

# Impact of first-principles calculations on the QCD phase diagram and equation of state

**Yuki Fujimoto**  
**(RIKEN & UC Berkeley)**



## References:

- [1] [Y. Fujimoto](#), K. Fukushima, L. McLerran, M. Praszalowicz, PRL129 (2022) [2207.06753]
- [2] [Y. Fujimoto](#), S. Reddy, PRD 109 (2024) (selected as Editors' Suggestion) [2310.09427]
- [3] [Y. Fujimoto](#), PRD109 (2024) [2312.11443]; [2408.12514]

# Neutron stars: why do we study now?

Now is the most exciting period because of...

- Recent advances in astrophysics
- Recent advances in QCD — the first principle

# Recent advances in QCD

- Higher-order computations of weak-coupling QCD EoS

Freedman,McLerran(1978); Baluni(1979); Kurkela,Romatschke,Vuorinen (2009);  
Gorda,Säppi,Paatelainen,Seppänen,Österman,Schicho,Navarrete (2018-)

- Nuclear EoS from chiral effective field theory ( $\chi$ EFT)

Tews,Krüger,Hebeler,Schwenk(2013);Drischler,Furnstahl,Melendez,Philips(2020);  
Keller,Hebeler,Schwenk(2022); ... many others

- Lattice simulations of QCD at finite isospin density

Kogut,Sinclair (2002); NPLQCD collaboration (2007-);  
Brandt,Chelnokov,Cuteri,Endrodi,... (2014-);

- Lattice simulations of two-color QCD at finite baryon density

e.g. Iida,Itou,Murakami,Suenaga (2024)

- Hadron-hadron interaction from the lattice QCD

HAL QCD collaboration (2006-)

- Hamiltonian lattice simulations of QCD in (1+1)-dimensions

Hayata,Hidaka,Nishimura (2023)

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Freedman,McLerran(1978); Baluni(1979); Kurkela,Romatschke,Vuorinen (2009);  
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# QCD at finite isospin density

Alford,Kapustin,Wilczek (1999); Kogut,Sinclair (2002-);  
Beane,Detmold,Savage et al. (2007-);  
Endrodi et al. (2014-)...

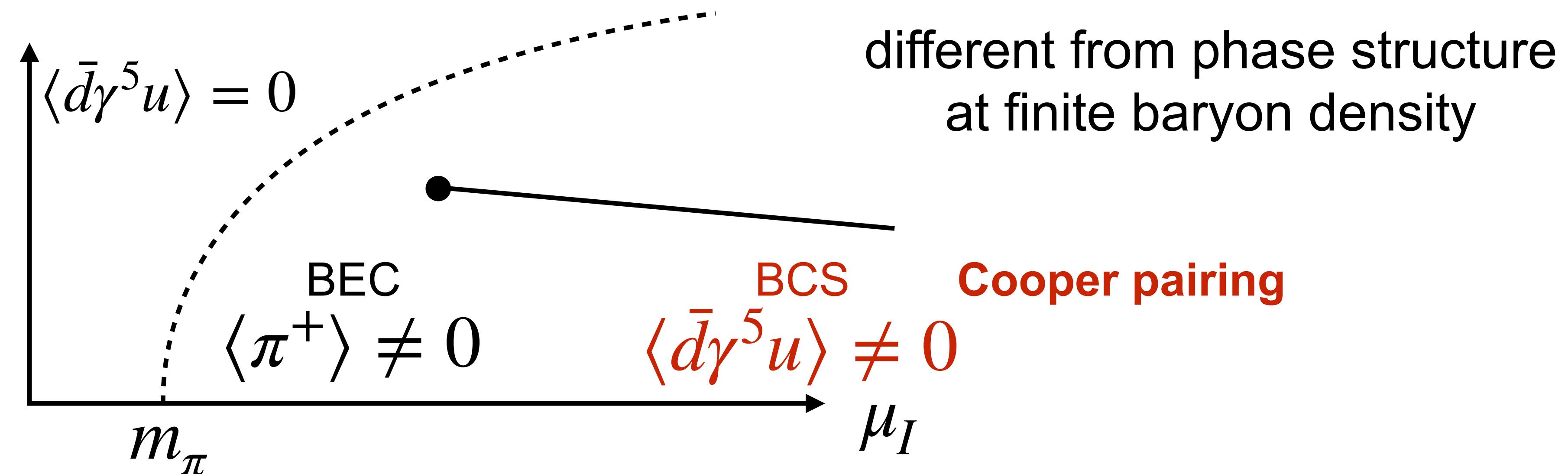
- **No sign problem** → can be simulated on the lattice!

- Isospin chemical potential (conjugate to isospin density  $I_3$ ):

$$\mu_u = \frac{\mu_I}{2}, \quad \mu_d = -\frac{\mu_I}{2} \dots \text{Fermi surface of } u \text{ & } \bar{d}$$

- Phase structure:

Son,Stephanov (2000)

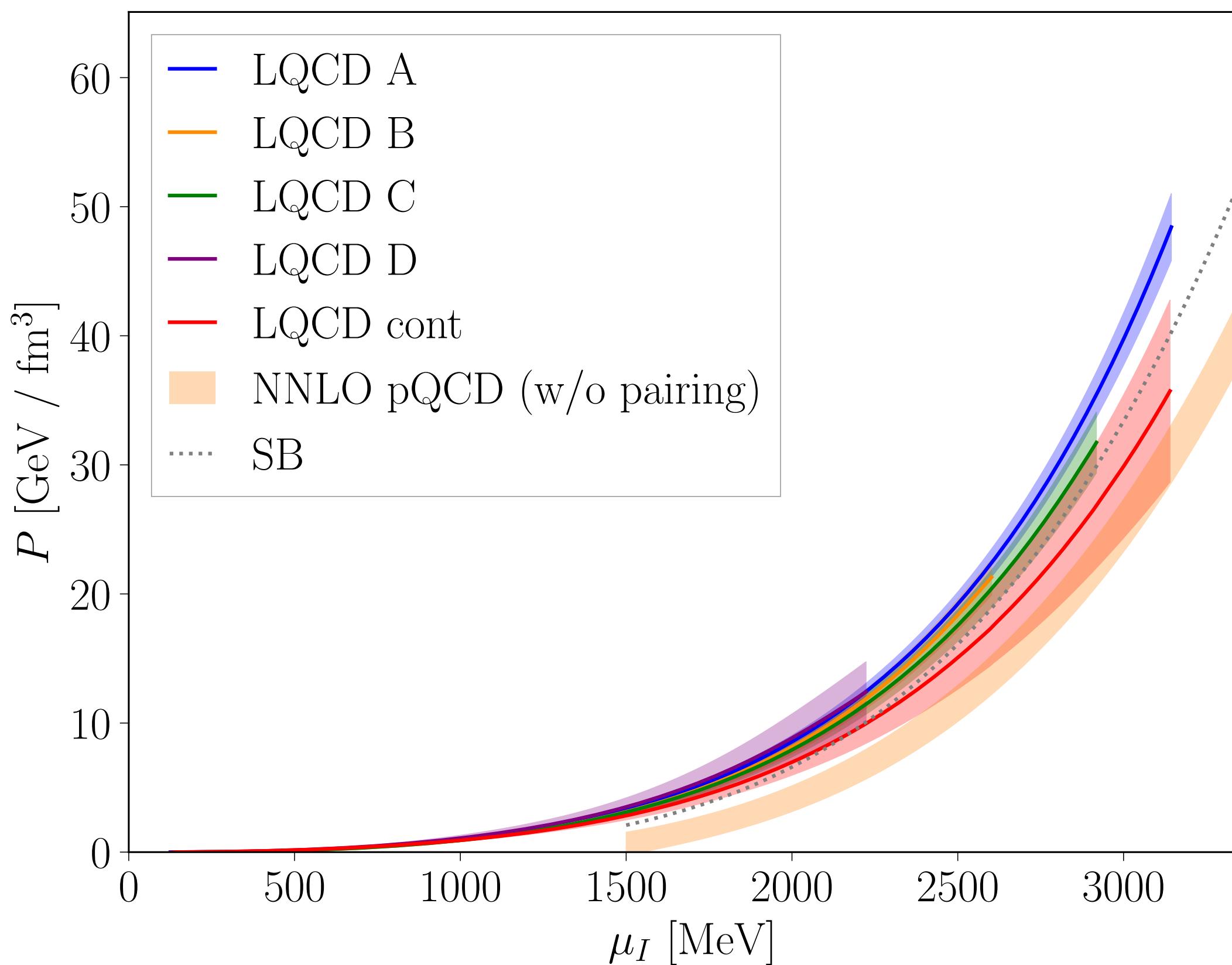


# QCD at finite isospin density

**Recent impact:**

Abbott et al. (NPLQCD) (2023, 24)

EoS is calculated up to  $\mu_I \sim 3$  GeV by lattice QCD in the continuum limit



