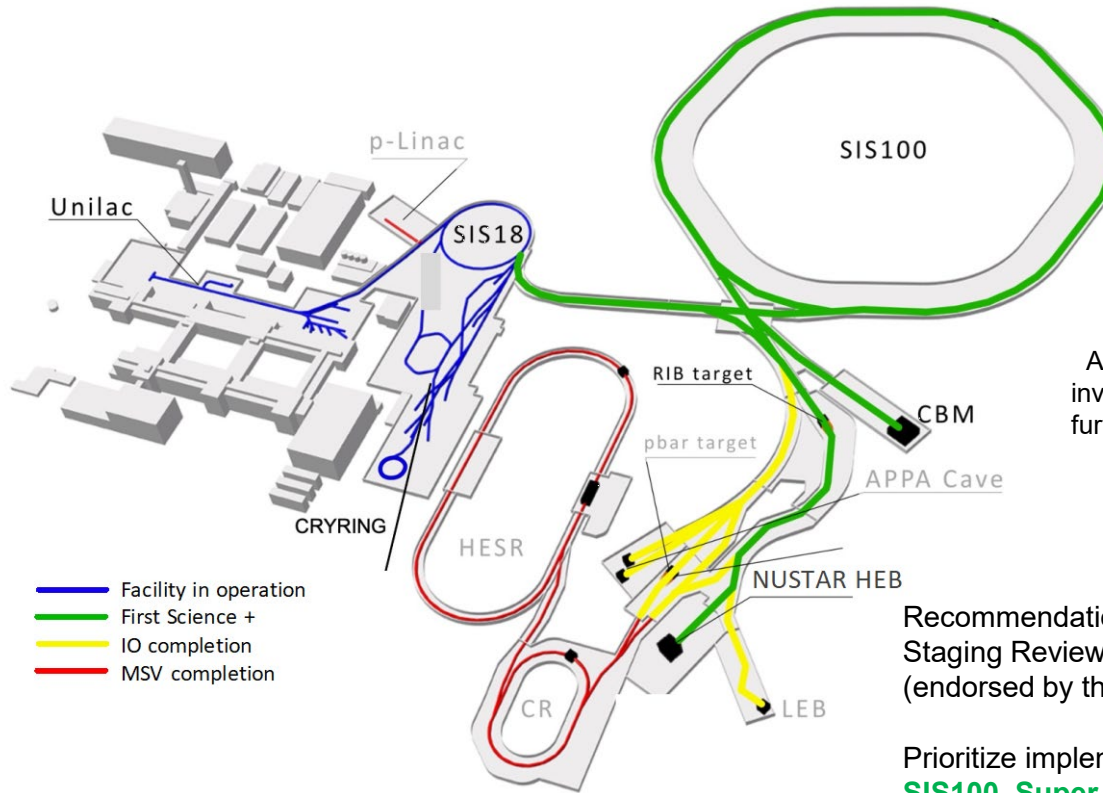


# Facility for Antiproton and Ion Research



2018 start of FAIR Phase-0  
~3 month beamtime/year

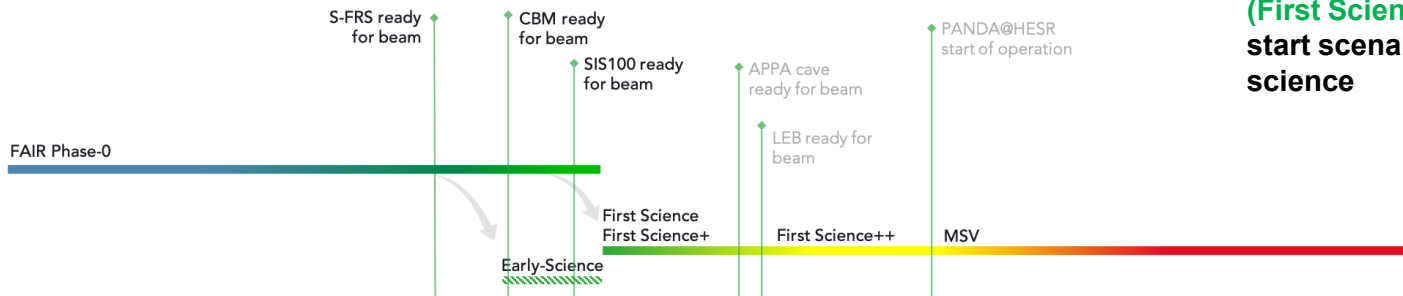
Intermediate forefront research program at GSI with improved beams and FAIR detectors



APPA, and PANDA investigate options for further experiments at green lines

Recommendation of the "First Science and Staging Review of the FAIR Project" (endorsed by the FAIR Council in Oct. 2022):

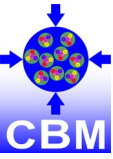
Prioritize implementation of Scenario #3: **SIS100, Super-FRS-HEB, CBM (First Science+)** – the most appropriate start scenario to achieve world-leading science



