# **Exploring chirality structure** in nucleon decay

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  - with
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  - JHEP 01 (2025) 175
  - The Frontier of Particle Physics 2025@YITP







#### 

Decay channel:  $p \rightarrow \pi^0 e^+$  and  $p \rightarrow \eta e^+$ 

#### $\mathcal{L}_{\text{eff}} = C_{RL}^{\ell} \left[ \epsilon_{abc} (u_R^a d_R^b) (u_L^c \ell_L) \right] + C_{LR}^{\ell} \left[ \epsilon_{abc} (u_L^a d_L^b) (u_R^c \ell_R) \right] + C_{LL}^{\ell} \left[ \epsilon_{abc} (u_L^a d_L^b) (u_L^c \ell_L) \right] + C_{RR}^{\ell} \left[ \epsilon_{abc} (u_R^a d_R^b) (u_R^c \ell_R) \right]$



#### Take-home: Nucleon decay → Chirality structure

Decay channel:  $p \rightarrow \pi^0 e^+$  and  $p \rightarrow \eta e^+$ 

$$\mathcal{L}_{\text{eff}} = C_{RL}^{\ell} \left[ \epsilon_{abc} (u_R^a d_R^b) (u_L^c \ell_L) \right] + C_{LR}^{\ell} \left[ \epsilon_{abc} (u_L^a d_L^b) (u_L^c \ell_L) \right]$$





#### $u_R^c \ell_R) \Big] + C_{LL}^\ell \Big[ \epsilon_{abc} (u_L^a d_L^b) (u_L^c \ell_L) \Big] + C_{RR}^\ell \Big[ \epsilon_{abc} (u_R^a d_R^b) (u_R^c \ell_R) \Big]$



### Take-home: Nucleon decay → Chirality structure

Decay channel:  $p \rightarrow \pi^0 e^+$  and  $p \rightarrow \eta e^+$ 

$$\mathcal{L}_{\text{eff}} = C_{RL}^{\ell} \left[ \epsilon_{abc} (u_R^a d_R^b) (u_L^c \ell_L) \right] + C_{LR}^{\ell} \left[ \epsilon_{abc} (u_L^a d_L^b) (u_L^c \ell_L) \right]$$







#### Plan:

- Introduction
- EFT discussion
- Application to SUSY GUTs



#### **Baryon number violation**

Baryon number is expected to be violated in physics beyond the standard model.

• Baryon asymmetry of the universe

Sakharov's three conditions  $\rightarrow B$  violation A. D. Sakharov (1966)

#### **Baryon number violation**

Baryon number is expected to be violated in physics beyond the standard model (SM).

- Baryon asymmetry of the universe Sakharov's three conditions  $\rightarrow B$  violation
- Gauge coupling unification:  $SU(3)_C \otimes SU(2)_L \otimes U(1)_Y \rightarrow SU(5)$

nicely incorporated in Grand Unified Theory (GUT)

- Quark-lepton unification

$$\mathbf{10}_{i} = \frac{1}{\sqrt{2}} \begin{pmatrix} 0 & u_{Ri3}^{\dagger} & -u_{Ri2}^{\dagger} & u_{Li}^{1} & d_{Li}^{1} \\ -u_{Ri3}^{\dagger} & 0 & u_{Ri1}^{\dagger} & u_{Li}^{2} & d_{Li}^{2} \\ u_{Ri2}^{\dagger} & -u_{Ri1}^{\dagger} & 0 & u_{Li}^{3} & d_{Li}^{3} \\ -u_{Li}^{1} & -u_{Li}^{2} & -u_{Li}^{3} & 0 & e_{Ri}^{\dagger} \\ -d_{Li}^{1} & -d_{Li}^{2} & -d_{Li}^{3} & -e_{Ri}^{\dagger} & 0 \end{pmatrix}$$

- B violation

A. D. Sakharov (1966)



#### Nucleon decay

Nucleon decay has been the main probe of baryon-number violation.

#### e.g. $\tau(p \to \pi^0 e^+) > 2.4 \times 10^{34}$ years

<sup>4</sup> years Super-Kamiokande Collaboration (2020)

#### **Nucleon decay**

Nucleon decay has been the main probe of baryon-number violation.

e.g. 
$$\tau(p \to \pi^0 e^+) > 2.4 \times 10^3$$

Next-generation nucleon decay experiments:

• Hyper-Kamiokande (HK)



• JUNO



• DUNE

# DEEP UNDERGROUND NEUTRINO EXPERIMENT

<sup>,4</sup> years Super-Kamiokande Collaboration (2020)