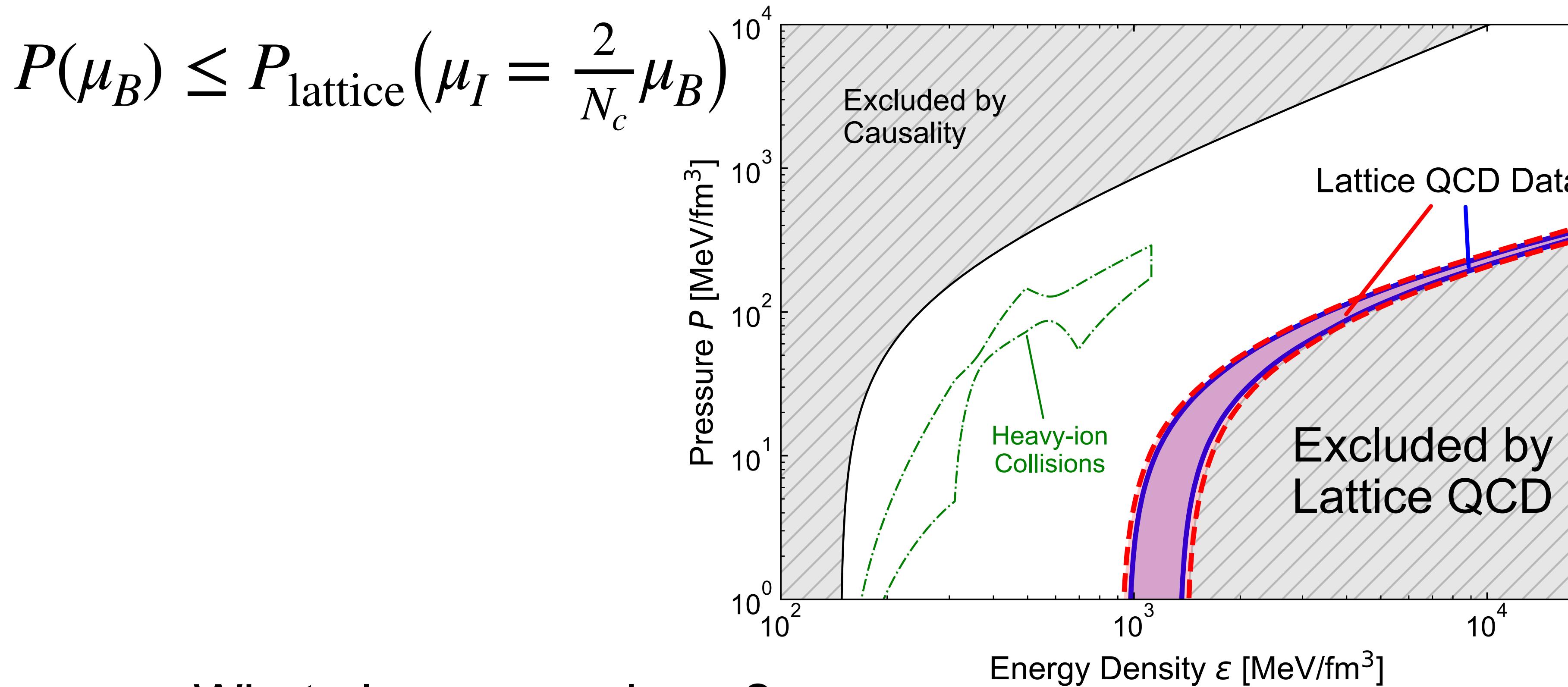
→ comparison with the weak-coupling results feasible

# QCD at finite isospin density

Abbott et al. (NPLQCD) (2023, 24)

What can we learn from this?

→ Bounds on the symmetric nuclear matter EoS from QCD inequality



Cohen (2003);  
Fujimoto, Reddy, PRD 109 (2023)

cf. Moore, Gorda (2023)

What else can we learn?

→ Weak-coupling results can be constrained by comparison w/ lattice QCD results

# Outline

## 1. Conformal EoS in neutron star cores

- useful observables: trace anomaly
- Bayesian inference from astrophysical data
- role of the pQCD constraint

## 2. Interplay between weak-coupling and lattice QCD

Following issues about the weak-coupling results:

- applicability
- fixing the undetermined parameter in pQCD
- color-superconducting gap

# **Part 1: Conformal EoS in NS cores**

# Trace anomaly equation

Related to scale/conformal nature of matter:

$$j_D^\nu = x_\mu T^{\mu\nu} \rightarrow \partial_\nu j_D^\nu = T_\mu^\mu \left\{ \begin{array}{l} = 0 \\ \neq 0 \end{array} \right. \quad \begin{array}{l} \text{Classical YM} \\ \text{in QFT (RG effect)} \end{array}$$

# Finite- $\mu_B$ part of the trace anomaly (interaction measure):

$\langle T_\mu^\mu \rangle_{\mu_B} = \varepsilon - 3P$  ... can be read out from NS EoS

**Normalized trace anomaly:**  $\Delta_{\text{tr}} \equiv \frac{\langle T^\mu_\mu \rangle_{\mu_B}}{3\varepsilon} = \frac{\varepsilon - 3P}{3\varepsilon}$

# Conformal limit

Limit  $\alpha_s \rightarrow 0$  (conformal theory) is achieved when  $\varepsilon \rightarrow \infty$ .

The weak-coupling EoS in this **conformal limit** is:

**Trace anomaly:**  $\varepsilon - 3P \sim -\beta_0 \mu^4 \left( \frac{\alpha_s}{\pi} \right)^2 \rightarrow 0^+$

**Sound speed:**  $v_s^2 = \frac{dP}{d\varepsilon} \sim \frac{1}{3} \frac{1}{1 - \beta_0 \left( \frac{\alpha_s}{\pi} \right)^2} \rightarrow \frac{1}{3} - 0^+ \quad (\beta_0 < 0)$

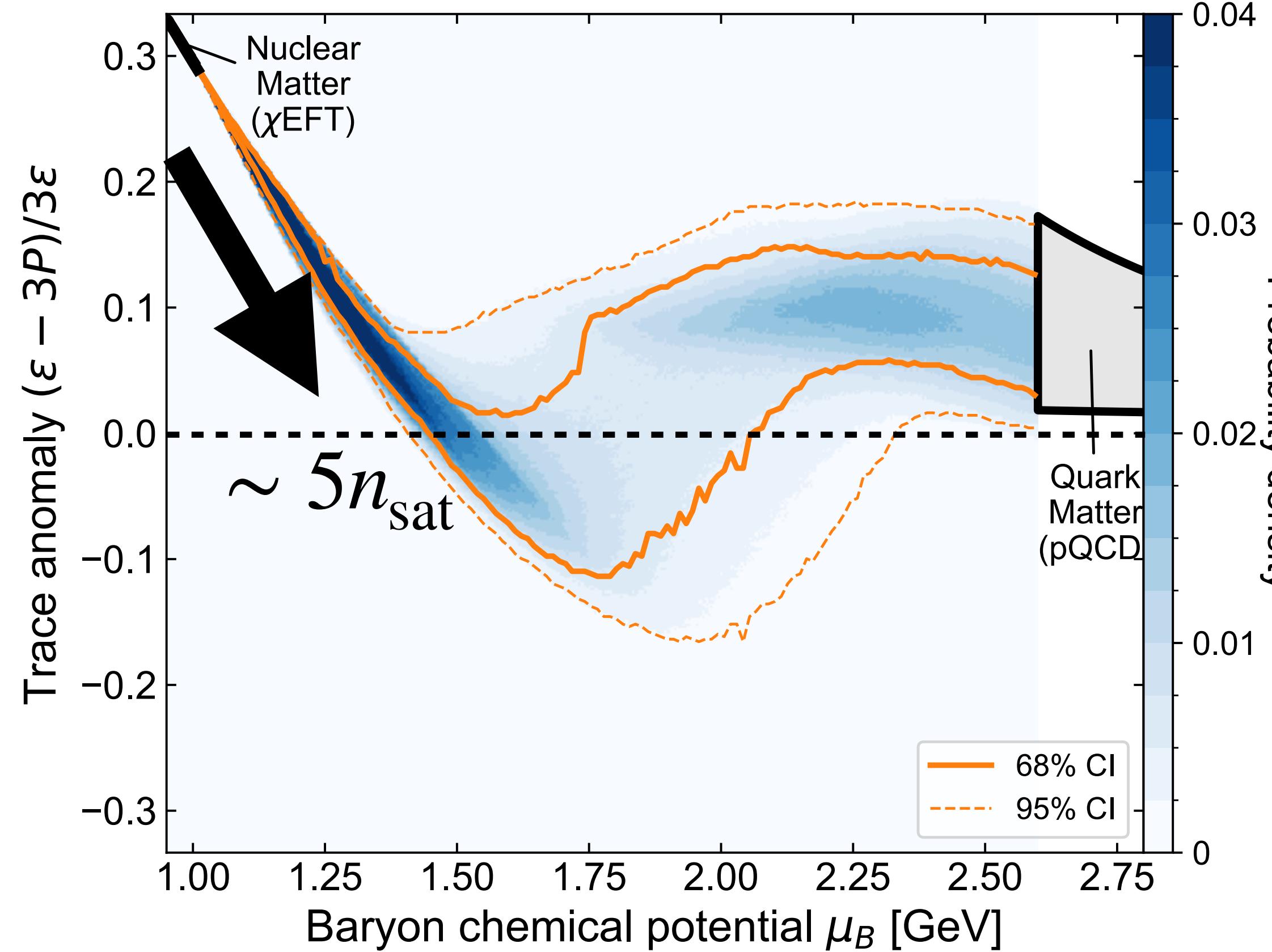
At the intermediate density,  $\varepsilon - 3P = 0$  and  $v_s^2 = 1/3$  are **different conditions**

# Trace anomaly from neutron star data

Fujimoto, Fukushima, McLerran, Praszalowicz, PRL129 (2022)

Normalized trace anomaly:  $\Delta_{\text{tr}} \equiv \frac{\langle T_{\mu}^{\mu} \rangle_{\mu_B}}{3\varepsilon} = \frac{\varepsilon - 3P}{3\varepsilon}$

Results of Bayesian inference: Method: Annala et al. (2019); Altiparmak, Ecker, Rezzolla (2022)



$\Delta_{\text{tr}} \sim 0$  already at  $\sim 5n_{\text{sat}}$   
→ **rapid approach to  
conformal EoS  $P \approx \varepsilon/3$**

Suggests the existence of  
strongly-coupled  
conformal matter?

