

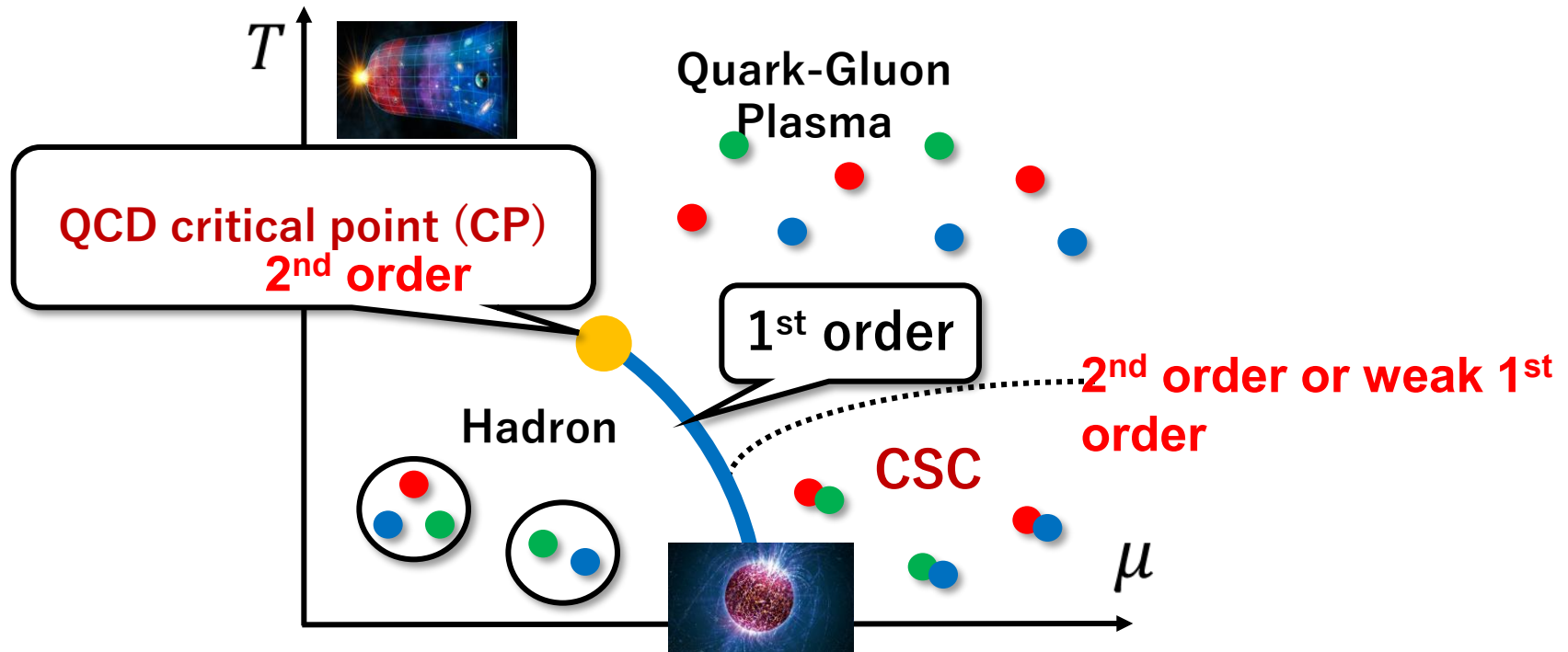
Nature of the soft mode of QCD critical point and Its effects on dilepton production rate and electric conductivity

Teiji Kunihiro (YITP, Kyoto)

Based on joint works with M. Kitazawa (YITP) and **T. Nishimura** (Osaka)

YITPWS 'QCD Critical Point and Hydrodynamic Evolution',
June-1-4, @Maskawa Hall, YITP, Kyoto U.

QCD phase diagram



How can we reveal this structure of the phase diagram, in particular the existence of QCD CP, experimentally?

The notion of the soft modes of 2nd-order ph.tr. and its detectability

The soft mode:

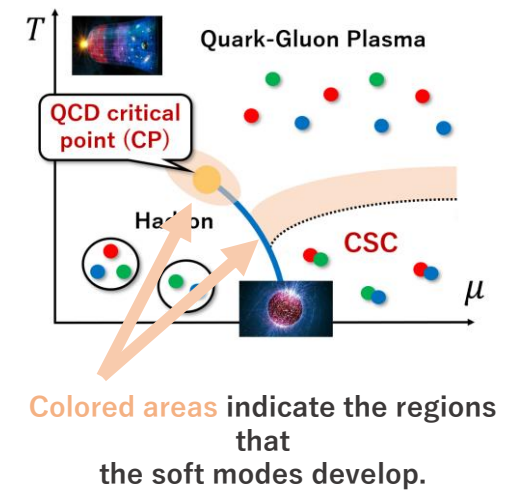
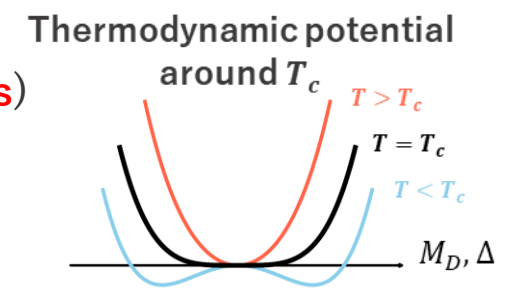
Quantum **amplitude fluctuations of the order parameters**
due to **2nd-order** phase transitions

- CSC : fluctuation of **diquark condensates (preformed pairs)**
Kitazawa, Koide, Kunihiro, Nemoto (2005)
- QCD CP : fluctuation of chiral condensate (σ')
Fujii, Ohtani (2004); Dam T. Son and M. Stephanov (2004)