2023				2024				2025				2026				2027				2028				2029				2030				2031				2032				2033				2034				2035				2036							
01	02	03	04	01	02	03	04	01	02	03	04	01	02	03	04	01	02	03	04	01	02	03	04	01	02	03	04	01	02	03	04	01	02	03	04	01	02	03	04	01	02	03	04	01	02	03	04	01	02	03	04	01	02	03	04	01	02	03	04

JUL 2023

# Civil Work Completed



APPA  
cave

NUSTAR HEB

NUSTAR LEB

S-FRS

CRYO 2

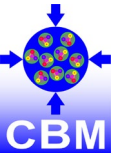
CBM

SIS100

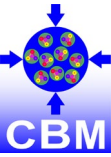
April 2024



# Civil Work Completed



# Compressed Baryonic Matter experiment



Fixed-target experiment → highest rates achievable

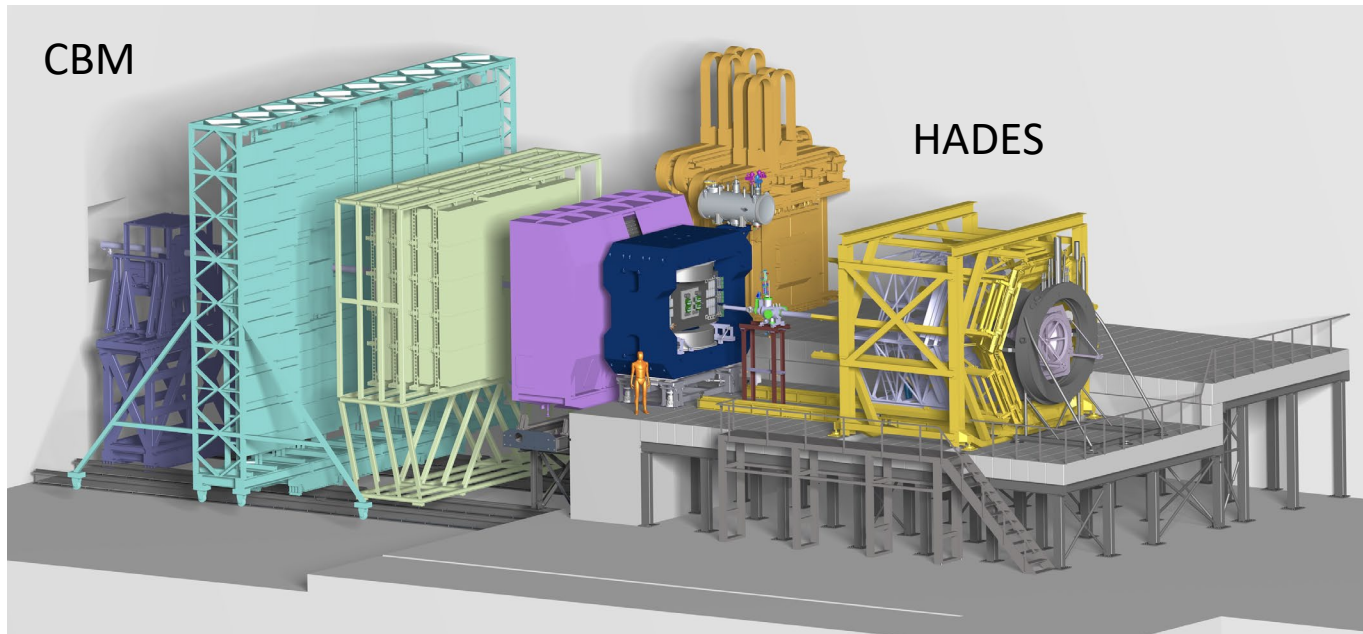
First beams in 2028/2029

Versatile subsystems → tailored for the physics program

Silicon-based tracking → fast and precise

Free-streaming front-end-electronics (FEE) → minimal dead-time while data acquisition

Online event selection → advanced data taking focused on customized needs



CBM

HADES

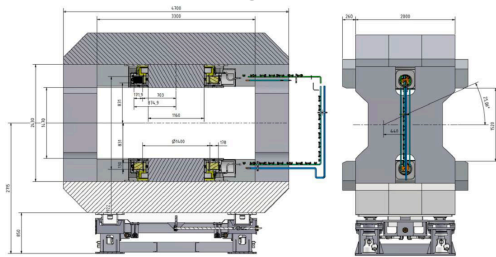
315 full members from 10 countries  
47 full member institutions  
10 associated member institutions

# CBM subsystems are on the verge of series production

→ pre-production is ongoing in all systems

## Superconducting dipole magnet

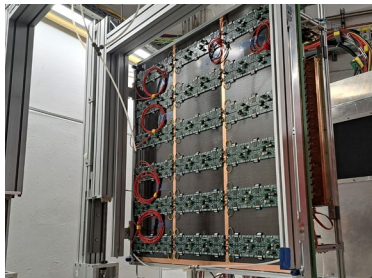
award of contract to Bilfinger Noell GmbH 20.12.2023