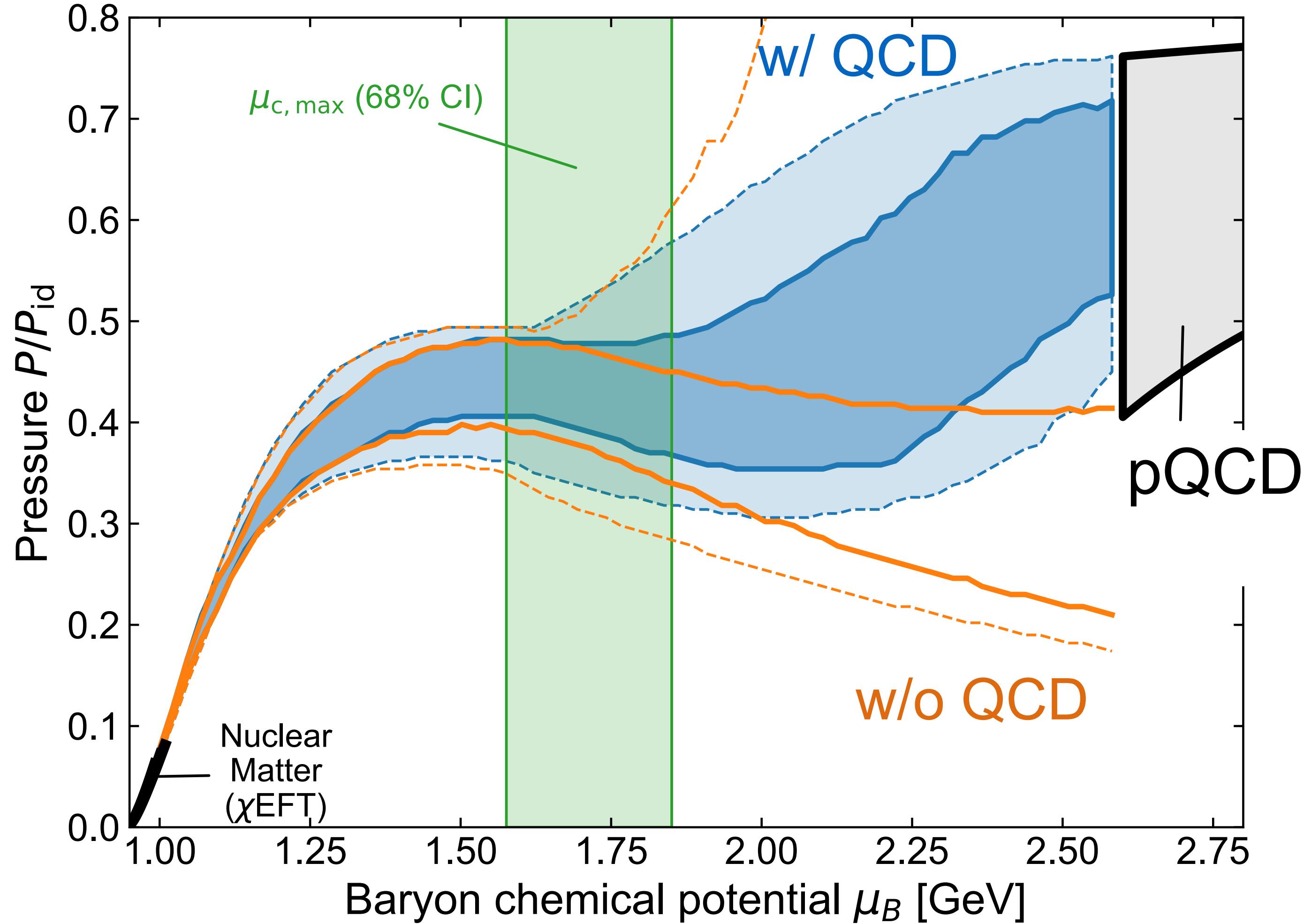
# Trace anomaly, effective d.o.f. & role of pQCD

Fujimoto, Fukushima, McLerran, Praszalowicz, PRL129 (2022)

- Trace anomaly:  
related to the changes in the effective  
degrees of freedom  $\nu$

$$\varepsilon - 3P \propto \frac{d\nu}{d \ln \mu}$$
$$(\nu = P/P_{\text{ideal}})$$

- It is natural to assume  $\nu$  is increasing  
 $\rightarrow$  positive  $\varepsilon - 3P$
- **Indeed, positive  $\varepsilon - 3P$  favored  
by QCD effect**



See also: Annala, Gorda, Hirvonen, Komoltsev, Kurkela, Nättilä, Vuorinen (2023)

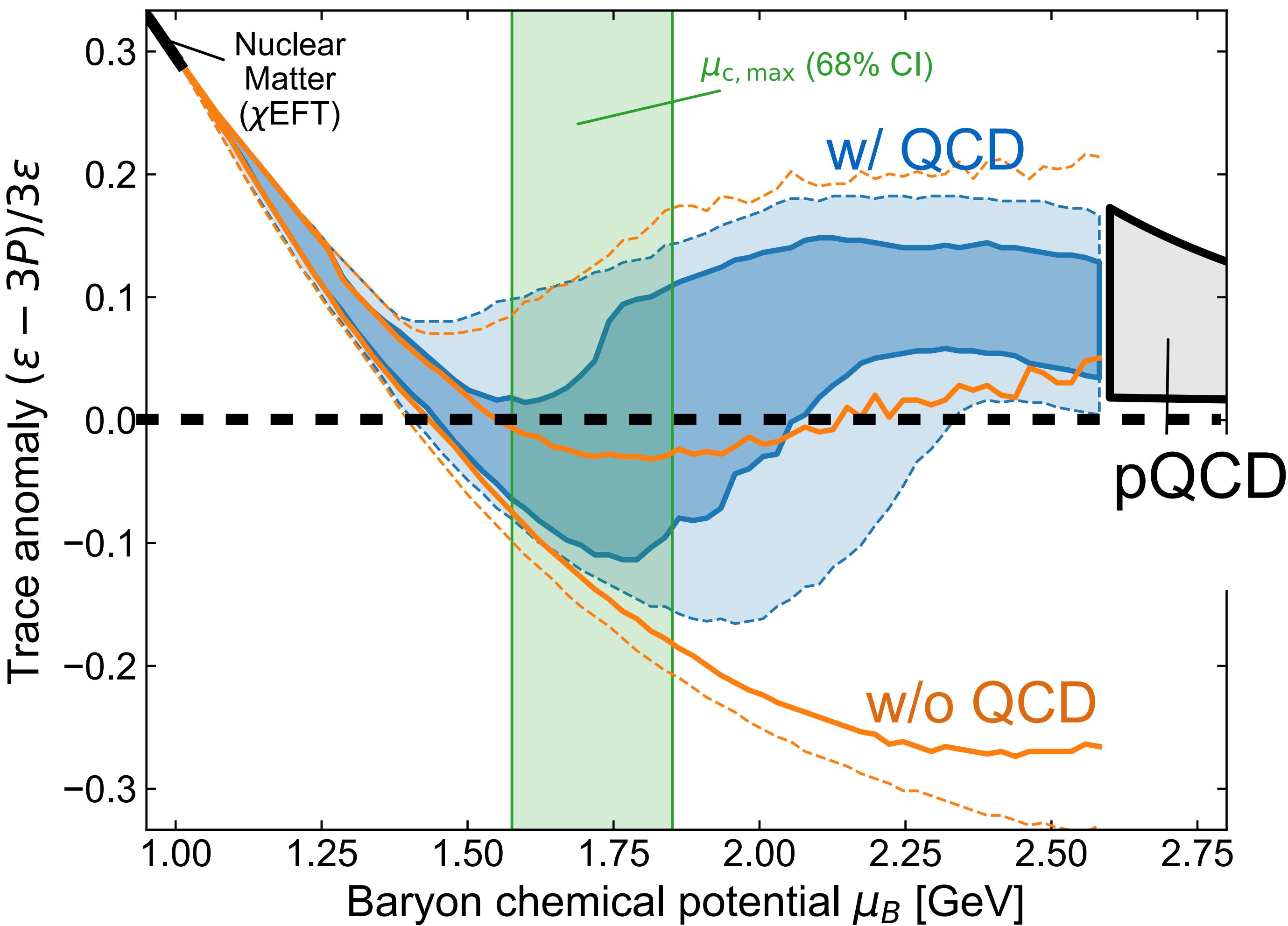
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# Trace anomaly and speed of sound

Fujimoto, Fukushima, McLellan, Praszalowicz, PRL129 (2022)

Speed of sound can be decomposed into  $\Delta_{\text{tr}}$  and its derivative:

$$\begin{aligned} v_s^2 &= \varepsilon \frac{d}{d\varepsilon} \left( \frac{P}{\varepsilon} \right) + \frac{P}{\varepsilon} \\ &= \underbrace{\varepsilon \frac{d\Delta_{\text{tr}}}{d\varepsilon}}_{\text{Derivative component}} + \underbrace{\left( \frac{1}{3} - \Delta_{\text{tr}} \right)}_{\text{Non-derivative component}} \end{aligned}$$

This explains why  $\varepsilon - 3P = 0$  and  $v_s^2 = \frac{1}{3}$  are the different conditions