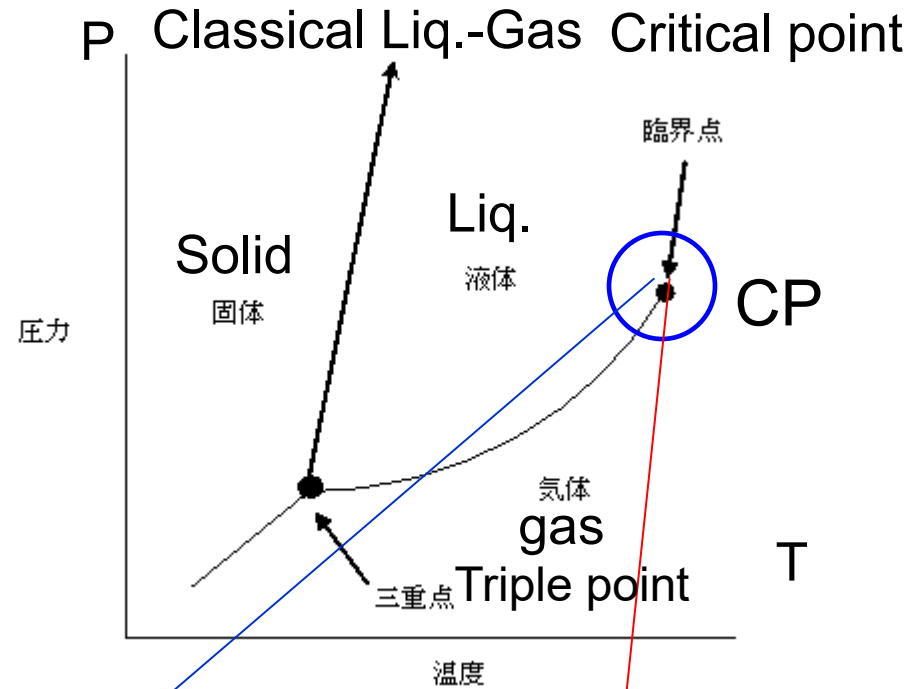
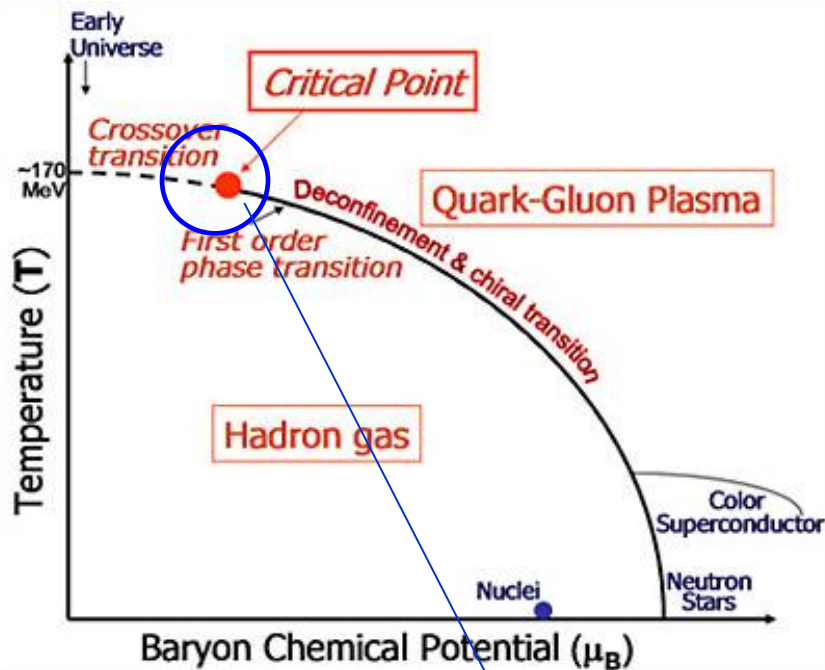
- ✓ They develop prominently around the critical point or line.
- ✓ They are excited with low energy & momentum (< 300 MeV).

These soft modes can contribute significantly to observables.

If the contributions are observed in HIC, they will provide experimental signals of QCD CP as well as CSC.



The case of QCD critical point



The same universality class; Z2

H. Fujii, PRD 67 (03) 094018; H. Fujii and M. Ohtani, Phys.Rev.D70(2004)
Dam. T. Son and M. A. Stephanov, PRD70 ('04) 056001

Collective p-h excitations is the soft mode of QCD critical point!

In classical liquid, a random light scattering due to density fluctuations
(critical opalescence)

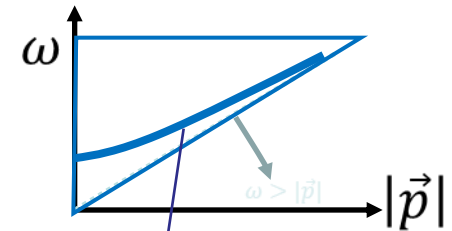
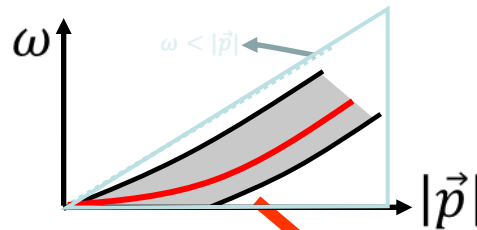
What is the soft mode at CP?

nonzero m_q and charge conjugation symm. is broken by finite chemical potential

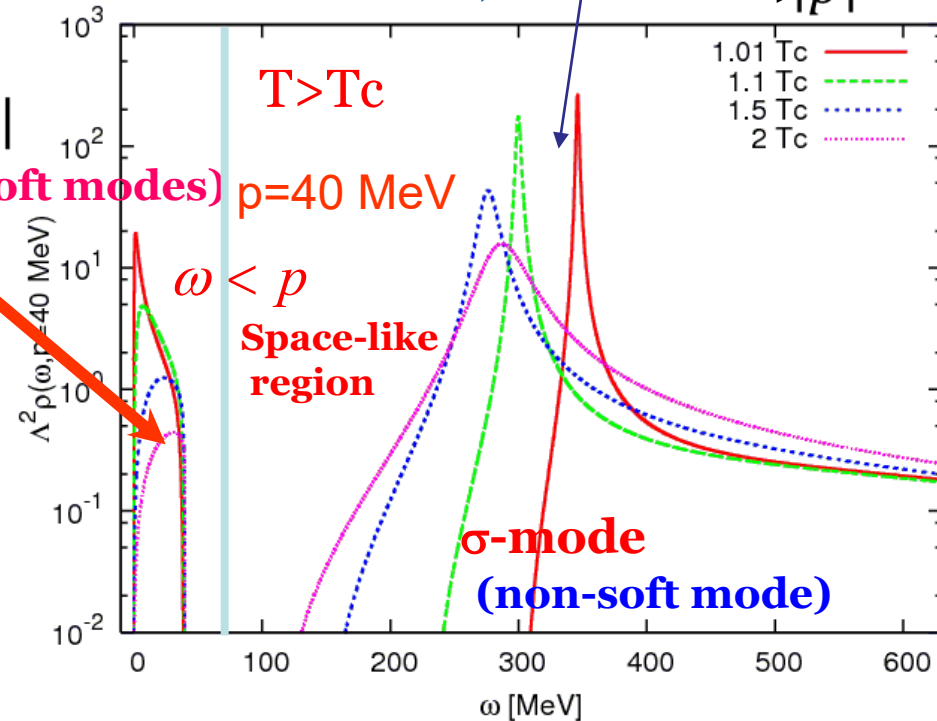
Spectral function in the scalar(-vector) channel

H. Fujii (2003)

H. Fujii and M.Ohtani(2004)



(the soft modes) $p=40$ MeV



Phonon mode in the space-like region softens at CP.

couples to hydrodynamical modes,

leading to interesting dynamical critical phenomena.

See also, D. T. Son and M. Stephanov (2004)

The σ meson has still a non-zero mass at CP, and is not the relevant soft mode.