## Beam monitoring system



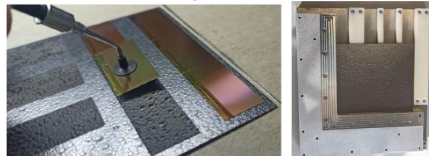
## Transition Radiation Detector



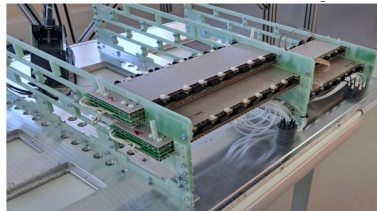
pre-production modules of 1D and 2D options ready

## Micro Vertex Detector

sensor/module integration



## Time of flight detector



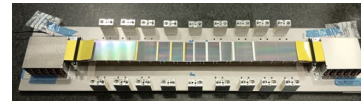
module pre-production concluded

## MUon Chamber system



test of full-size GEM and RPC prototypes

## Silicon Tracking System

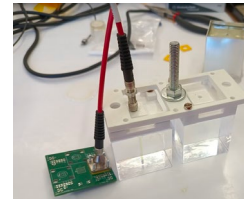


first STS series ladder



> 100 modules assembled

## Forward Spectator Detector

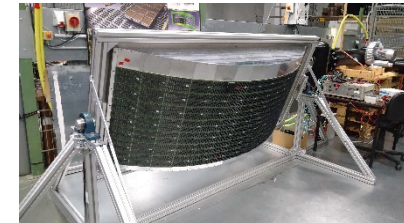


ZnS scintillators and LYSO crystals read-out via SiPM or/and PMT

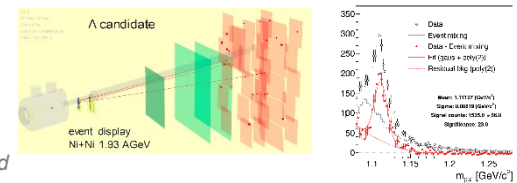
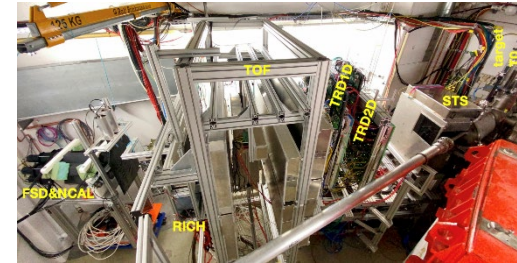
## Ring Imaging Cherenkov detector

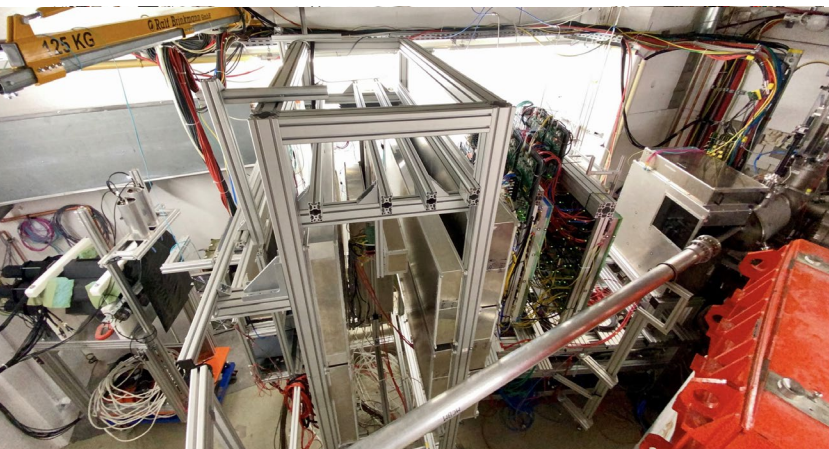
1 of 2 photo cameras ready

50% FEE produced



## Prototype of CBM online data processing tests with mCBM





FLES entry nodes  
CRI FPGA  
**μSlice building**  
(DAQ container)



FLES processing nodes  
time slice building  
event reconstruction & selection  
archiving