S. Weinberg (1979) F. Wilczek and A. Zee (1979) L. F. Abbott and M. B. Wise (1980)

 $(3)_C \otimes SU(2)_L \otimes U(1)_Y$ 

$$\mathcal{L}_{\mathrm{SM,eff}} = \sum_{I,ijkl} C_{(I)}^{ijkl} \mathcal{O}_{ijkl}^{(I)} + \text{h.c.} \qquad (SU(3))$$

$$\mathcal{O}_{ijkl}^{(1)} = \epsilon_{abc} \epsilon^{\alpha\beta} (u_{Ri}^{a} d_{Rj}^{b}) (Q_{Lk\alpha}^{c} L_{Ll\beta})$$

$$\mathcal{O}_{ijkl}^{(2)} = \epsilon_{abc} \epsilon^{\alpha\beta} (Q_{Li\alpha}^{a} Q_{Lj\beta}^{b}) (u_{Rk}^{c} e_{Rl})$$
Mixed

$$\mathcal{O}_{ijkl}^{(3)} = \epsilon_{abc} \epsilon^{\alpha\beta} \epsilon^{\gamma\delta} (Q_{Li\alpha}^a Q_{Lj\gamma}^b) (Q_{Lk\delta}^c L_{Ll\beta})$$
$$\mathcal{O}_{ijkl}^{(4)} = \epsilon_{abc} (u_{Ri}^a d_{Rj}^b) (u_{Rk}^c e_{Rl})$$



S. Weinberg (1979) F. Wilczek and A. Zee (1979) L. F. Abbott and M. B. Wise (1980)

 $)_C \otimes SU(2)_L \otimes U(1)_Y$ 

d-type

$$\begin{aligned} \mathcal{L}_{\mathrm{SM,eff}} &= \sum_{I,ijkl} C_{(I)}^{ijkl} \mathcal{O}_{ijkl}^{(I)} + \mathrm{h.c.} & \text{SU(3)}_{C} \\ \mathcal{O}_{ijkl}^{(1)} &= \epsilon_{abc} \epsilon^{\alpha\beta} (u_{Ri}^{a} d_{Rj}^{b}) (Q_{Lk\alpha}^{c} L_{Ll\beta}) \\ \mathcal{O}_{ijkl}^{(2)} &= \epsilon_{abc} \epsilon^{\alpha\beta} (Q_{Li\alpha}^{a} Q_{Lj\beta}^{b}) (u_{Rk}^{c} e_{Rl}) \end{aligned} \quad \text{Mixed-} \\ \mathcal{O}_{ijkl}^{(3)} &= \epsilon_{abc} \epsilon^{\alpha\beta} \epsilon^{\gamma\delta} (Q_{Li\alpha}^{a} Q_{Lj\gamma}^{b}) (Q_{Lk\delta}^{c} L_{Ll\beta}) \\ \mathcal{O}_{ijkl}^{(4)} &= \epsilon_{abc} (u_{Ri}^{a} d_{Rj}^{b}) (u_{Rk}^{c} e_{Rl}) \end{aligned} \quad \text{Pure-t}$$



 $\otimes$  SU(2)<sub>L</sub>  $\otimes$  U(1)<sub>Y</sub>

S. Weinberg (1979) F. Wilczek and A. Zee (1979) L. F. Abbott and M. B. Wise (1980)

-type

type

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H. Georgi and S. L. Glashow (1974) Gauge boson exchange

## Dominant in non-SUSY GUTs

![](_page_14_Figure_3.jpeg)

**Dimension-six Kähler** 

Mixed-type

 $C^{ijkl}_{(2)}$ 

 $C^{ijkl}$ 

(1)

H. Georgi and S. L. Glashow (1974) Gauge boson exchange

## Dominant in non-SUSY GUTs

![](_page_15_Figure_3.jpeg)

**Dimension-six Kähler** 

Mixed-type

 $C_{(2)}^{ijkl}$ 

 $C^{ijkl}$ 

S. Weinberg (1982) Color-triplet Higgs exchange N. Sakai and T. Yanagida (1982)

![](_page_15_Figure_6.jpeg)

H. Georgi and S. L. Glashow (1974) Gauge boson exchange

![](_page_16_Figure_2.jpeg)

#### Chirality structure is sensitive to UV physics.

S. Weinberg (1982) Color-triplet Higgs exchange N. Sakai and T. Yanagida (1982)

![](_page_16_Figure_5.jpeg)

#### Chirality structure

Mixed-type Pure-type

#### UV physics

**4** - - - - - **>** 

Mediator, mass spectrum

#### Chirality structure

Mixed-type Pure-type

Q. How to probe chirality structure?

#### UV physics

• - - - - - • •

Mediator, mass spectrum

#### Chirality structure

Mixed-type

Pure-type

Q. How to probe chirality structure?

A. By means of ratios of branching fractions

#### UV physics

#### Mediator, mass spectrum

![](_page_19_Figure_8.jpeg)