Possible observables of the precursory soft modes of QCD-CP in HI-collisions: Dilepton production rate (DPR)

T.Nishimura, M. Kitazawa, TK, PTEP 2023 (2023) 5, 053D01;
Annals Phys. 469 (2024) 169768

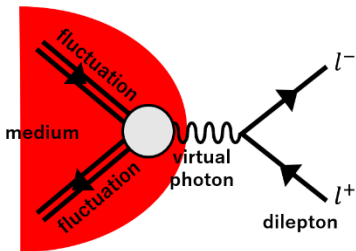
- ✓ Good observable in HIC
- ✓ Computable theoretically

$$\frac{d^4\Gamma}{dk^4}(\mathbf{k}, \omega) = -\frac{\alpha}{12\pi^4} \frac{1}{\omega^2 - \mathbf{k}^2} \frac{1}{e^{\beta\omega} - 1} \rho(\mathbf{k}, \omega)$$

$$\rho(\mathbf{k}, \omega) = g_{\mu\nu} \text{Im}\Pi^{\text{R}\mu\nu}(\mathbf{k}, \omega)$$

Spectral function in the vector channel

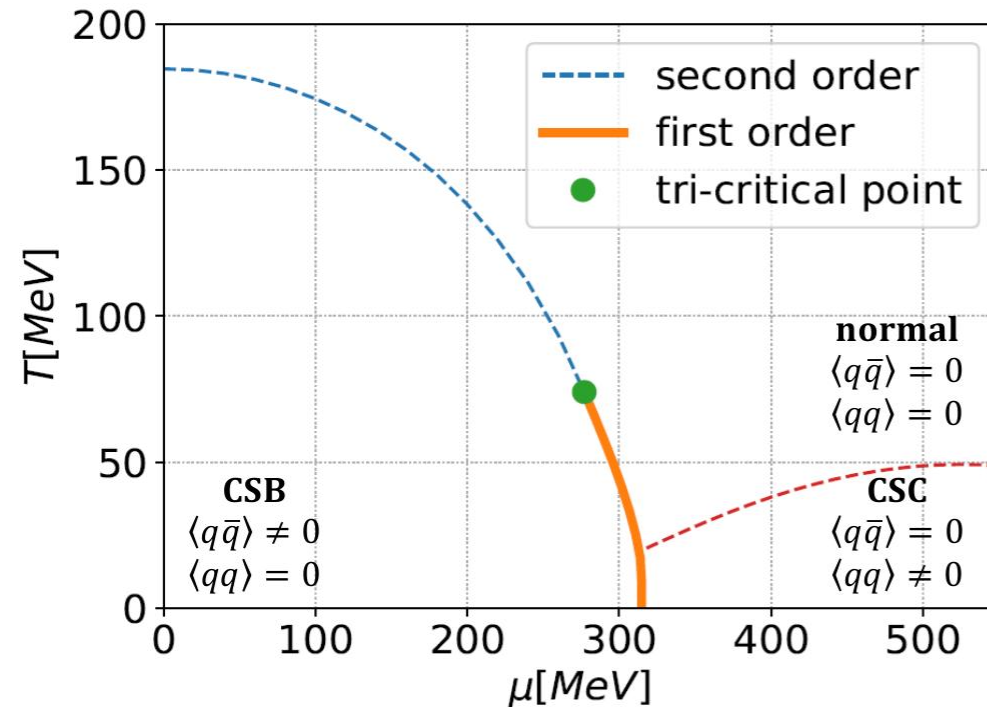
The soft modes associated to the QCD-CP modify the photon self-energy through the mechanism established to account for the para-conductivity above T_c in metal superconductivity.



Lesson from the case of 2SC color superconductivity and condensed matter physics

■ Color superconductivity (CSC)

QCD phase diagram



* Induced by **diquark** condensation

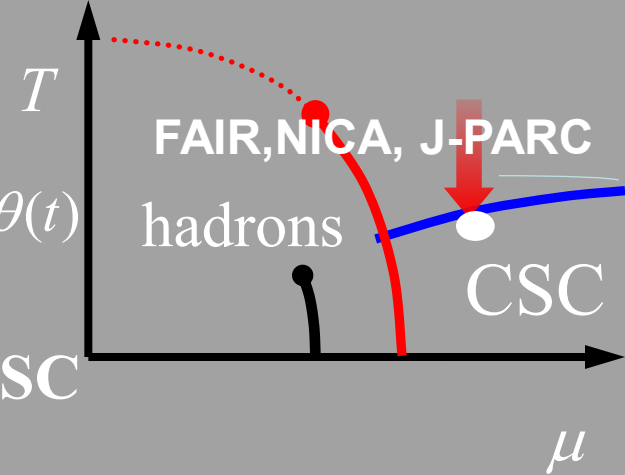
* **Low-temperature** and **high-density**

* It is difficult to observe **CSC** in experiment.

* We focus on **2SC**.

Review : Alford, Schmitt, Rajagopal, Schäfer(2008)

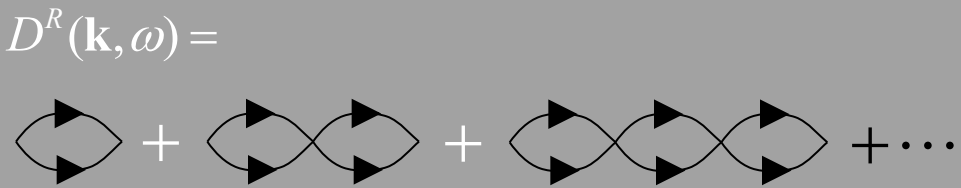
Diquark pair fluctuations prior to CSC



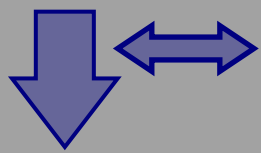
$$D^R(\mathbf{x}, t) = -2G_C \left\langle \left[\bar{\psi}(x) i\gamma_5 \tau_2 \lambda_2 \psi^C(x), \bar{\psi}(0) i\gamma_5 \tau_2 \lambda_2 \psi^C(0) \right] \right\rangle \theta(t)$$

The diquark excitations as the soft mode of CSC

$$\rho(\mathbf{k}, \omega) = -\frac{1}{\pi} \text{Im} D^R(\mathbf{k}, \omega)$$



The strong coupling nature of CSC leads to a sharp peak even at $\epsilon \sim \underline{0.2}$

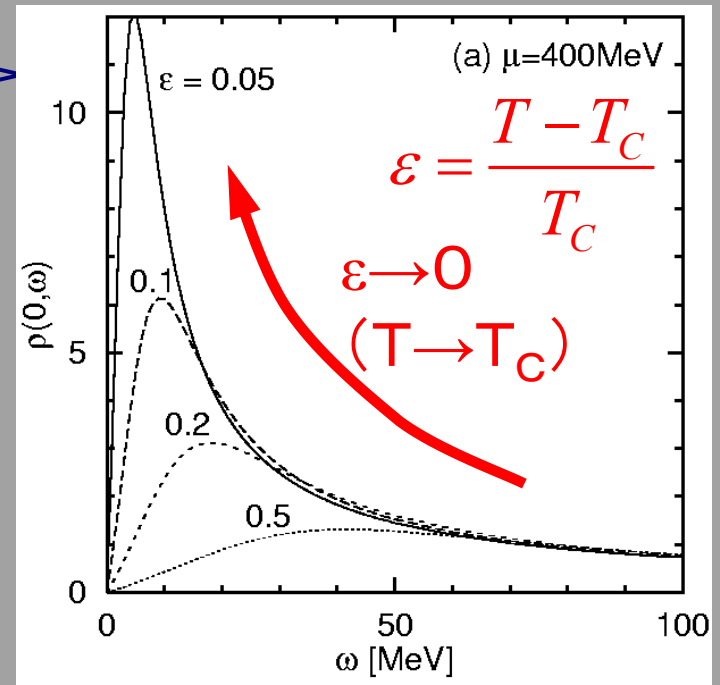


electric SC: $\epsilon \sim \underline{0.005}$
even in 2d-SC

Existence of large pair fluctuations

M.Kitazawa, T. Koide, T. K., Y. Nemoto, PRD 65, 091504 (2002)

It may affect various observables even well above T_c .