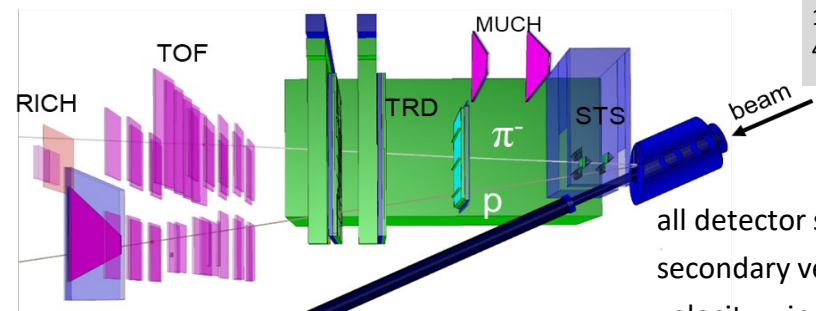


## Free-streaming CBM data transport

Pre-series productions of all CBM detector systems

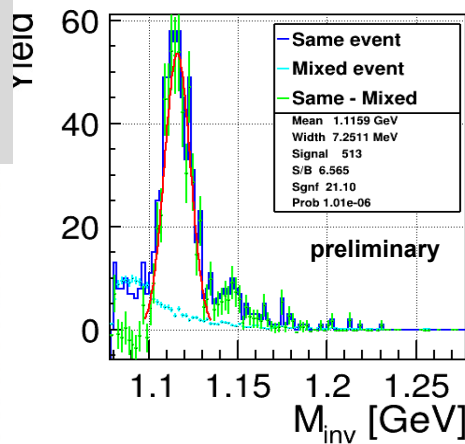
High-rate studies up to 10 MHz coll. rate in nucleus-nucleus collisions

## Rare signal reconstructed: $\Lambda \rightarrow p \pi^-$



**Ni+Ni 1.93 AGeV**  
run 2391 (May '22):  
 $10^9$  collisions, 1:57h  
400 kHz av. coll. rate

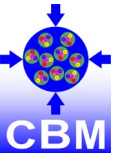
all detector systems involved  
secondary vertex  
velocity windows for p and  $\pi^-$   
candidate



## Campaign 2024:

high-rate studies  
online reconstruction and selection  
 $\Lambda$  baryons in Ni+Ni at 1.0 - 1.93 AGeV

# Compressed Baryonic Matter experiment



Fixed-target experiment → highest rates achievable

Versatile subsystems → tailored for the physics program

Silicon-based tracking → fast and precise

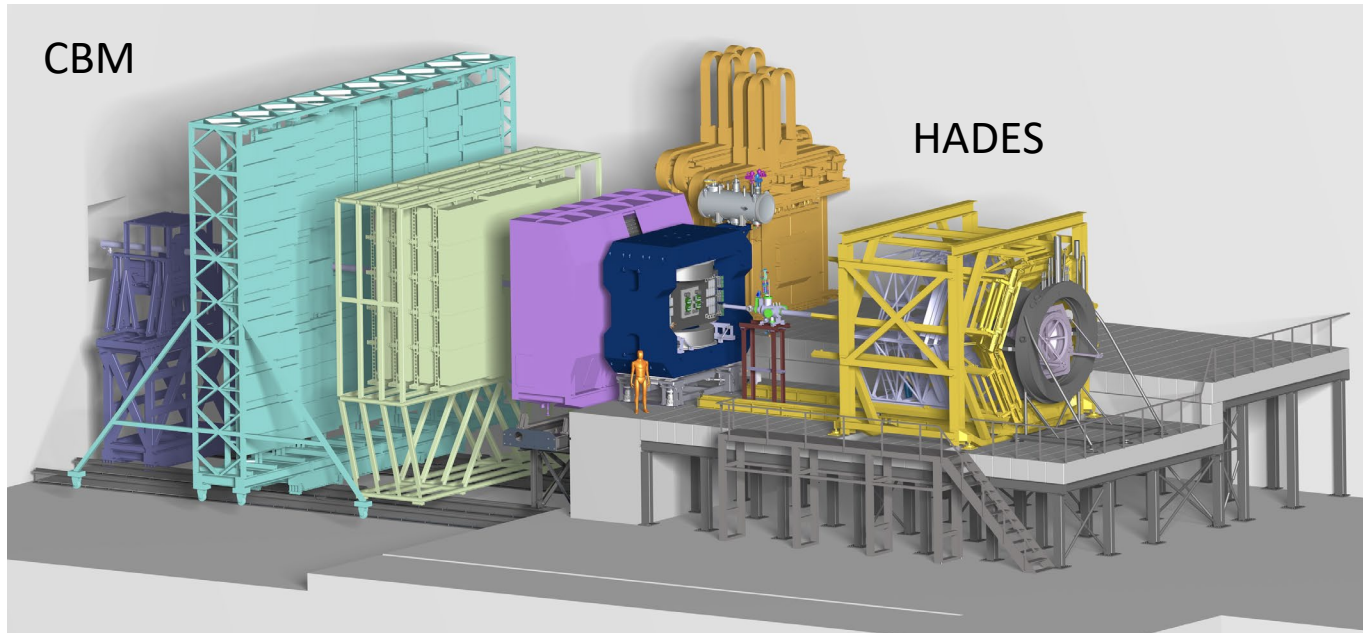
Free-streaming front-end-electronics (FEE) → minimal dead-time while data acquisition

Online event selection → advanced data taking focused on customized needs

First beams in 2028/2029

**Years 1-3:** first energy scan, improved statistical uncertainties of factor 10 with respect to STAR

**Years 4-8:** high-statistics measurements: di-lepton IMR, ultra-rare probes

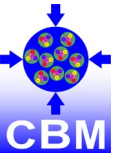


CBM

HADES

315 full members from 10 countries  
47 full member institutions  
10 associated member institutions

# Road map



What are we pursuing and why?