![](_page_20_Figure_0.jpeg)

Solution Case 1: mixed-type only case,  $C_{LL}^{\ell} = C_R^{\ell}$  $\Gamma(N \to \Pi \ell^+) = \frac{m_N}{32\pi} \left(1 - \frac{m_{\Pi}^2}{m_N^2}\right)^2 |W_{N\Pi\ell,0}^{LR}|^2$ 

$$T_{RR}^{\ell} = 0$$
, with  $m_{\ell} = 0$ .

$$|^{2}\left[|C_{RL}^{\ell}|^{2}+|C_{LR}^{\ell}|^{2}
ight]$$

![](_page_22_Figure_1.jpeg)

$$T_{RR}^{\ell} = 0$$
, with  $m_{\ell} = 0$ .

$$|^{2}\left[|C_{RL}^{\ell}|^{2}+|C_{LR}^{\ell}|^{2}
ight]$$

Solution Case 1: mixed-type only case, 
$$C_{LL}^{\ell} = C_{RR}^{\ell} = 0$$
, with  $m_{\ell} = 0$ .  
 $\Gamma(N \to \Pi \ell^+) = \frac{m_N}{32\pi} \left(1 - \frac{m_\Pi^2}{m_N^2}\right)^2 \frac{|W_{N\Pi\ell,0}^{LR}|^2}{|W_{N\Pi\ell,0}^{LR}|^2} \left[|C_{RL}^{\ell}|^2 + |C_{LR}^{\ell}|^2\right]$ 

Hadron matrix elements

$$\langle \Pi(\boldsymbol{p}) | \mathcal{O}_{\chi\chi'} | N(\boldsymbol{k}) \rangle = P_{\chi'} \left[ \frac{W_0^{\mathcal{O}}(q^2) + \frac{\not q}{m_N} W_1^{\mathcal{O}}(q^2) \right] u_N(\boldsymbol{k}) \qquad \mathcal{O}_{\chi\chi'} = \epsilon_{abc} (q_\chi^a q_\chi'^b) q_{\chi'}'^c$$

$$q^\mu = k^\mu - p^\mu$$

**Case 1: mixed-type** only case,  $C_{LL}^{\ell} = C_{R}^{\ell}$  $\Gamma(N \to \Pi \ell^+) = \frac{m_N}{32\pi} \left( 1 - \frac{m_\Pi^2}{m_N^2} \right)^2 |W_{N\Pi\ell,0}^{LR}|^2$ 

Hadron matrix elements

$$\langle \Pi(\boldsymbol{p}) | \mathcal{O}_{\boldsymbol{\chi}\boldsymbol{\chi}'} | N(\boldsymbol{k}) \rangle = P_{\boldsymbol{\chi}'} \left[ \frac{W_0^{\mathcal{O}}(q^2)}{m_N} + \frac{\not q}{m_N} W_1^{\mathcal{O}}(q^2) \right] u_N(\boldsymbol{k}) \qquad \mathcal{O}_{\boldsymbol{\chi}\boldsymbol{\chi}'} = \epsilon_{abc} (q_{\boldsymbol{\chi}}^a q_{\boldsymbol{\chi}}'^b) q_{\boldsymbol{\chi}'}''^c$$

We use the form factors obtained by <u>lattice simulations</u>

$$f_{RR}^{\ell} = 0$$
, with  $m_{\ell} = 0$ .

$${}^{2}\left[|C_{RL}^{\ell}|^{2}+|C_{LR}^{\ell}|^{2}\right]$$

Y. Aoki et al. (2017) J. S. Yoo, et al. (2022)

Uncertainty:  $\mathcal{O}(10)$  % 

**Case 1:** mixed-type only case,  $C_{LL}^{\ell} = C_{R}^{\ell}$  $\Gamma(N \to \Pi \ell^+) = \frac{m_N}{32\pi} \left( 1 - \frac{m_\Pi^2}{m_M^2} \right)^2 |W_{N\Pi\ell,0}^{LR}|^2$ 

- Wilson coefficients which reflect UV physics.

$$m_{RR}^{\ell}=0$$
, with  $m_{\ell}^{\prime}=0$ .