We explore the possibility that the diquark fluctuations as the soft modes of CSC anomalously affect dilepton-pair production rate in H-I collisions.

How do they ?

A hint is given by the **para conductivity** (i.e., anomalous enhancement of electric conductivity) through the **Aslamazov-Larkin** as well as the **Maki-Thompson** terms and DOS. A.Aslamasov and A. Larkin, Sov. Phys. SS **10**,875('68)

K. Maki, Prog. Theor. Phys. 39, 897(1968);

R. S. Thompson, Phys. Rev. B1, 327 (1970)

Solid line

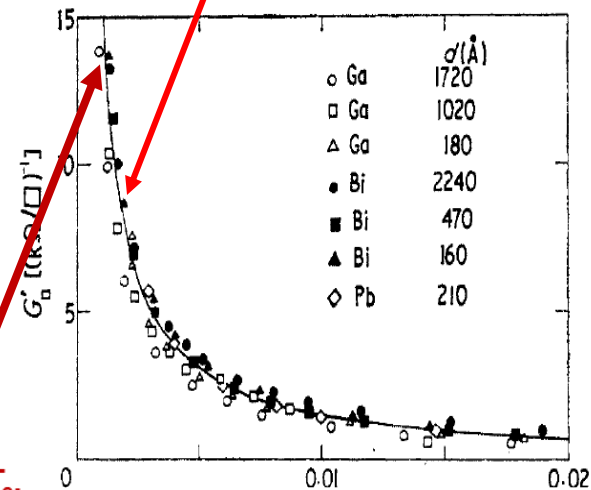
Paraconductivity in metallic SC's near but above T_c

vertical axis : Electric conductivity
horizontal axis : $\epsilon = (T - T_c)/T_c$

$$\Pi_{AL}^{\mu\nu} = \hat{e}\gamma^\mu \text{ [diagram] } \hat{e}\gamma^\nu$$

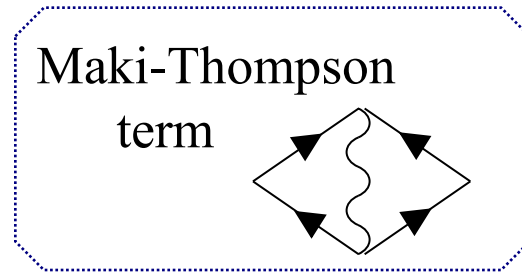
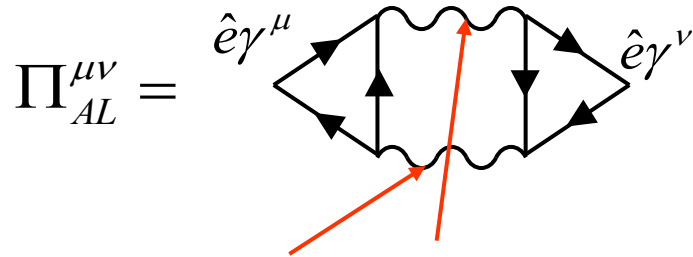
preformed pairing oscillation

The conductivity shows a diverging behavior at $T = T_c$ due to the fluctuations of the electron pair fields above T_c .



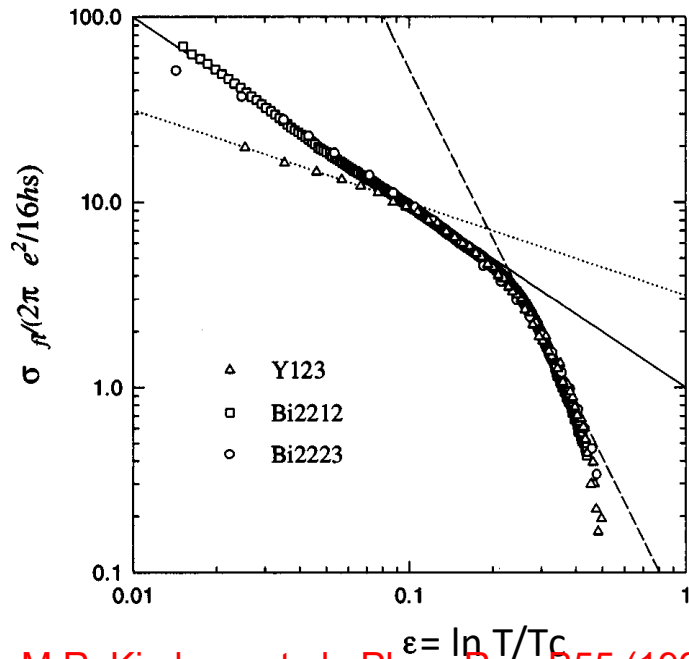
W.J.Skopal and M.Tinkham,
Rep. Prog. Phys.38 (1975),1049

The relevant diagrams



electron pairing oscillations \longrightarrow Diquark oscillations

@ finite q and ω .



the solid line is ϵ^{-1} , the dashed line $0.055/\epsilon^{-3}$, and the dotted line is $3.2/\epsilon^{-1/2}$.

M.R. Kimbere et al, *Phys. Rev. B* 55 (1997):

see also

A. Larkin and A. Varlamov, cond-matt/0109177

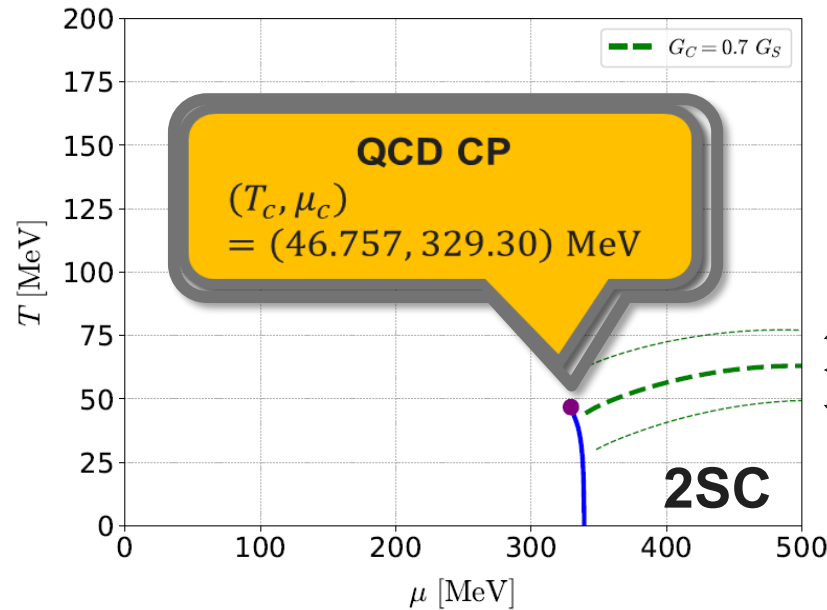
2-flavor NJL model

$$\mathcal{L} = \bar{\psi}i(\gamma^\mu\partial_\mu - m)\psi + \mathcal{L}_S + \mathcal{L}_C$$