Physics motivation

Where are we now?

Current status

Who is involved?

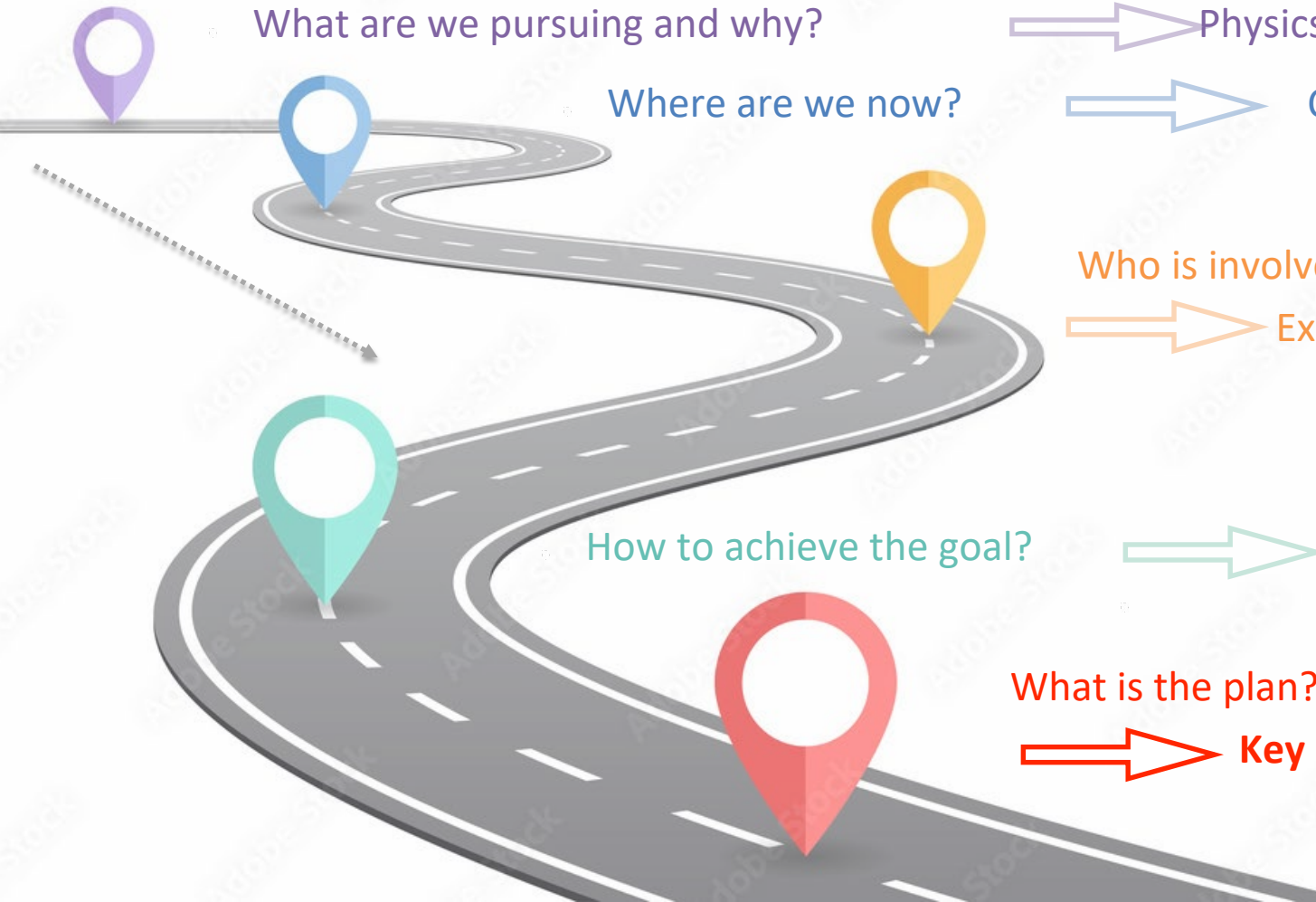
Experimental facilities

How to achieve the goal?

Compressed Baryonic Matter

What is the plan?