$${}^{2}\left[|C_{RL}^{\ell}|^{2} + |C_{LR}^{\ell}|^{2}\right]$$

• Does not depend on whether  $p \to \pi^0 e^+$  or  $p \to \eta e^+$ 

**Case 1: mixed-type** only case,  $C_{LL}^{\ell} = C_{LL}^{\ell}$  $\Gamma(N \to \Pi \ell^+) = \frac{m_N}{32\pi} \left( 1 - \frac{m_\Pi^2}{m_N^2} \right)^2 |W_{N\Pi\ell,0}^{LR}|^2 \left[ |C_{RL}^{\ell}|^2 + |C_{LR}^{\ell}|^2 \right]$ The ratio  $\Gamma(p \to \eta e^+)/\Gamma(p \to \pi^0 e^+)$  are solely determined by low-energy quantities.  $\frac{\Gamma(p \to \eta e^+)}{\Gamma(p \to \pi^0 e^+)} = \frac{(1 - m_\eta^2 / m_p^2)^2}{(1 - m_\pi^2 / m_p^2)^2} \cdot \frac{|W_{p\eta\ell}^{\chi\chi'}|^2}{|W_{n\pi^0\ell}^{\chi\chi'}|^2}$ 

$$m_{RR}^{\ell} = 0$$
, with  $m_{\ell} = 0$ .

$$\simeq \begin{cases} 0.0013 & (Mixed only) \\ 0.51 & (Pure only) \end{cases}$$

$$\begin{array}{l} \textcircledlength{\textcircled[]{\label{eq:Case 1: mixed-type only case, } C_{LL}^{\ell} = C_{RR}^{\ell} = 0, \mbox{ with } m_{\ell} = 0. \\ & \Gamma(N \to \Pi \ell^+) = \frac{m_N}{32\pi} \left( 1 - \frac{m_\Pi^2}{m_N^2} \right)^2 |W_{N\Pi\ell,0}^{LR}|^2 \left[ |C_{RL}^{\ell}|^2 + |C_{LR}^{\ell}|^2 \right] \\ \textcircledlength{\textcircled[]{\label{eq:Case 1: mixed-type only case, } C_{IL}^{\ell} = \pi_N^0 e^+ \end{tabular}} \\ & \Gamma(N \to \Pi \ell^+) = \frac{m_N}{32\pi} \left( 1 - \frac{m_\Pi^2}{m_N^2} \right)^2 \cdot \frac{|W_{N\Pi\ell,0}^{LR'}|^2}{|W_{p\eta\ell}^{\chi\chi'}|^2} \simeq \begin{cases} 0.0013 & (\text{Mixed only}) \\ 0.51 & (\text{Pure only}) \end{cases} \\ & \bullet \text{ Parity invariance of QCD plays the role:} \end{cases} \\ & \overline{W_{N\Pi\ell,I}^{RL} = W_{N\Pi\ell,I}^{LR}}, \qquad W_{N\Pi\ell,I}^{LL} = W_{N\Pi\ell,I}^{RR} \end{cases} \quad \langle \Pi(p) | \mathcal{O}_{\chi\chi'} | N(k) \rangle = I \end{cases} \end{array}$$

- low-energy quantities.

$$\left\langle \Pi(\boldsymbol{p}) | \mathcal{O}_{\boldsymbol{\chi}\boldsymbol{\chi}'} | N(\boldsymbol{k}) \right\rangle = P_{\boldsymbol{\chi}'} \left[ W_0^{\mathcal{O}}(q^2) + \frac{\not q}{m_N} W_1^{\mathcal{O}}(q^2) \right] u_N(\boldsymbol{k})$$

![](_page_27_Picture_6.jpeg)

![](_page_28_Figure_1.jpeg)

![](_page_29_Figure_1.jpeg)

![](_page_30_Figure_1.jpeg)

![](_page_31_Figure_1.jpeg)

#### UV physics

#### Mediator, mass spectrum

## The ratios of branching fractions

#### Chirality structure

Mixed-type

Pure-type

![](_page_32_Figure_5.jpeg)

H. Georgi and S. L. Glashow (1974) Gauge boson exchange

![](_page_33_Figure_2.jpeg)

S. Weinberg (1982) Color-triplet Higgs exchange N. Sakai and T. Yanagida (1982)

#### Dominant in **SUSY GUTs**

#### Dimension-five superpotential

![](_page_33_Figure_6.jpeg)

 $M_{H_C} M_{\rm SUSY}^2$ 

•  $M_{\rm SUSY}$ : sfermion mass •  $m_{\tilde{g}} \leq M_{\mathrm{SUSY}}$ 

![](_page_33_Figure_9.jpeg)

#### The minimal SU(5) with high-scale SUSY No sfermion flavor violation

![](_page_34_Figure_1.jpeg)

 $\Gamma(p \to \eta \mu^+)/\Gamma(p \to \pi^0 \mu^+)$ : wino-exchange contributes (pure-type) as  $M_{H_c}$  decreases.