$$\left\{ \begin{array}{l} \mathcal{L}_S = G_S [(\bar{\psi}\psi)^2 + (\bar{\psi}i\gamma_5\boldsymbol{\tau}\psi)^2] \\ \mathcal{L}_C = G_C (\bar{\psi}i\gamma_5\boldsymbol{\tau}_2\lambda_A\psi^C)(\bar{\psi}^C i\gamma_5\boldsymbol{\tau}_2\lambda_A\psi) \end{array} \right.$$

($m = 5.5 \text{ MeV}, G_S = 5.50 \text{ MeV}, \Lambda = 631 \text{ MeV}$)

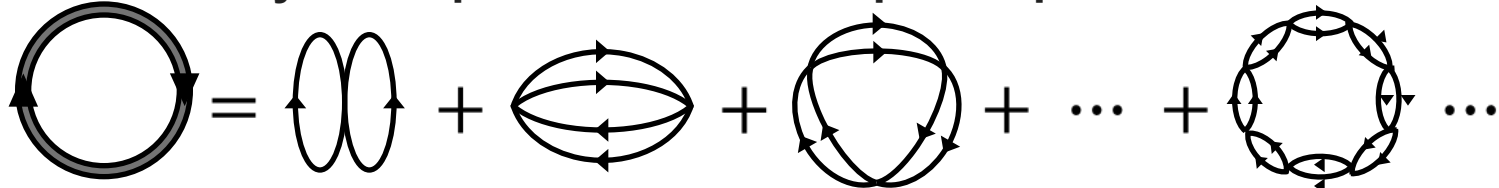
Kitazawa, Koide, Kunihiro, Nemoto (2002)



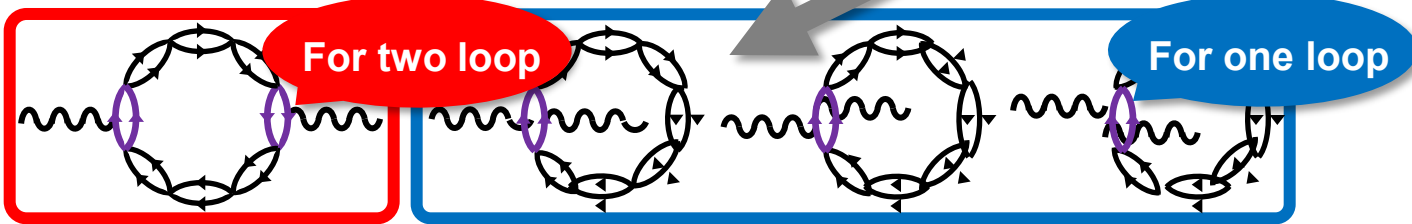
Systematic Computation of the effects of the soft mode on the photon self-energy with the Ward-Takahashi relation being satisfied

T.Nishimura, M. Kitazawa, TK, PTEP 2022 (2022) 9, 093D02

Thermodynamic potential : One loop of diquark fluctuations



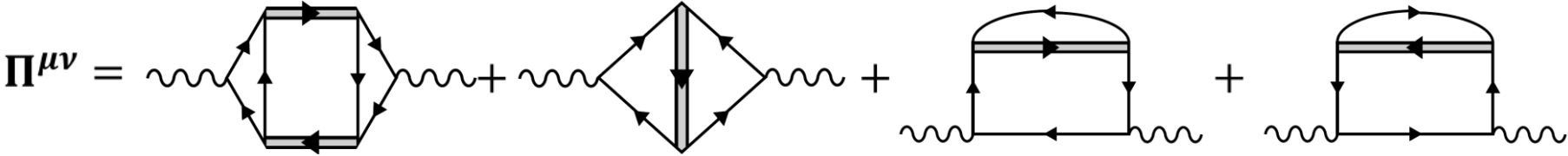
Attach two photons



Aslamazov-Larkin (AL) term

Maki-Thompson (MT) term

Density of states (DOS) term

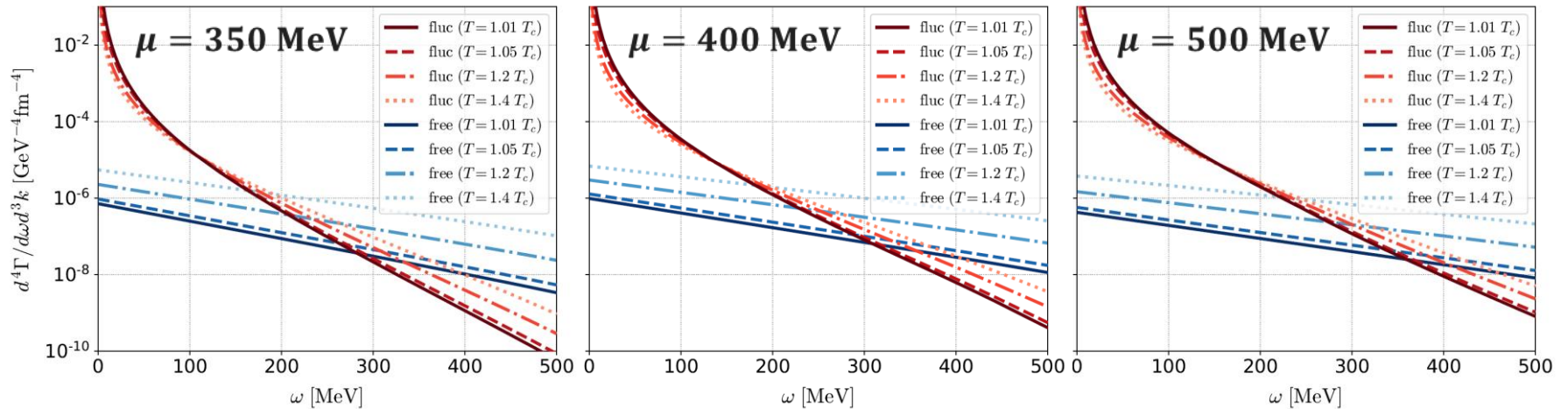


Ward identity $k_\mu \Pi^{\mu\nu} = 0$ is satisfied.

Contribution of diquark soft mode at $k = 0$ ($T > T_c$)

Red lines : Contribution of the soft modes
Blue lines : Contribution of the free quark gases

:



✓ The dilepton production rate is enhanced by the soft mode.

✓ As $T \rightarrow T_c$, the rate becomes bigger.

... This behavior is expected from the property of soft modes.