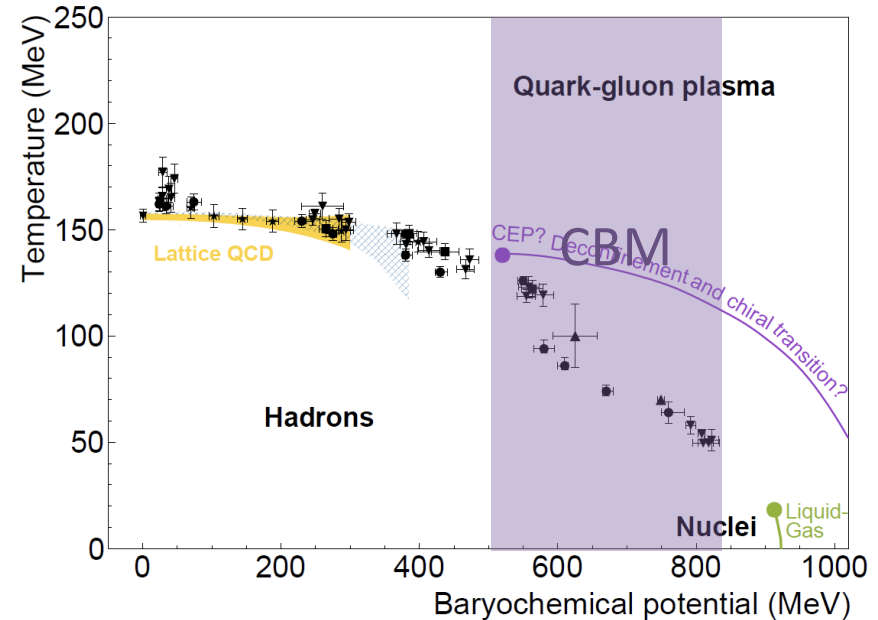
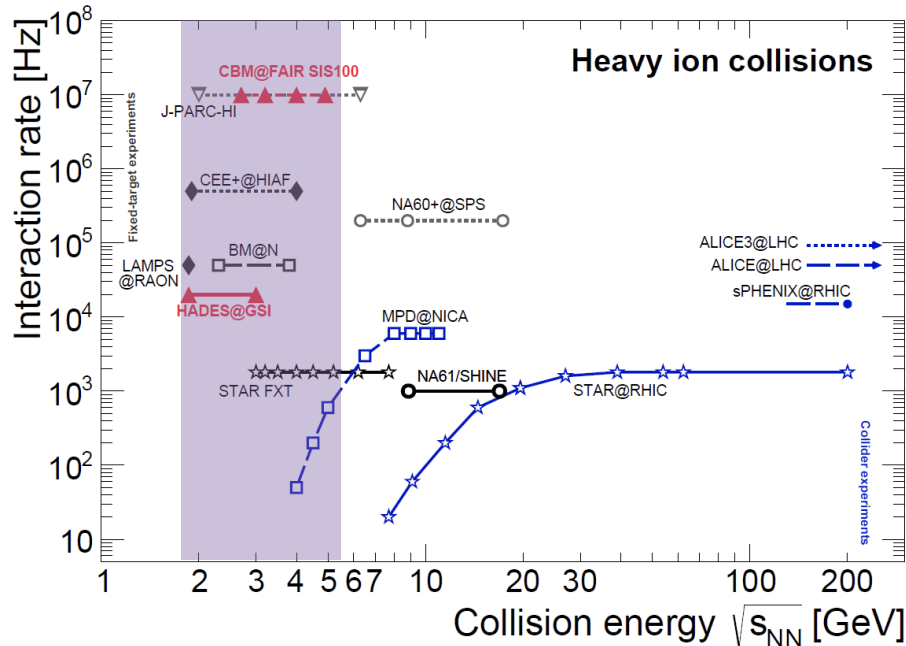
Key measurements



# Key observables

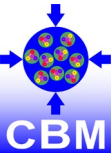
## Systematic measurements:

- **Fluctuations:** System alteration through first-order phase transition, critical point
- **Dileptons :** Emissivity: system's lifetime, temperature, density, in-medium characteristics
- **Hadrons (Strangeness, Charm, Hyper-nuclei, Bound states):** EOS: vorticity, collectivity, correlations: NN, YN, YY, multi-body interactions



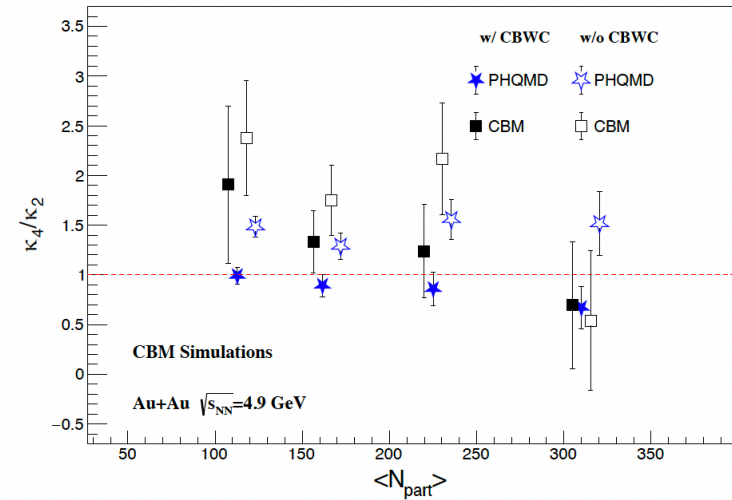
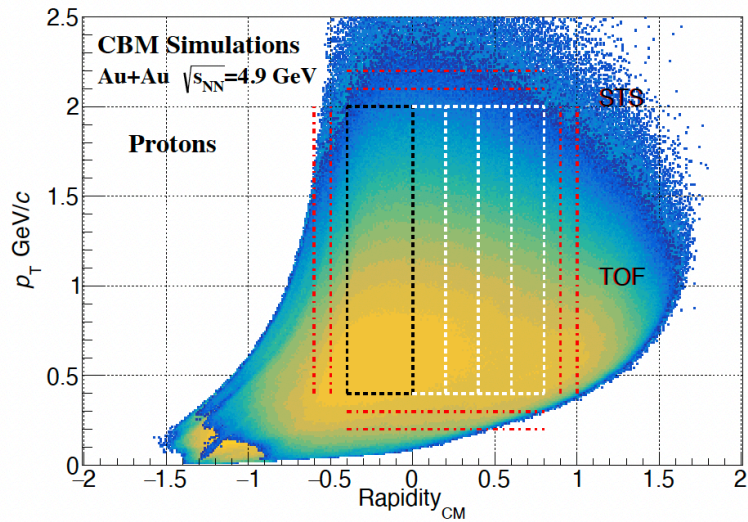
A high interaction rate is desired to reduce uncertainties and enable measurements that have so far been unattainable.