# Trace anomaly and speed of sound

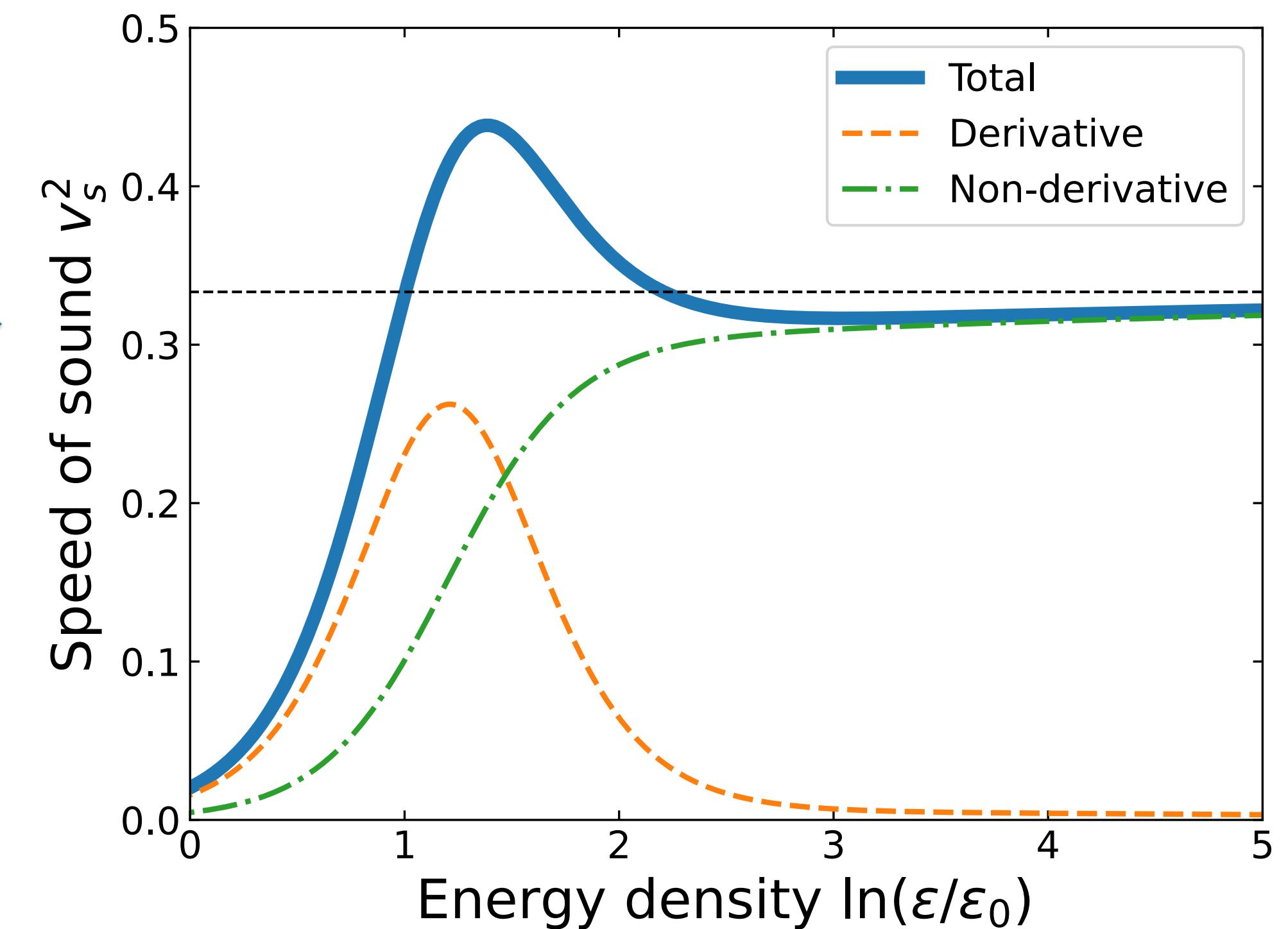
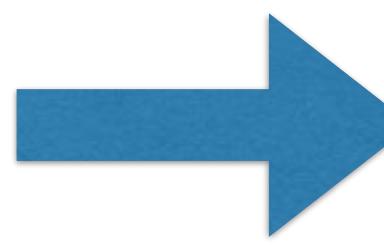
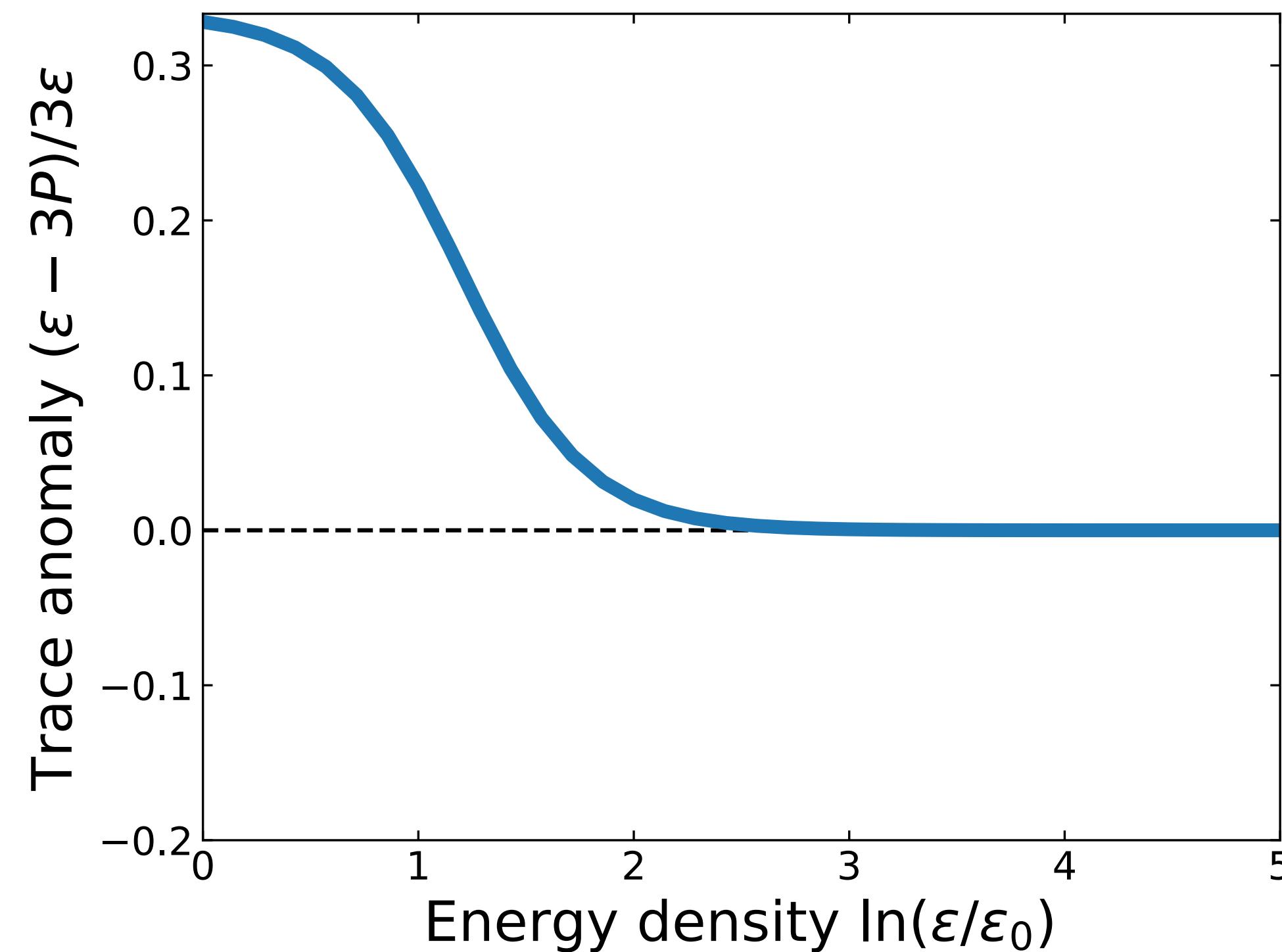
Fujimoto, Fukushima, McLellan, Praszalowicz, PRL129 (2022)

Rapid approach to  $\Delta_{\text{tr}} \rightarrow 0$  naturally spikes  $v_s^2$

$$\text{Trace anomaly } \Delta_{\text{tr}} = \frac{\varepsilon - 3P}{3\varepsilon}$$

$$\text{Sound velocity } v_s^2 = \varepsilon \frac{d\Delta_{\text{tr}}}{d\varepsilon} + \left( \frac{1}{3} - \Delta_{\text{tr}} \right)$$

Derivative      Non-derivative

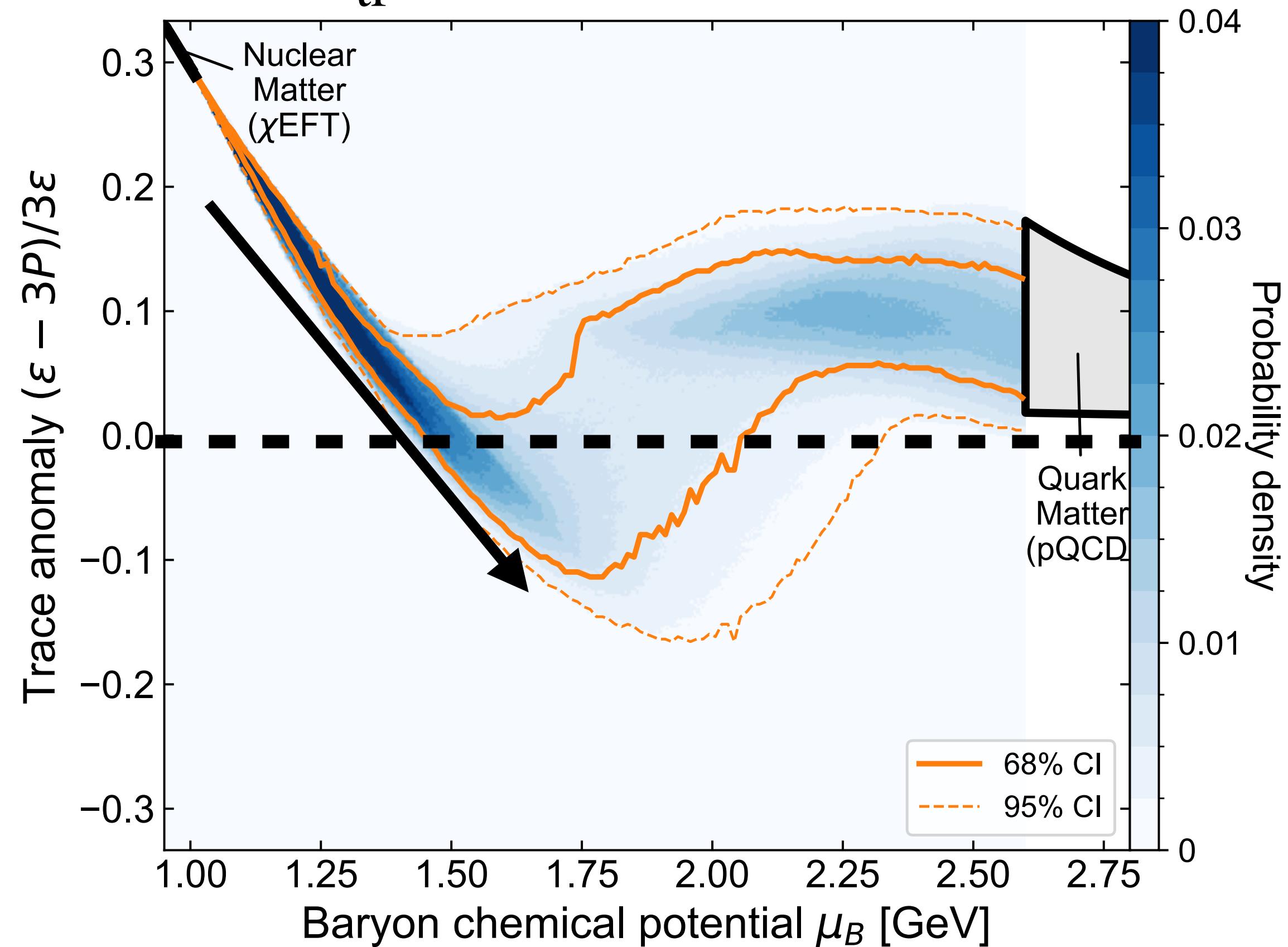


# Trace anomaly and speed of sound

Fujimoto, Fukushima, McLerran, Praszalowicz, PRL129 (2022)