Dielectron Invariant mass spectra

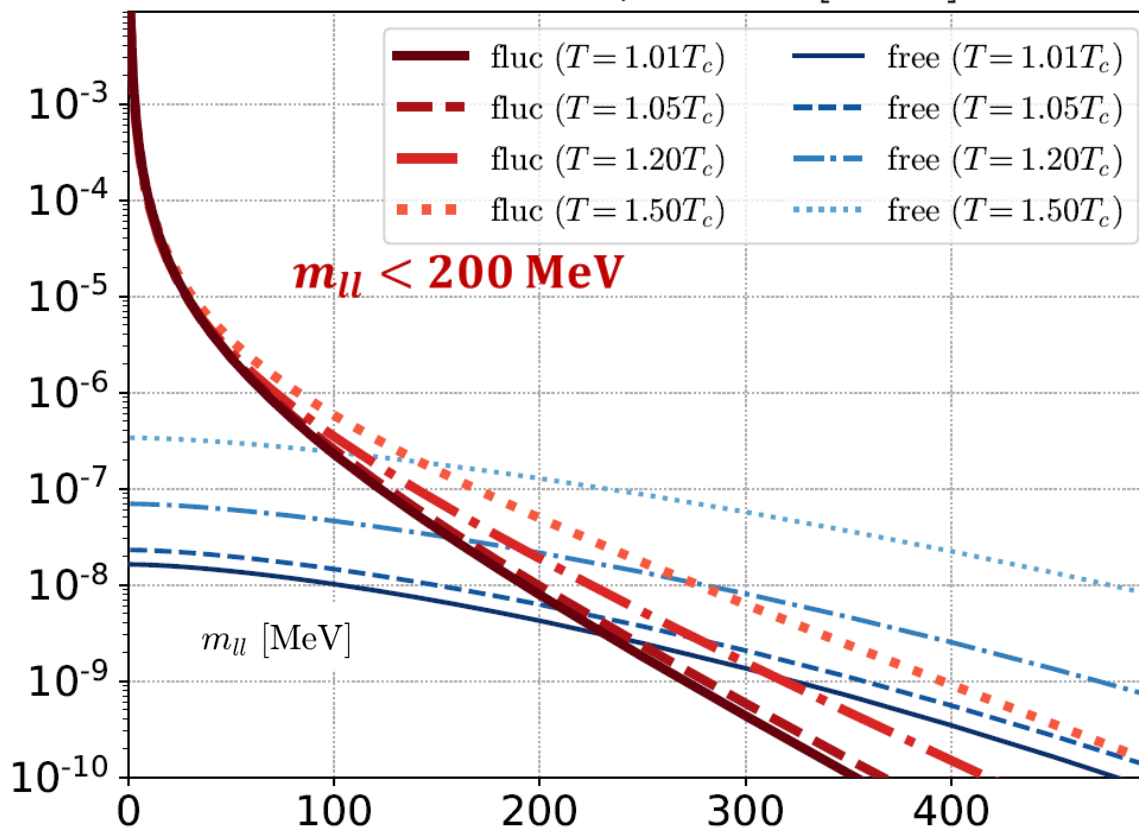
$$\frac{d\Gamma}{dM^2} = \int d^3k \frac{1}{2\omega} \frac{d^4\Gamma}{d^4k} \Big|_{\omega=\sqrt{k^2+M^2}}$$

If the enhancement is confirmed,
it may possibly give an experimental evidence
of the phase transition to CSC !

CSC

$G_C = 0.7G_S, \mu = 350$ [MeV]

$d\Gamma/dm_{ll}^2$ [GeV⁻²fm⁻⁴]



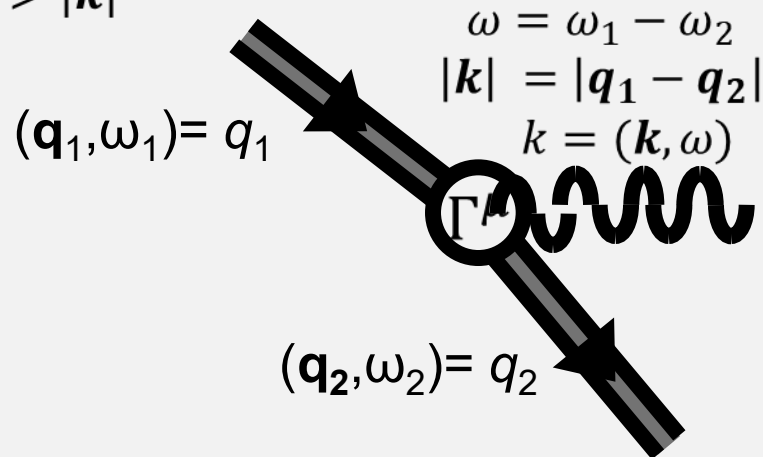
Mechanism of lepton-pair production by two space-like modes

Virtual photon creation in the time-like region by the diquark soft modes

$$|\mathbf{q}_1| > \omega_1$$

$$|\mathbf{q}_2| > \omega_2$$

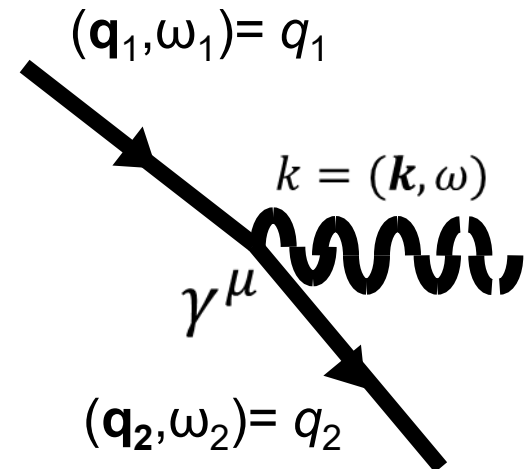
$$\omega > |\mathbf{k}|$$



Although the diquark soft modes have spectral support in the space-like region, a combined excitation of the two modes can have spectral support in the time-like region.

* MT&DOS cancel. → Scattering in AL.

Scattering process with free quarks



The scattering process only produces virtual photon in the space-like region, while the $q\bar{q}$ annihilation contributes dilepton production in time-like.