# Fluctuations



Corrections for volume fluctuations and conservation laws

- Event-by-event changes of efficiency
- Proper selection of  $y - p_T$  interval
- (Net-)baryons vs. protons, neutrons, nuclei

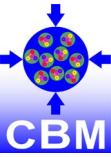


Expectations after ~3 years of running

- Full coverage of  $\kappa_4(E)$  for protons
- First results of  $\kappa_6$
- Possible addition of strangeness:  $\kappa_4(\Lambda)$



# Dileptons



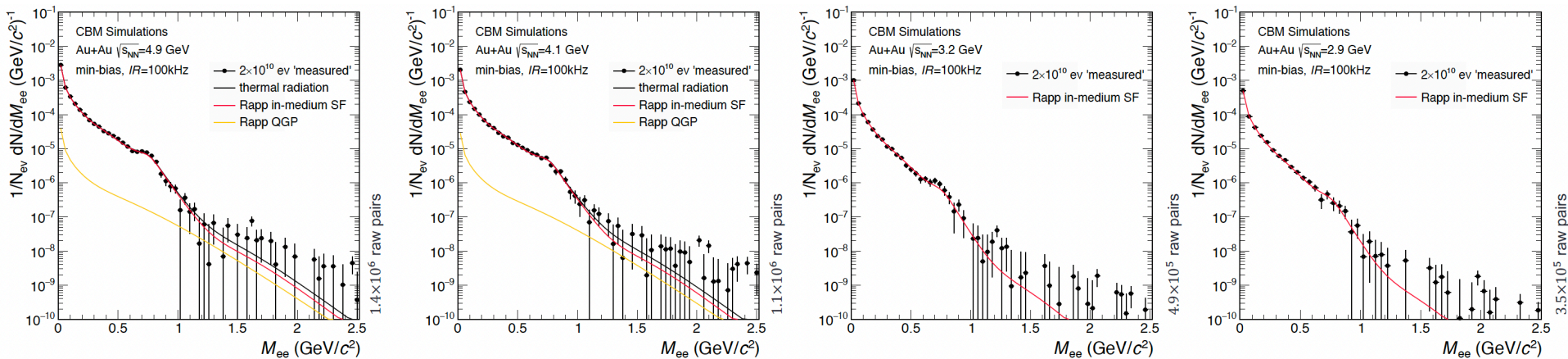
Electron thermal radiation, corrected for acceptance and efficiency,