Normalized trace anomaly:

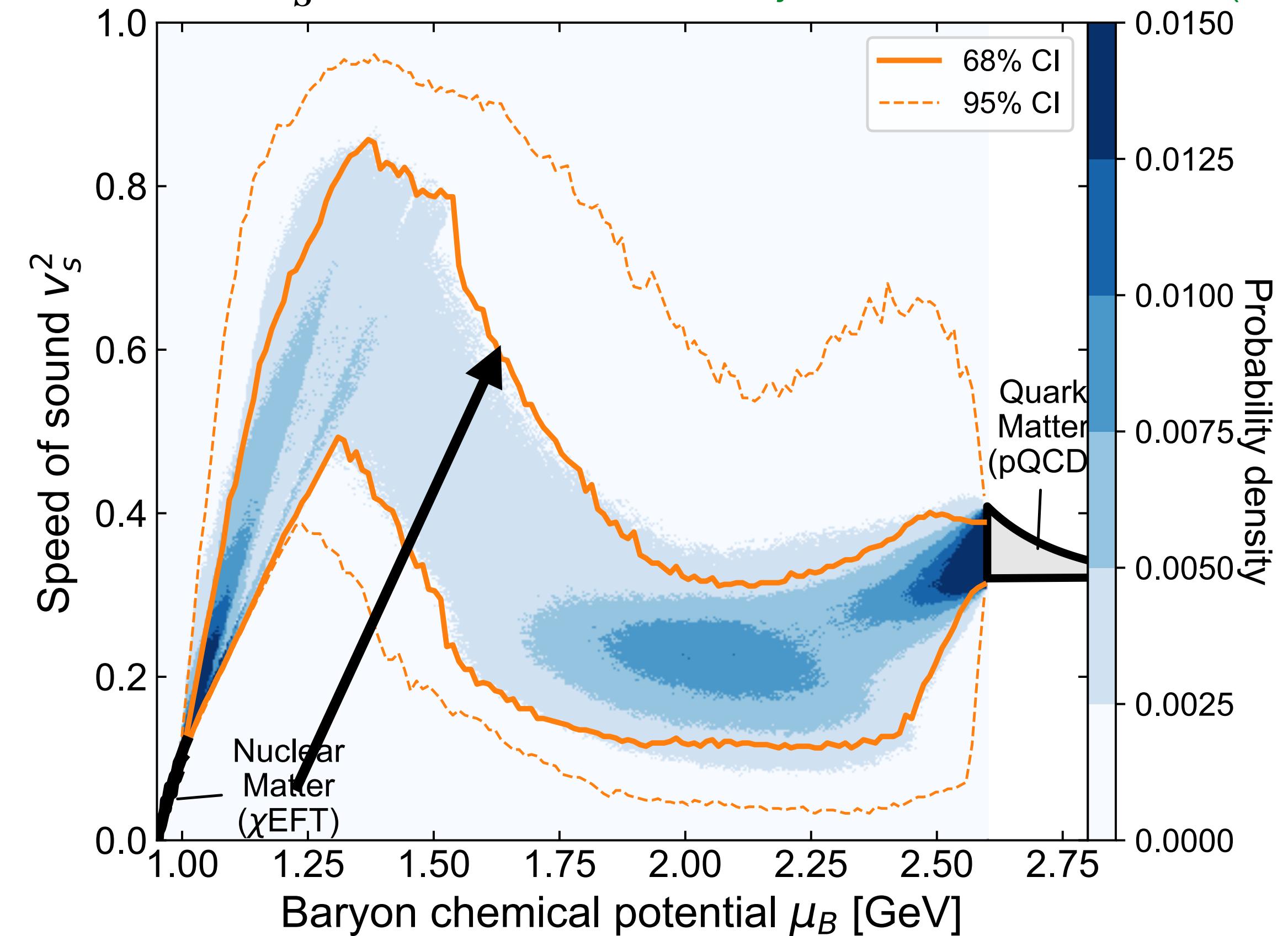
$$\Delta_{\text{tr}} = (\varepsilon - 3P)/3\varepsilon$$



Sound speed:

$$v_s^2 = dP/d\varepsilon$$

see also: Bedaque,Steiner (2014);  
Tews,Carlson,Gandolfi,Reddy (2018);  
Fujimoto,Fukushima,Murase (2019)



Rapid approach to  $\varepsilon - 3P \rightarrow 0$  drives the peak in  $v_s^2$  in actual data

# **Part 2: Interplay between weak-coupling and lattice QCD at high density**

# Notation

- $\text{QCD}_{\textcolor{red}{I}}$ : QCD at finite  $\mu_I$  and zero  $\mu_B$
- $\text{QCD}_{\textcolor{red}{B}}$ : QCD at finite  $\mu_B$  and zero  $\mu_I$

# Weak-coupling results in high-density QCD

Freedman, McLerran (1978); Kurkela et al. (2009-)

**QCD EoS in weak-coupling  $\alpha_s$  expansion:**

$$P_{\text{QCD}}(\mu) = \frac{3\mu^4}{4\pi^2} [1 - \mathcal{O}(\alpha_s)] + \frac{3\mu^2 \Delta^2}{2\pi^2} [1 + \mathcal{O}(\alpha_s^{1/2})]$$

where

$$\ln\left(\frac{\Delta_{\text{gap}}}{\mu}\right) = -b_{-1} \left(\frac{\alpha_s}{\pi}\right)^{-1/2} - b_0$$

Son (1999), Brown,Liu,Ren (1999); Wang,Rischke (2001)  
Review: Alford,Rajagopal,Schafer,Schmitt (2008);

**Applicability at low  $\mu$ ?**

- Usually, it is used down to  $\mu \sim 0.9$  GeV for the input of neutron stars

Kurkela,Fraga,Vuorinen (2014)

Weak-coupling formula is universal for  $\text{QCD}_B$  and  $\text{QCD}_I$  up to  $\mathcal{O}(\alpha_s^2)$

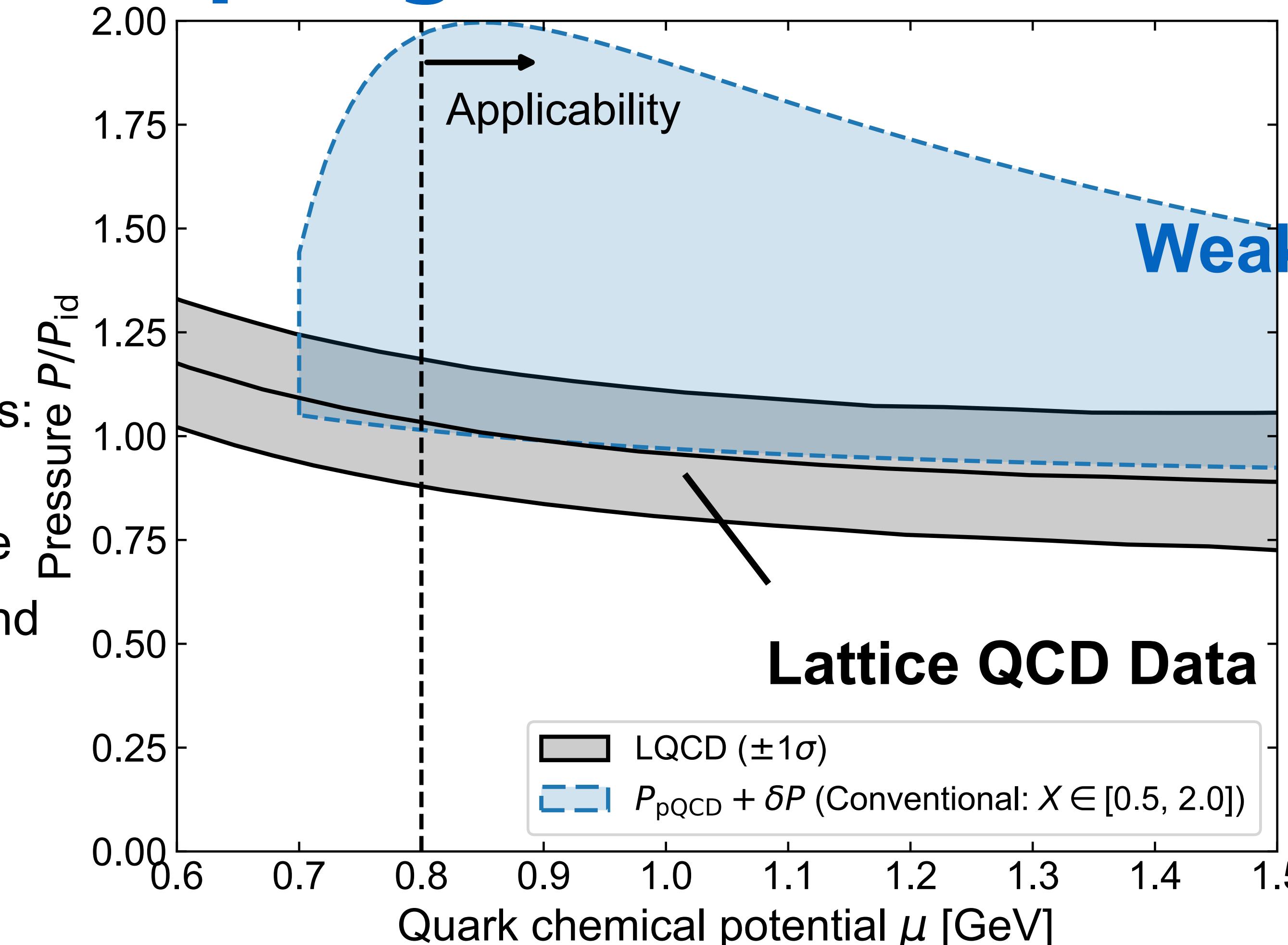
→ Lattice  $\text{QCD}_I$  can be used as a benchmark