Transport coefficients: electric conductivity

Photon self-energy $\rho(\mathbf{k}, \omega) = g_{\mu\nu} \text{Im}\Pi^{R\mu\nu}(\mathbf{k}, \omega)$

Dilepton production rate

$$\frac{d^4\Gamma}{dk^4}(\mathbf{k}, \omega) = -\frac{\alpha}{12\pi^4} \frac{1}{\omega^2 - \mathbf{k}^2} \frac{1}{e^{\beta\omega} - 1} \rho(\mathbf{k}, \omega)$$

If the hydrodynamic picture is valid in the low energy-momentum region...

$$\rho(\mathbf{k}, \omega) = g_{\mu\nu} \text{Im}\Pi^{R\mu\nu}(\mathbf{k}, \omega) \simeq \frac{\sigma\omega(\omega^2 - \mathbf{k}^2)}{(\tau\omega^2 - D\mathbf{k}^2)^2 + \omega^2} + 2\frac{\sigma\omega}{\tau^2\omega^2 + 1}$$

Kadanoff, Martin (1963)

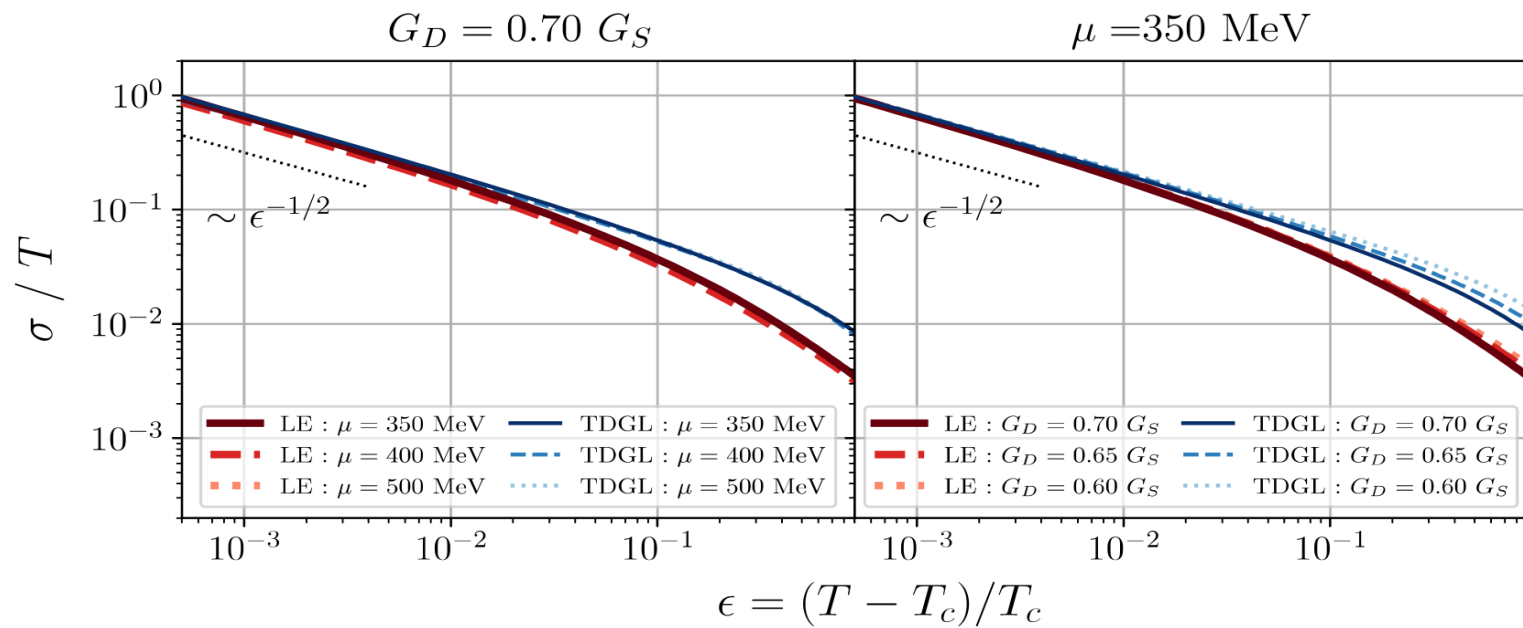
$$\rightarrow \rho(\mathbf{0}, \omega) = -\sum_{i=1,2,3} \text{Im}\Pi^{Rii}(\mathbf{0}, \omega) = 3\frac{\sigma\omega}{\tau^2\omega^2 + 1}$$

Electric conductivity $\sigma = \frac{1}{3} \left. \frac{\partial \rho(\mathbf{0}, \omega)}{\partial \omega} \right|_{\omega=0}$

Relaxation time

$$\tau = \sqrt{-\frac{1}{3!} \frac{\partial^3 \rho(\mathbf{0}, \omega)}{\partial \omega^3} \Big|_{\omega=0} / \frac{\partial \rho(\mathbf{0}, \omega)}{\partial \omega} \Big|_{\omega=0}}$$

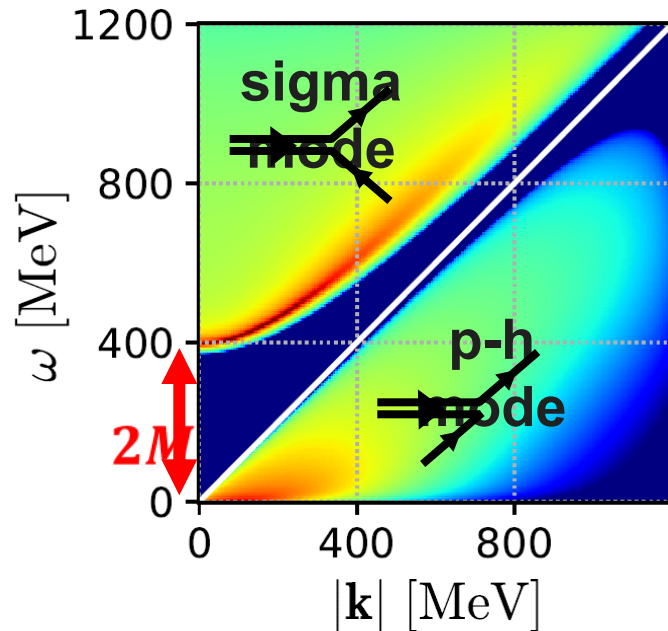
Para-conductivity associated to 2SC



The soft modes at the QCD critical point

Soft modes for QCD^{CP}

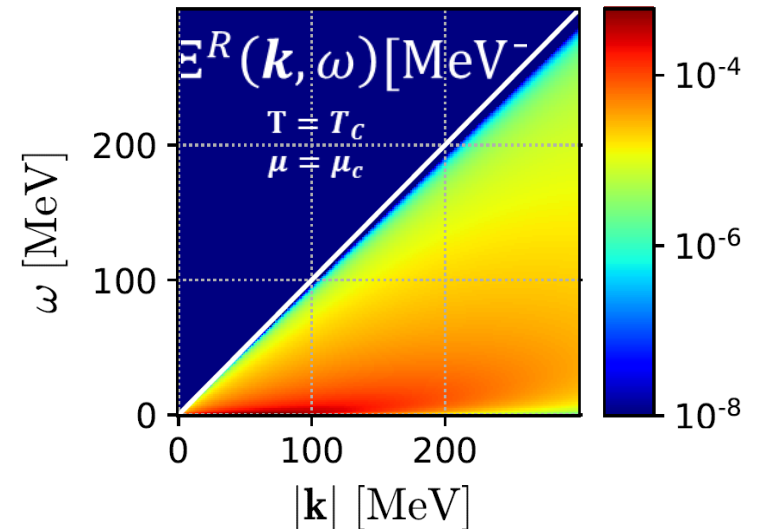
In our work, the soft mode of QCD CP is the particle-hole (p-h) mode.