Dominated by  $\rho$  contribution at LMR,

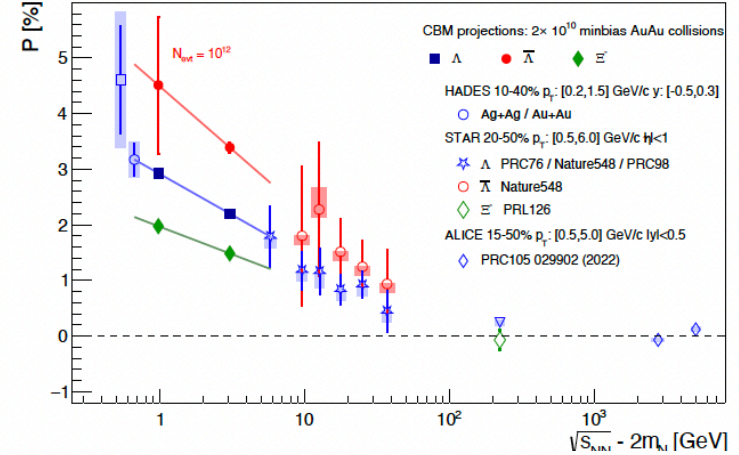
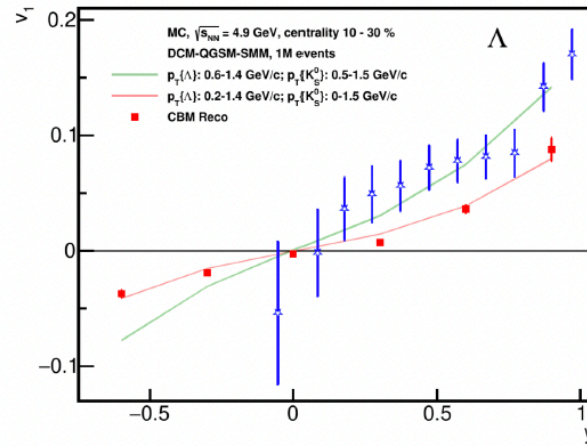
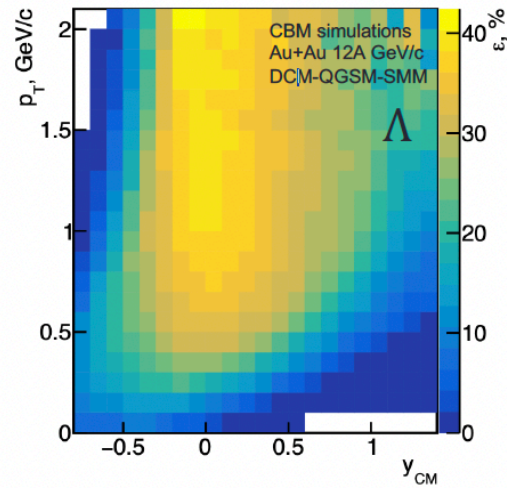
Can be reconstructed with 1.5-4.5% of precision,

Gives access to the fireball lifetime and electrical conductivity (transport properties)

$T$  vs. baryon density effects from partonic to hadronic fireballs



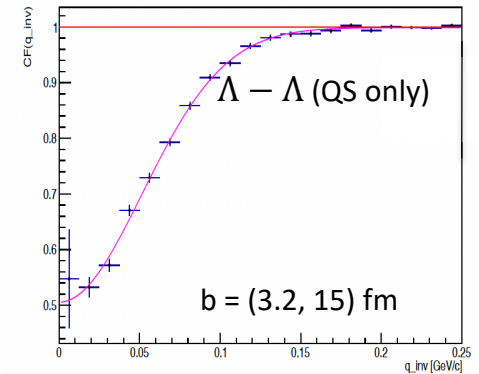
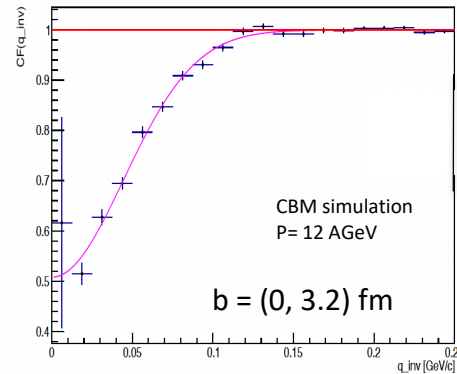
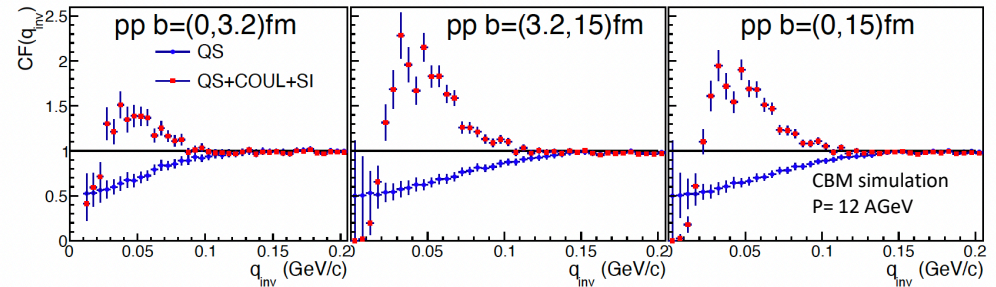
# Flow, polarization, correlations

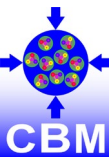


Excellent acceptance coverage

Reconstruction efficiency  $\sim 30\%$

- Precise measurements of flow for  $S=1, 2, 3$
- Polarization measurement with precision of  $\sim 5\%$
- Thorough multi body N and Y correlations of  $S=1, 2, 3$  achievable





CBM aims to answer fundamental questions about the structure of the QCD phase diagram at high  $\mu_B$

What are we pursuing and why?

Where are we now?

Already operating at high  $\mu_B$  experiments are complete and exploration of new physics needs higher interaction rates

Who is involved?

Many world-wide existing and planned facilities complement each other programs

How to achieve the goal?

Compressed Baryonic Matter experiment with high interaction rates will explore the region of the energies of the highest importance

What is the plan?

To start these exploration in 2028 and to answer fundamental questions in the first year of CBM running



CBM Collaboration Meeting,  
Prague, September 2024

The Future is  
**Bright**

Be the  
light

CBM is open for new  
participation



Thank You for your attention