# Weak-coupling results vs lattice QCD<sub>I</sub> data

Lattice data: Abbott et al. (2023, 24);  
Fujimoto (2023, 24)

Uncertainty in  
weak-coupling results:  
varying the  
renormalization scale  
 $\bar{\Lambda}$  by a factor 2 around  
its typical scale

$$\bar{\Lambda} = 2\mu$$



Empirical evidence for the dense-QCD weak-coupling results  
to be applicable down to  $\mu \sim 0.8$  GeV

At least the magnitude is correct

# “Uncertainty” in pQCD

Fraga,Pisarski,Schaffner-Bielich(2001);  
Kurkela,Romatschke,Vuorinen (2009)

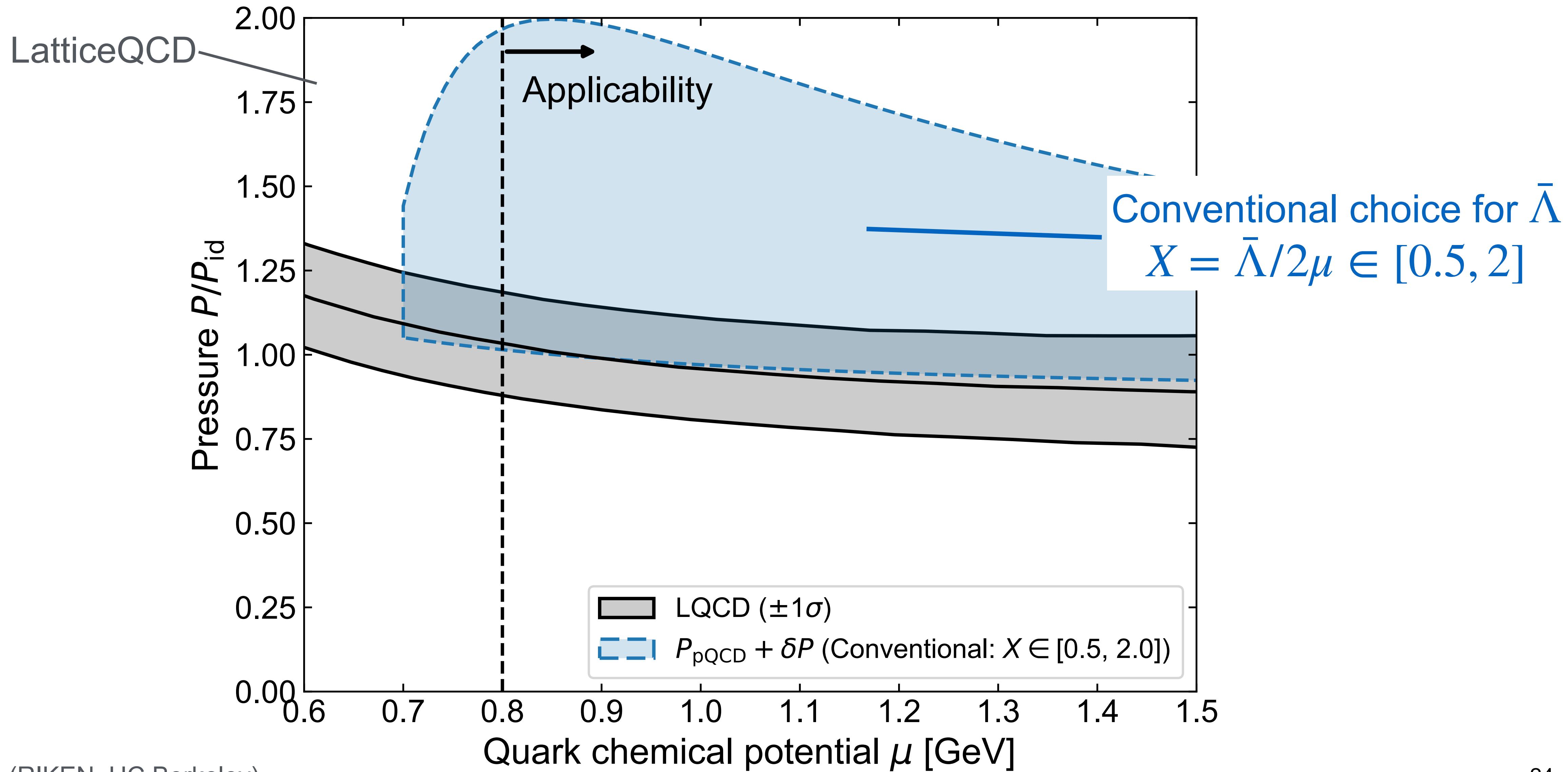
$$\alpha_s \simeq \frac{1}{\beta_0 \ln \left( \frac{\bar{\Lambda}}{\Lambda_{\overline{\text{MS}}}} \right)}$$

- $\bar{\Lambda}$ : **renormalization scale**
  - ... only ambiguity in pQCD from perturbative series truncation
- Canonical choice:  $\bar{\Lambda} = 2\mu$  (typical hard interaction scale)
- “Uncertainty” quantified by varying by factor 2
  - i.e.  $X \in [0.5, 2]$  with  $X \equiv \bar{\Lambda}/(2\mu)$
  - ... ad hoc procedure, purely based on historical practice

cf. Gorda,Komoltsev,Kurkela,Mazeliauskas (2022)

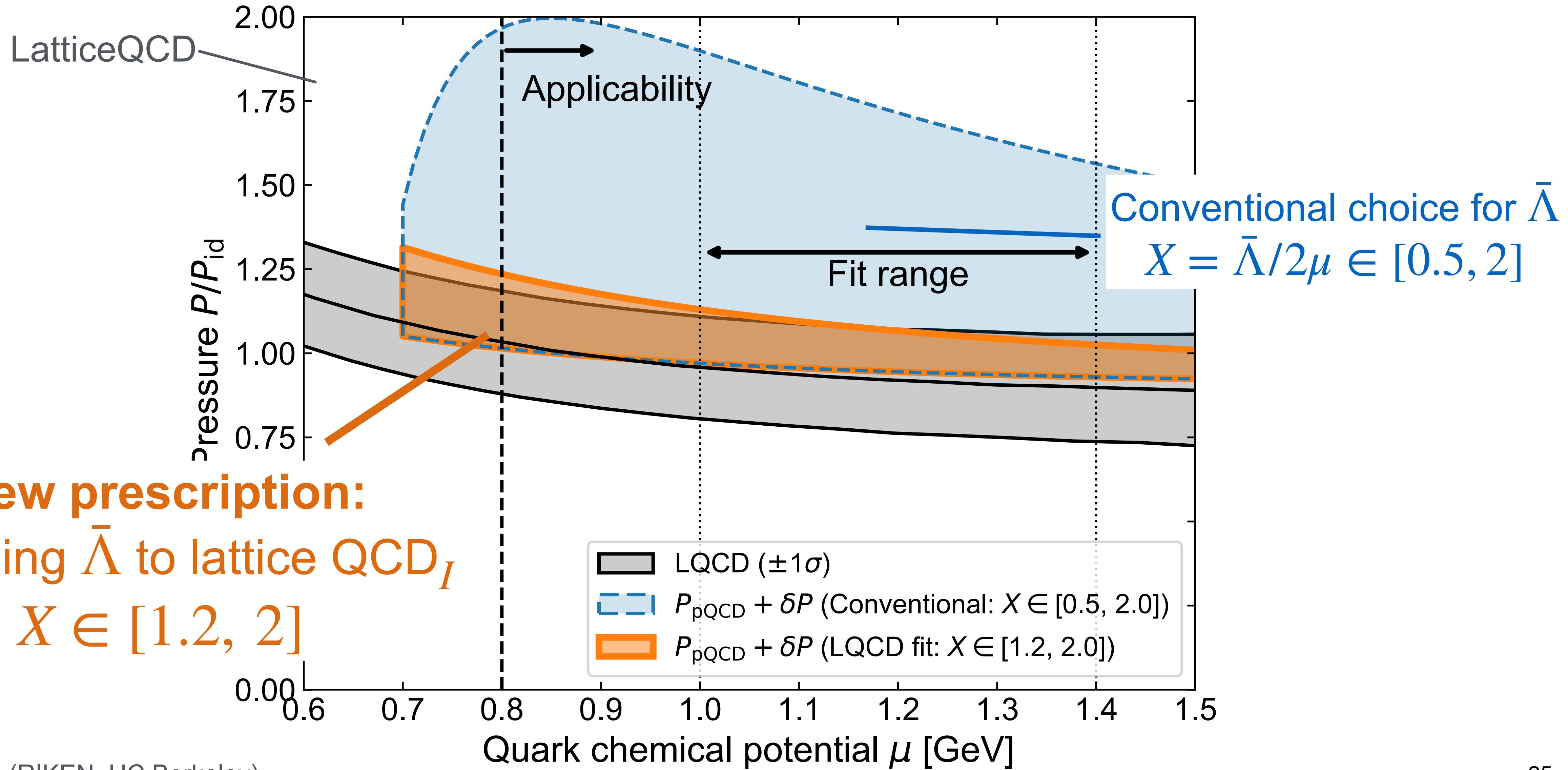
# Prescription for $\bar{\Lambda}$ determination