![](_page_34_Picture_3.jpeg)

- $M_X = 10^{16} \, \mathrm{GeV}$   $M_2 = 1 \, \mathrm{TeV}$
- $M_3 = 10 \, {\rm TeV}$
- $M_{\rm SUSY} = 100 \,{\rm TeV}$
- $\tan \beta = 3$

#### Mini-split type mass spectrum

L. J. Hall, Y. Nomura and S. Shirai (2012) M. Ibe, S. Matsumoto and T. T. Yanagida (2012)

- $--- \Gamma(p \rightarrow \eta \mu^{+}) / \Gamma(p \rightarrow \pi^{0} \mu^{+})$ 
  - $\cdots \Gamma(n \to \eta \bar{\nu}) / \Gamma(n \to \pi^0 \bar{\nu})$
  - $- \Gamma(n \rightarrow \pi^0 \bar{\nu}) / \Gamma(p \rightarrow \pi^0 e^+)$
  - $\Gamma(n \rightarrow \pi^0 \bar{\nu}) / \Gamma(p \rightarrow \pi^0 \mu^+)$

![](_page_34_Figure_16.jpeg)

![](_page_35_Figure_1.jpeg)

![](_page_35_Picture_3.jpeg)

#### No sfermion flavor violation

![](_page_36_Figure_1.jpeg)

#### No sfermion flavor violation

![](_page_37_Figure_1.jpeg)

#### No sfermion flavor violation

![](_page_38_Figure_9.jpeg)

![](_page_38_Picture_10.jpeg)

![](_page_39_Figure_1.jpeg)

![](_page_40_Figure_1.jpeg)

![](_page_41_Picture_0.jpeg)

## The ratios of nucleon decay branching fractions are sensitive to UV physics.

![](_page_41_Figure_2.jpeg)

![](_page_42_Picture_0.jpeg)

### **SUSY dimension-five nucleon decay operators**

![](_page_43_Figure_1.jpeg)

The dimension-five contribution is proportional to  $m_{\tilde{g}}/(M_{H_C}M_{SUSY}^2)$ 

P. Nath and R. L. Arnowitt (1988) J Hisano, H. Murayama and T. Yanagida (1993)

![](_page_43_Picture_4.jpeg)

### **Current bounds and future prospects**

Decay Mode	Current [years]	HK sensitivity
$p \to \pi^0 e^+$	$2.4 \times 10^{34}$ [25]	$7.8  imes 10^{34}$ [11]
$p \to \pi^0 \mu^+$	$1.6 \times 10^{34}$ [25]	$7.7  imes 10^{34} \ [11]$
$p  ightarrow \eta e^+$	$1.0 \times 10^{34}$ [26]	$4.3  imes 10^{34} \ [11]$
$p  ightarrow \eta \mu^+$	$4.7 \times 10^{33}$ [26]	$4.9  imes 10^{34} \ [11]$
$p \to \pi^+ \bar{\nu}$	$3.9  imes 10^{32}$ [27]	
$n  ightarrow \pi^- e^+$	$5.3 \times 10^{33}$ [26]	$2.0  imes 10^{34} [11]$
$n  ightarrow \pi^- \mu^+$	$3.5 \times 10^{33}$ [26]	$1.8  imes 10^{34} \ [11]$
$n  ightarrow \pi^0 \bar{ u}$	$1.1 \times 10^{33}$ [27]	
$n  ightarrow \eta ar{ u}$	$1.6 \times 10^{32}$ [28]	

[11]: Hyper-Kamiokande Collaboration [arXiv: 1805.04163] [25]: Super-Kamiokande Collaboration [arXiv: 2010.16098] [26]: Super-Kamiokande Collaboration [arXiv: 1705.07221] [27]: Super-Kamiokande Collaboration [arXiv: 1305.4391] [28]: C. McGrew et al. (1999)

- [years]
- Units:  $10^{33}$  years
- ▶ 90% CL
- 1.9 Megaton-year exposure is assumed for the prospect.

![](_page_44_Picture_8.jpeg)

#### **Event Reconstruction**

![](_page_45_Figure_1.jpeg)

Eta mass reconstruction at SK

![](_page_45_Figure_3.jpeg)

- Left:  $\eta \rightarrow 2\gamma$ , branching ratio = 39%
- Right:  $\eta \rightarrow 3\pi^0$ , branching ratio = 33%
- Open histogram: Monte-Carlo events
- Hatched histogram:
  - left: true  $\eta \rightarrow 2\gamma$
  - right: true  $\eta \rightarrow 3\pi^0$

![](_page_45_Picture_11.jpeg)

- Assume the tree-level exchange of a scalar or vector boson -
  - Vector boson  $V_{\mu}$  can induce  $\mathcal{O}_{ijk\ell}^{(1)}$ ,  $\mathcal{O}_{ijk\ell}^{(2)}$ Renormalizable interaction:  $V_{\mu}\psi_{R\dot{\alpha}}\bar{\sigma}^{\mu\dot{\alpha}\alpha}\chi_{L\alpha}$   $\psi_{R}$  include the conjugate of  $\psi_{L}$ Scalar boson *S* can induce  $\mathcal{O}_{ijk\ell}^{(1)}$ ,  $\mathcal{O}_{ijk\ell}^{(2)}$ ,  $\mathcal{O}_{ijk\ell}^{(3)}$ ,  $\mathcal{O}_{ijk\ell}^{(4)}$ . Renormalizable interactions:  $S(\psi_L \chi_L)$ ,  $S(\psi_R \chi_R)$
- In non-SUSY GUTs, the gauge boson exchange typically dominates: mixed-type
- In SUSY-GUT, the one-loop contribution can be significant: both can contribute

S. Weinberg (1979)

D. V. Nanopoulos and S. Weinberg (1979)

$$\mathcal{O}_{ijk\ell}^{(1)} \equiv \epsilon_{abc} \epsilon_{\alpha\beta} \left( u_{Ri}^{a} d_{Rj}^{b} \right) \left( Q_{k}^{c\alpha} L_{\ell}^{\beta} \right) \\\mathcal{O}_{ijk\ell}^{(2)} \equiv \epsilon_{abc} \epsilon_{\alpha\beta} \left( Q_{i}^{a\alpha} Q_{j}^{b\beta} \right) \left( u_{Rk}^{c} e_{R} \right) \\\mathcal{O}_{ijk\ell}^{(3)} \equiv \epsilon_{abc} \epsilon_{\alpha\beta} \epsilon_{\gamma\delta} \left( Q_{i}^{a\alpha} Q_{j}^{b\gamma} \right) \left( Q_{k}^{c\delta} \right) \\\mathcal{O}_{ijk\ell}^{(4)} \equiv \epsilon_{abc} \left( u_{Ri}^{a} d_{Rj}^{b} \right) \left( u_{Rk}^{c} e_{R\ell} \right),$$

![](_page_46_Picture_10.jpeg)

#### The direct method

We use hadron matrix elements evaluated by lattice simulations.

![](_page_47_Figure_3.jpeg)

$$\langle \Pi(\boldsymbol{p}) | \mathcal{O}_{\chi\chi'} | N(\boldsymbol{k}) \rangle = P_{\chi'} \left[ W_0^{\mathcal{O}}(q^2) + \frac{\not q}{m_N} W_1^{\mathcal{O}}(q^2) \right] u_N(\boldsymbol{k}) \qquad \mathcal{O}_{\chi\chi'} = \epsilon_{abc} (q_\chi^a q_\chi'^b) q_{\chi'}'^{\prime c}$$

Y. Aoki et al. (2017) J. S. Yoo, et al. (2022)

	1		
_	LR	LL	
51	0.169	-0.134	
13	0.044	-0.044	

• 
$$W^{RR} = W^{LL}, W^{RL} = W^{LL}$$

• In units of  $\text{GeV}^2$ 

•  $\mathcal{O}(m_{\ell}/m_p) \rightarrow \text{only relevant for anti-muon}$ 

![](_page_47_Picture_11.jpeg)

## Positron channels: $\Gamma(p \rightarrow \eta e^+)/\Gamma(p \rightarrow \pi^0 e^+)$

To understand the behavior, let us use the expressions calculated from chiral lagrangian (i.e., the so-called indirect method

$$\begin{split} \Gamma(p \to \pi^0 e^+) &= \frac{m_p}{32\pi} \left( 1 - \frac{m_\pi^2}{m_p^2} \right)^2 \frac{(1+D+F)^2 \alpha^2}{2f^2} \left[ |C_{RL}^e - C_{LL}^e|^2 + |C_{LR}^e - C_{RR}^e|^2 \right] \\ \Gamma(p \to \eta e^+) &= \frac{m_p}{32\pi} \left( 1 - \frac{m_\eta^2}{m_p^2} \right)^2 \frac{\alpha^2}{6f^2} \left[ |(1+D-3F)C_{RL}^e + (3-D+3F)C_{LL}^e|^2 + |(1+D-3F)C_{LR}^e + (3-D+3F)C_{RR}^e|^2 \right] \\ &+ \left| (1+D-3F)C_{LR}^e + (3-D+3F)C_{RR}^e \right|^2 \right], \end{split}$$

- $p \rightarrow \pi^{\vee} e'$  is suppressed when  $C_{LL}^e / C_{RL}^e = C_{RR}^e / C_{LR}^e = 1$

M. Claudson, M. B. Wise, and L. J. Hall (1982)

•  $p \to \eta e^+$  is suppressed when  $C_{II}^e / C_{RI}^e = C_{RR}^e / C_{IR}^e = -(1 + D - 3F) / (3 - D + 3F) \simeq -0.1$ 

![](_page_48_Figure_8.jpeg)

![](_page_49_Figure_0.jpeg)

![](_page_49_Picture_3.jpeg)

![](_page_50_Figure_0.jpeg)