Fujimoto, 2408.12514



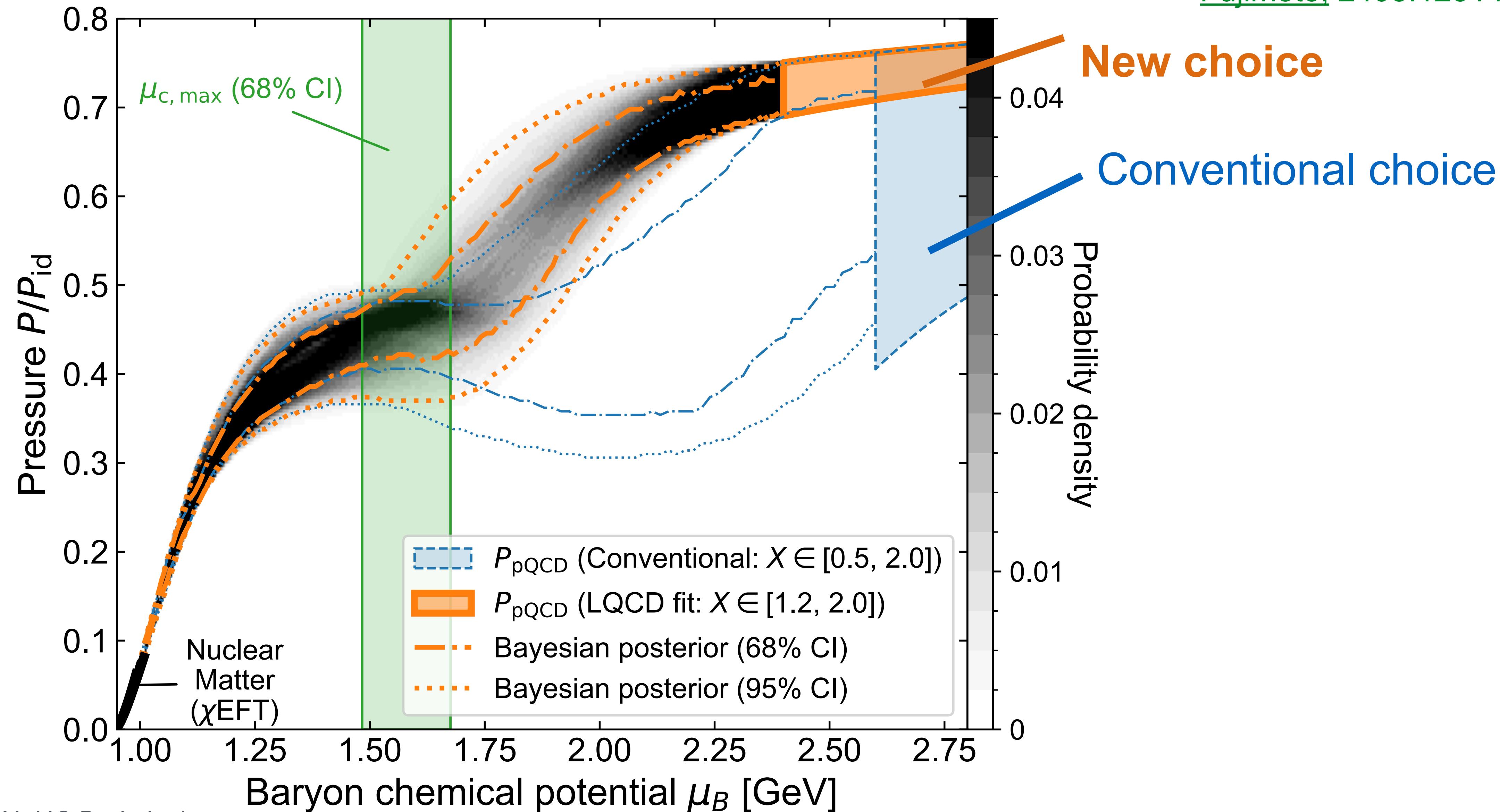
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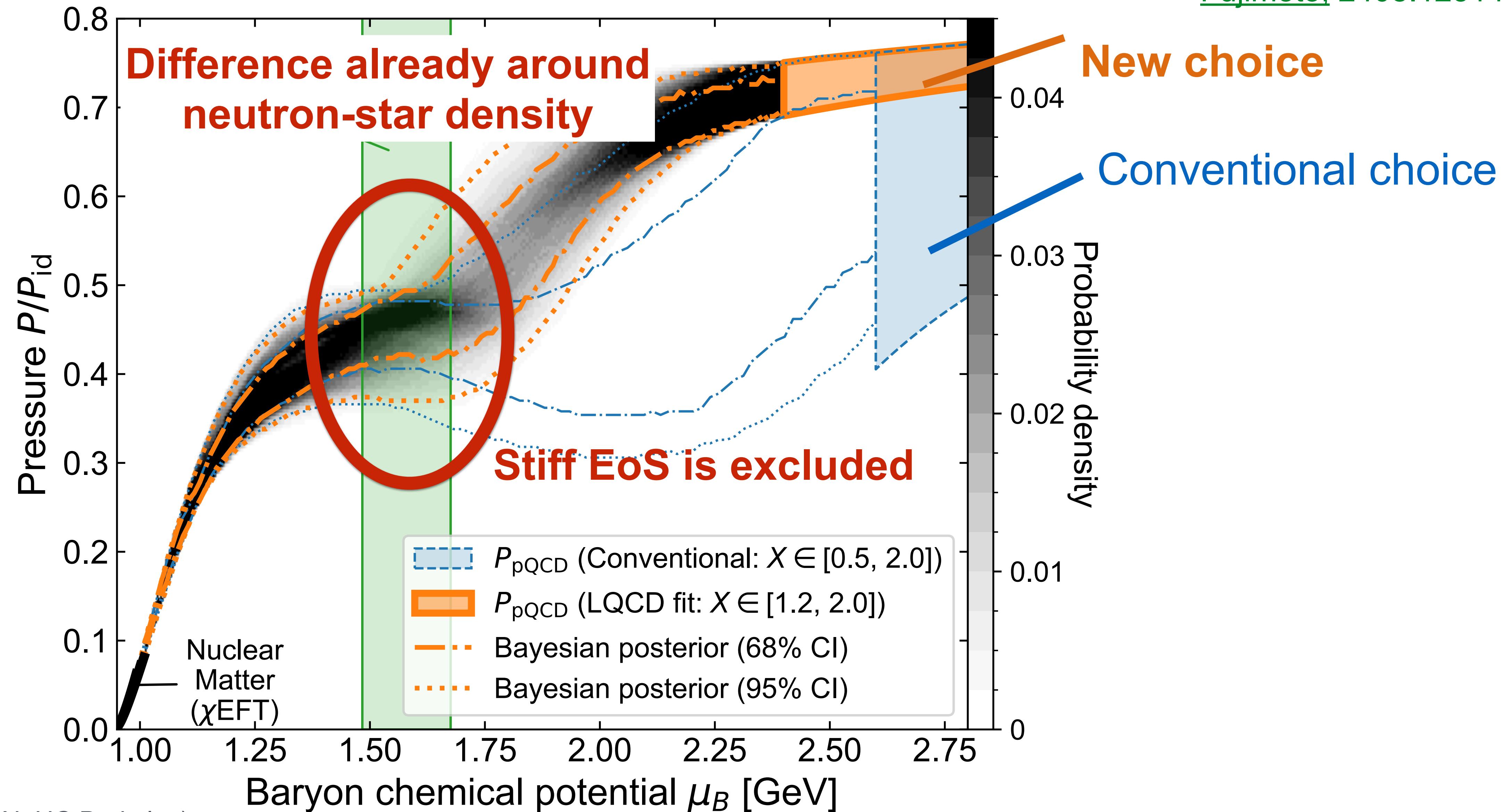
# Effect on $\zeta_{QCD_B}$ : NS phenomenology

Fujimoto, 2408.12514



# Effect on $\xi_{QCD_B}$ : NS phenomenology

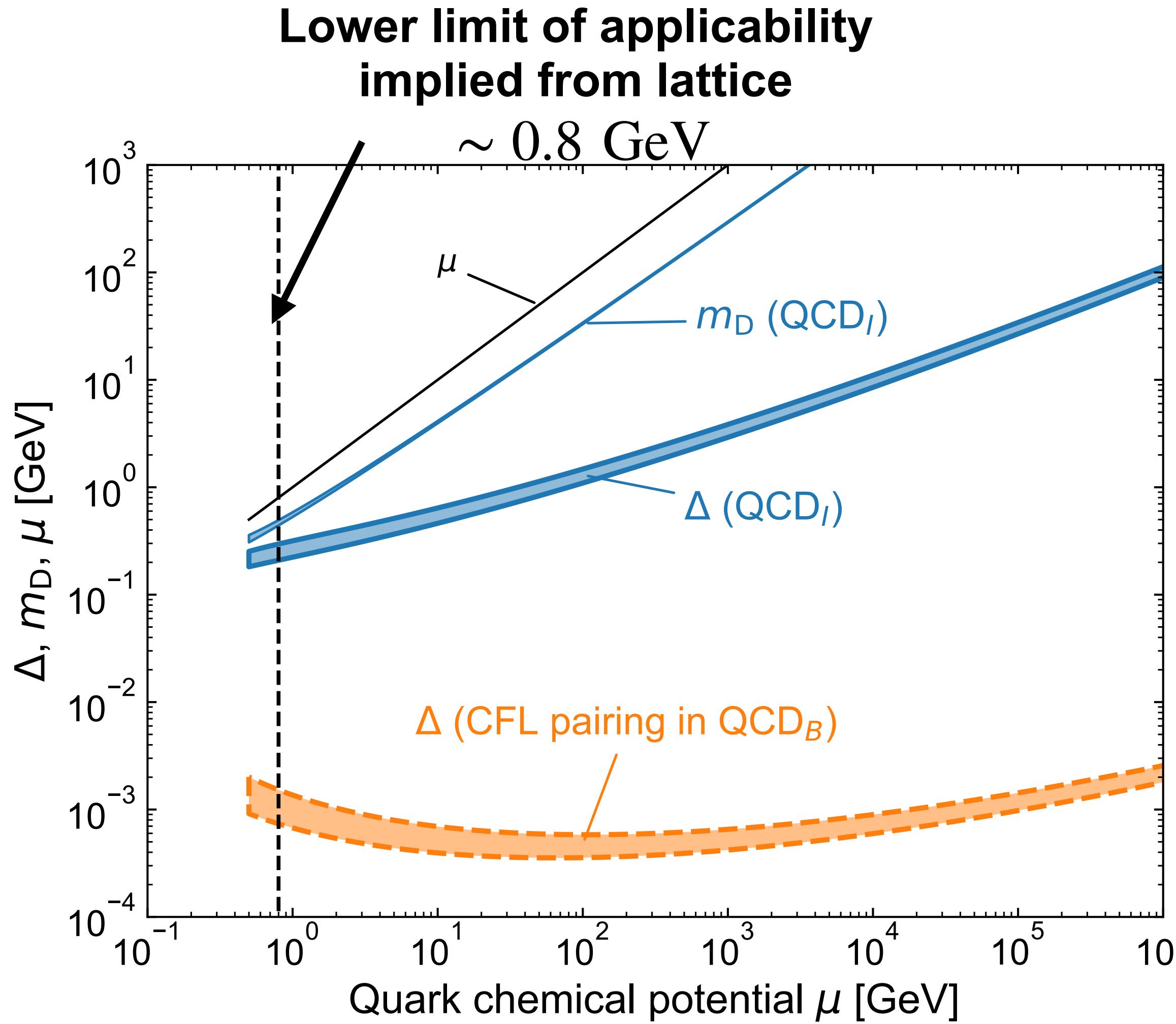
Fujimoto, 2408.12514



# Color superconductivity in weak coupling

Fujimoto, 2408.12514

Plot of weak-coupling Cooper pairing gap:



$$\Delta_{\text{CFL}} \sim 1 \text{ MeV at } \mu = 0.8 \text{ GeV}$$

cf. Kurkela,Rajagopal,Steinhorst (2024)