## Anti-muon channels: $\Gamma(p \rightarrow \eta \mu^+)/\Gamma(p \rightarrow \pi^0 \mu^+)$

Solution For the mixed-only case, up to  $\mathcal{O}(m_{\mu}/m_{p})$ ,  $\Gamma(p \to \Pi \mu^+) = \frac{m_p}{32\pi} \left( 1 - \frac{m_\Pi^2}{m_p^2} \right)^2 \left( W_{p\Pi\ell,0}^{LR} \right)^2 \left[ |C_{RL}^{\mu}|^2 + |C_{LR}^{\mu}|^2 \right] ,$  $\times \left[1 + \frac{4m_{\mu}}{m_{p}} \left\{ \frac{W_{p\Pi\ell,1}^{LR}}{W_{n\Pi\ell,0}^{LR}} + \left(1 - \frac{m_{\Pi}^{2}}{m_{n}^{2}}\right)^{-1} \right\} \frac{\operatorname{Re}(C_{RL}^{\mu}C_{LR}^{\mu*})}{|C_{RL}^{\mu}|^{2} + |C_{LR}^{\mu}|^{2}} \right]$ 

- The source of the dependence on  $|C_{LR}^e/C_{RL}^e|$ • This effect is enhanced when  $|W_{p\Pi\mu,1}^{LR}| \gg |W_{p\Pi\mu,0}^{LR}|$

![](_page_51_Figure_4.jpeg)

Form factors obtained by lattice QCD in units of  $\text{GeV}^2$ 

	1	
LL	LR	$\operatorname{LL}$
151	0.169	-0.134
.113	0.044	-0.044

Y. Aoki et al. (2017) J. S. Yoo, et al. (2022)

#### **Other channels**

or pure-only cases.

- same applies to  $\Gamma(n \to \pi^0 \bar{\nu}) / \Gamma(p \to \pi^0 \ell^+)$ .
- models.

The ratio of neutrino channels is related to that of charged leptons in either mixed-

Due to the presence of more Wilson coefficients (=six for the neutrino ratio), it is difficult to extract information about the Wilson coefficients on generic grounds. The

Those ratios can still become powerful probes when considering specific UV

![](_page_52_Picture_8.jpeg)

#### **High-scale SUSY**

![](_page_53_Figure_1.jpeg)

![](_page_53_Figure_2.jpeg)

M. Ibe, S. Matsumoto, T. T. Yanagida (2012) e and an eta

#### For $M_{\rm SUSY} = 10^2$ TeV, tan $\beta \sim 3$ is needed to reproduce the correct Higgs mass.

![](_page_54_Figure_1.jpeg)

![](_page_54_Picture_2.jpeg)

![](_page_55_Figure_1.jpeg)

![](_page_55_Picture_3.jpeg)

![](_page_56_Figure_1.jpeg)

![](_page_57_Figure_1.jpeg)

 $\Gamma(p \rightarrow \eta e^+)/\Gamma(p \rightarrow \pi^0 e^+)$ : gauge-boson exchange (mixed-type) dominates even when the GUT gauge boson is heavy.

• 
$$M_X = 10^{17} \text{ GeV}$$
  
•  $M_{H_C} = 10^{16} \text{ GeV}$   
•  $M_2 = 1 \text{ TeV}$   
•  $M_3 = 10 \text{ TeV}$   
•  $M_{SUSY} = 100 \text{ TeV}$   
•  $\tan \beta = 3$ 

![](_page_58_Figure_1.jpeg)

 $\Gamma(p \rightarrow \eta \mu^+)/\Gamma(p \rightarrow \pi^0 \mu^+)$ : wino contribution (mixed-type) dominates regardless of the higgsino mass.

 $\rightarrow$  this ratio is sensitive to dim-5 wino v.s. the GUT gauge boson competition.

![](_page_59_Figure_1.jpeg)

 $\Gamma(n \to \eta \bar{\nu})/\Gamma(n \to \pi^0 \bar{\nu})$ : higgsino contribution (mixed-type) becomes larger as its mass increases.

 $\rightarrow$  this ratio is most useful to discriminate the higgsino contribution from that of wino.

![](_page_60_Figure_1.jpeg)

![](_page_61_Figure_1.jpeg)

•  $M_X = 10^{16}$  GeV,  $M_2 = 1$  TeV,  $M_3 = 10$  TeV,  $M_{SUSY} = 100$  TeV,  $\tan \beta = 3$ 

![](_page_62_Figure_1.jpeg)

•  $M_X = 10^{17} \text{ GeV}, \ M_{H_C} = 10^{16} \text{ GeV}, \ M_2 = 1 \text{ TeV}, \ M_3 = 10 \text{ TeV}, \ M_{SUSY} = 100 \text{ TeV}, \ \tan \beta = 3$ 

#### **Sfermion flavor violation**

![](_page_63_Figure_1.jpeg)

 $M_X = 10^{17} \text{ GeV}, M_{H_C} = 10^{16} \text{ GeV}, \mu_H = 200 \text{ GeV},$  $M_1 = 5$  TeV,  $M_2 = 1$  TeV,  $M_3 = 10$  TeV,  $M_{SUSY} = 100$  TeV,  $\tan \beta = 3$ 

#### In the presence of flavor violation,

- various channels could be accessible by upcoming experiments, or
- even ruled out by current experimental limits for large  $\delta_{13}^{\tilde{f}}$

 $\rightarrow$  can readily be avoided if  $M_{\text{SUSY}}$  or  $M_{H_C}$  is larger

$$\widetilde{m}_{\tilde{f}}^2 = M_{\rm SUSY}^2 \begin{pmatrix} 1 & \delta_{12}^{\tilde{f}} & \delta_{13}^{\tilde{f}} \\ \delta_{12}^{\tilde{f}*} & 1 & \delta_{23}^{\tilde{f}} \\ \delta_{13}^{\tilde{f}*} & \delta_{23}^{\tilde{f}*} & 1 \end{pmatrix}$$

![](_page_63_Picture_8.jpeg)

![](_page_63_Picture_9.jpeg)

#### **Sfermion flavor violation**

![](_page_64_Figure_1.jpeg)

 $M_X = 10^{17} \text{ GeV}, M_{H_C} = 10^{16} \text{ GeV}, \mu_H = 200 \text{ GeV},$  $M_1 = 5$  TeV,  $M_2 = 1$  TeV,  $M_3 = 10$  TeV,  $M_{SUSY} = 100$  TeV,  $\tan \beta = 3$ 

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![](_page_64_Picture_8.jpeg)

![](_page_64_Picture_9.jpeg)

#### **Uncertainties in hadron matrix elements**

- As we have seen in Aoki san's talk, the uncertainties in the nucleon decay matrix elements obtained by the lattice QCD calculation are currently  $\sim 10 \,\%$ .
- At leading order in the chiral perturbation,

$$\begin{split} W_{p\pi^{+}\nu,0}^{LR} &= \frac{1+D+F}{f} \alpha , \qquad W_{p\pi^{+}\nu,1}^{LR} = -\frac{2(D+F)}{f} \alpha , \\ W_{p\pi^{+}\nu,0}^{LL} &= \frac{1+D+F}{f} \beta , \qquad W_{p\pi^{+}\nu,1}^{LL} = -\frac{2(D+F)}{f} \beta , \\ W_{p\pi^{\ell},0}^{LR} &= -\frac{1+D-3F}{\sqrt{6}f} \alpha , \qquad W_{p\pi^{\ell},1}^{LR} = \frac{2(D-3F)}{\sqrt{6}f} \alpha , \\ W_{p\eta^{\ell},0}^{LL} &= \frac{3-D+3F}{\sqrt{6}f} \beta , \qquad W_{p\eta^{\ell},1}^{LL} = \frac{2(D-3F)}{\sqrt{6}f} \beta , \\ \langle 0|\epsilon_{abc}(u_{R}^{a}d_{R}^{b})u_{L}^{c}|p\rangle \equiv \alpha P_{L}u_{p} , \\ \langle 0|\epsilon_{abc}(u_{L}^{a}d_{L}^{b})u_{L}^{c}|p\rangle \equiv \beta P_{L}u_{p} , \end{split}$$

We expect the uncertainties to be significantly reduced if, instead of the matrix elements themselves, the ratios of them are estimated directly by the calculation. This is because each matrix element is not independent, but correlated with each other.

 $\alpha = -0.01257(111) \text{ GeV}^3$ ,  $\beta = 0.01269(107) \text{ GeV}^3$ .  $\alpha \simeq -\beta$  to ~ 1 %

If we take their ratios,  $f, \alpha, \beta$  will cancel out.

# **Comments on** $W^{\chi\chi'}$

the following combination is evaluated in Y. Aoki et al. (2017).  $W^{\chi\chi'}_{p\eta\ell,\mu} \equiv W^{\chi\chi'}_{p\eta\ell,0} + \frac{m_{\mu}}{m_{N}} W^{\chi\chi'}_{p\eta\ell,1}$ We extract  $W_{pn\ell,1}^{\chi\chi'}(0)$  from  $W_{pn\ell,\mu}^{\chi\chi'}(m_{\mu}^2)$  by using  $W_{pn\ell,0}^{\chi\chi'}(0)$ , neglecting  $m_{\mu}^2$ dependence.

- In previous lattice simulations, the values of  $W_{pn\ell,1}^{\chi\chi'}(0)$  were not estimated. However,