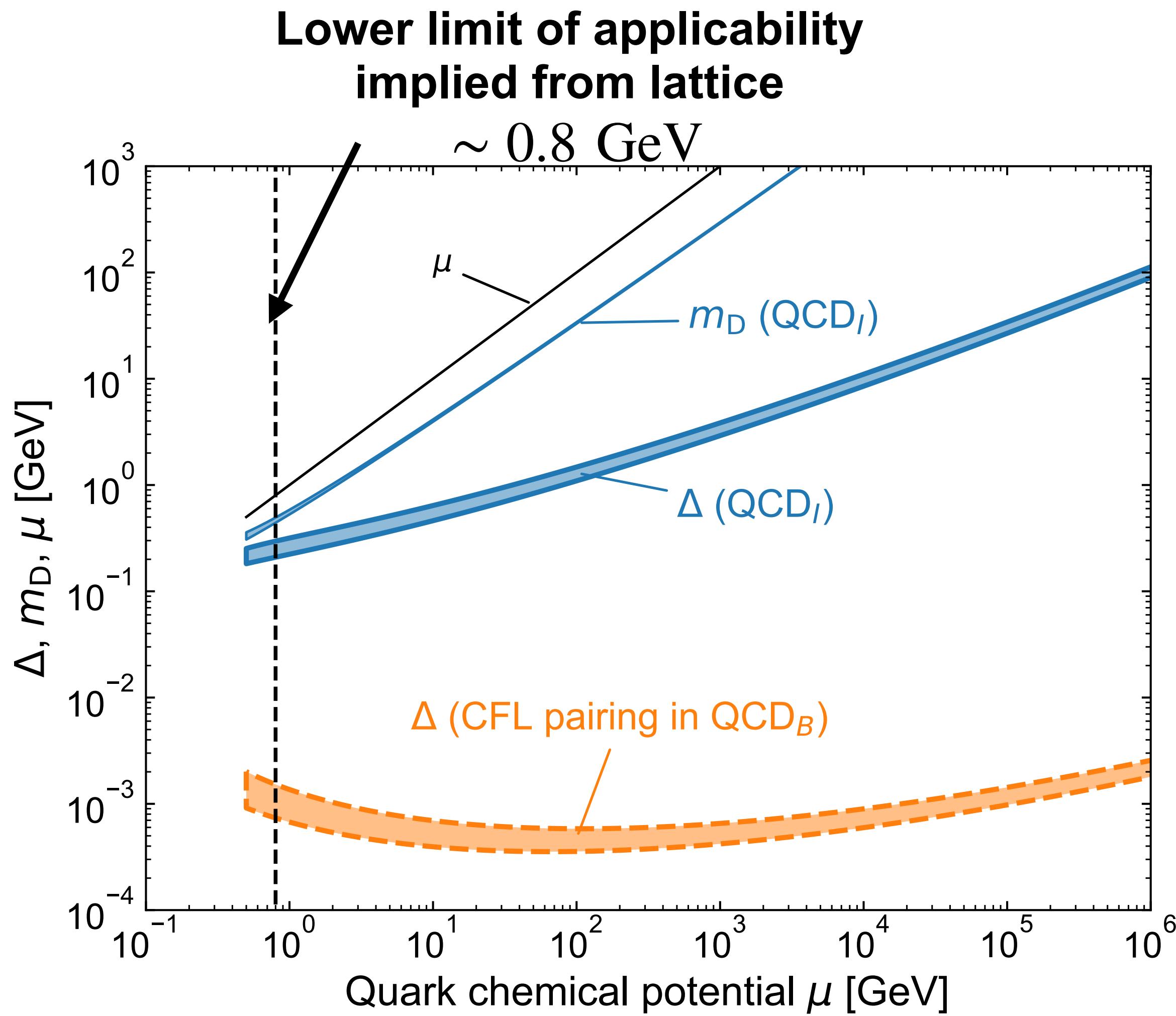
- A negligibly **small** contribution to bulk thermodynamics in weak-coupling

$$P_{\text{QCD}}(\mu) = \frac{3\mu^4}{4\pi^2} [1 - \mathcal{O}(\alpha_s)] + \frac{3\mu^2\Delta^2}{2\pi^2} [1 + \mathcal{O}(\alpha_s)]$$

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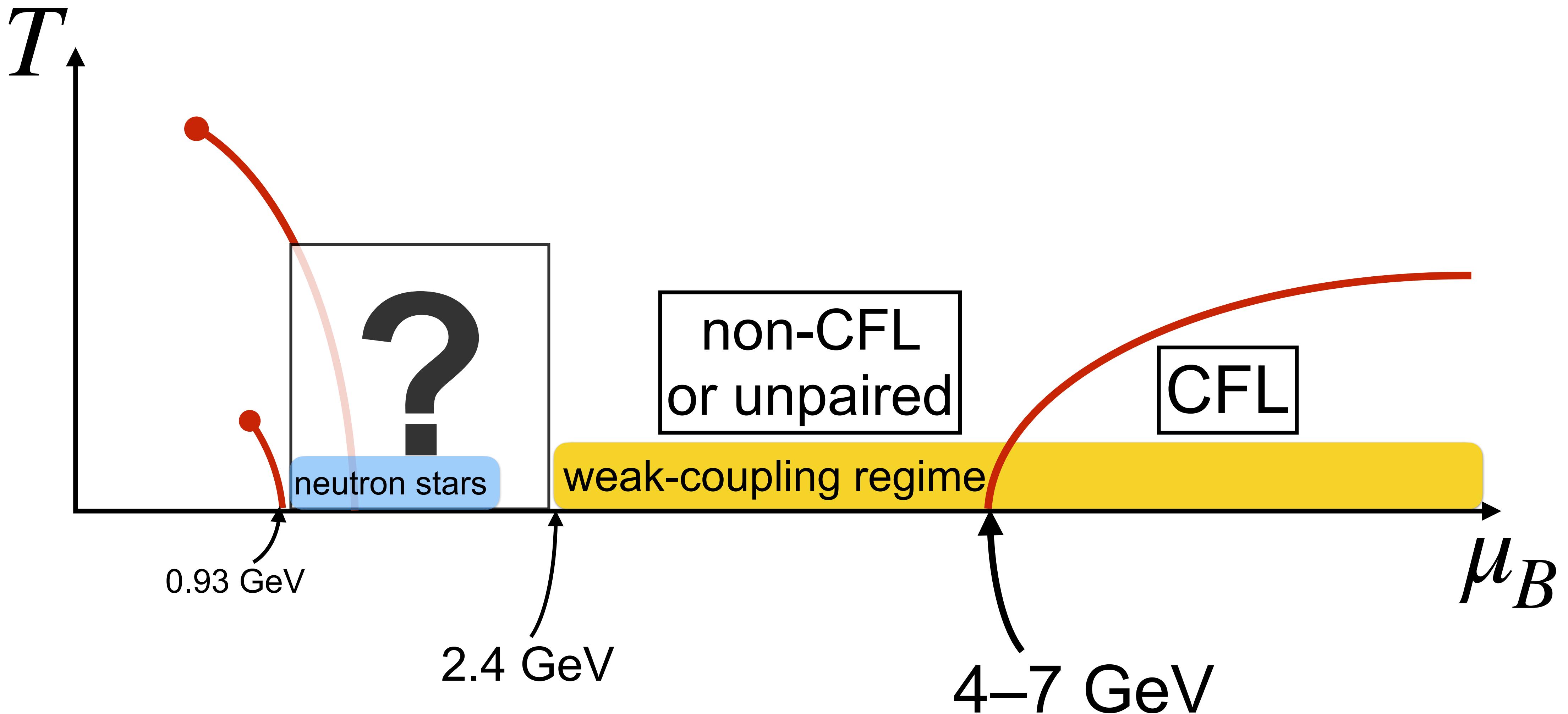
$$\Delta_{\text{CFL}} \sim 1 \text{ MeV at } \mu = 0.8 \text{ GeV}$$

- Comparable to the stress by strange quark mass:  $\Delta_{\text{CFL}} \sim m_s^2/4\mu$
- **CFL may not be the ground state even at  $\mu_B = 2.4 \text{ GeV}$**

NB: CFL and color superconductor may still be realized in NSs due to the nonperturbative enhancement from instantons

Alford,Rajagopal,Wilczek (1997);  
Rapp,Schafer,Shuryak,Velkovsky (1997)

# Effect on the QCD phase diagram



# Summary

- **Trace anomaly:** useful measure for conformal nature of matter.
- The behavior of the trace anomaly may signify that the conformal matter may be realized in the core of neutron stars (possibly quark matter?)
- **QCD at finite isospin density:** useful nonperturbative piece of information
- **Weak-coupling results:** Matches with lattice data at finite isospin density
  - Empirical evidence for the validity down to  $\mu \sim 0.8$  GeV.
  - Color-superconducting gap is negligible in the weak coupling limit.
  - CFL phase may be unstable against unpairing induced by the stress from strange quark mass