Particle-hole soft modes

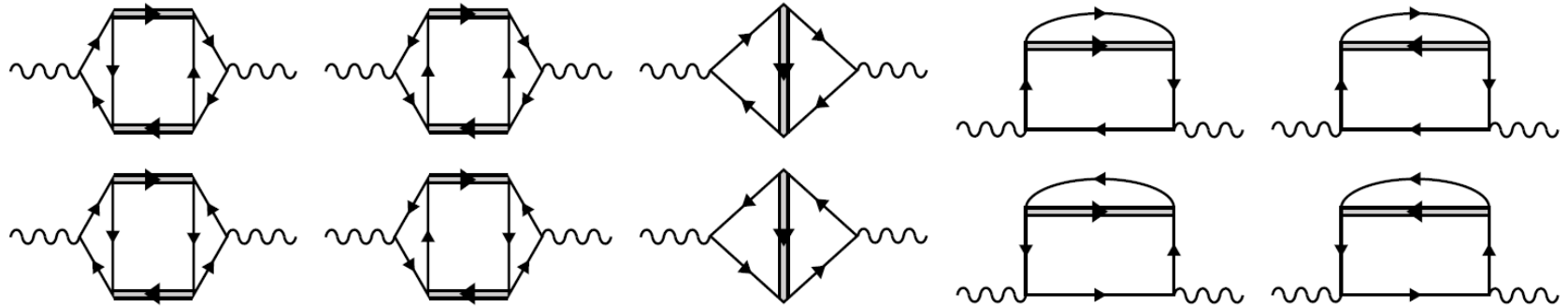
$$E(\mathbf{k}, \omega) \equiv G_S + \text{[T-matrix approx.]}$$

$$\equiv G_S + \text{[loop diagrams]} + \dots$$



Fujii, Ohtani (2004)

Photon self-energy $\Pi^{\mu\nu}(k, \omega)$ in case of p-h soft mode

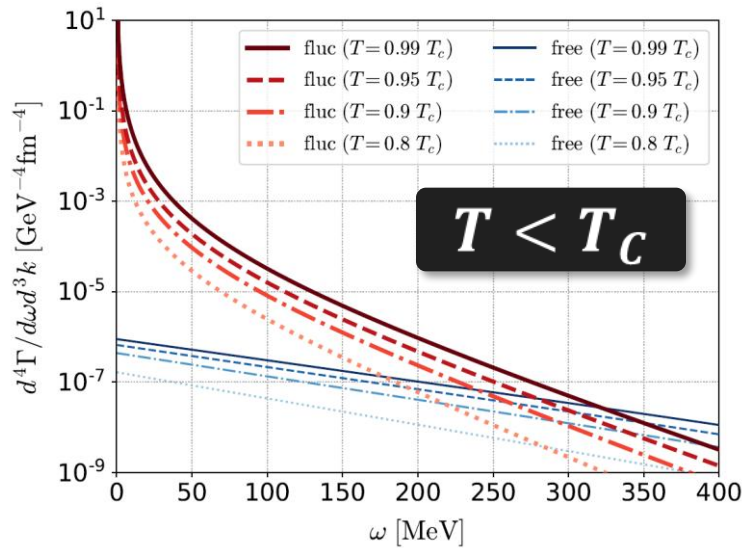


CS \searrow \swarrow QCD CP

$\Pi^{\mu\nu} =$
Aslamazov-Larkin (AL) term
Maki-Thompson (MT) term
Density of states (DOS) term

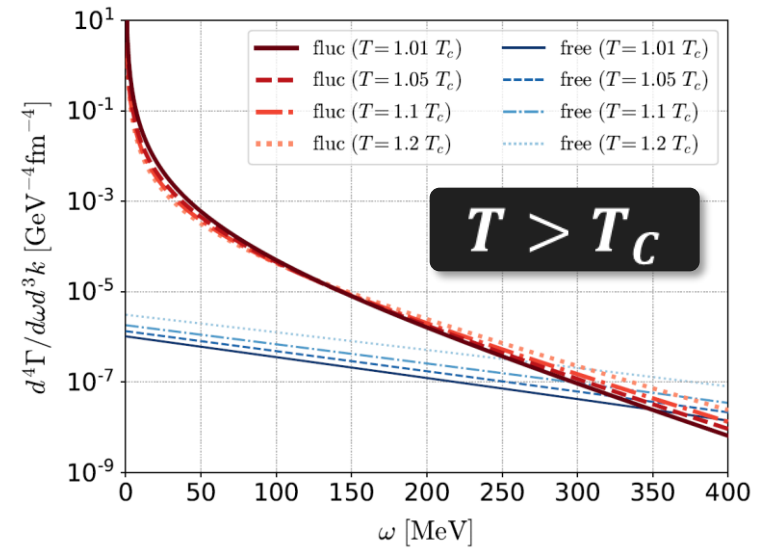
Ward identity of $\Pi^{\mu\nu}$ is satisfied.

Contribution of p-h (QCD CP) soft mode at $k = 0$ ($\mu = \mu_c$)



Bigger as $T \rightarrow T_c$

Bigger as T is bigger



Competition between
contributions of soft modes and
kinematical (temperature) effects

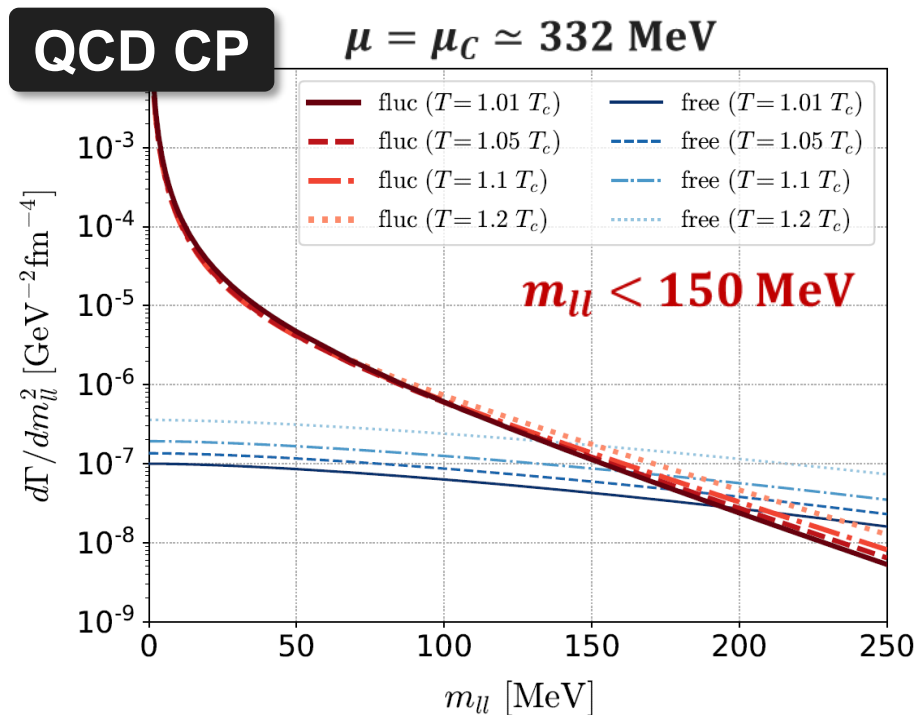
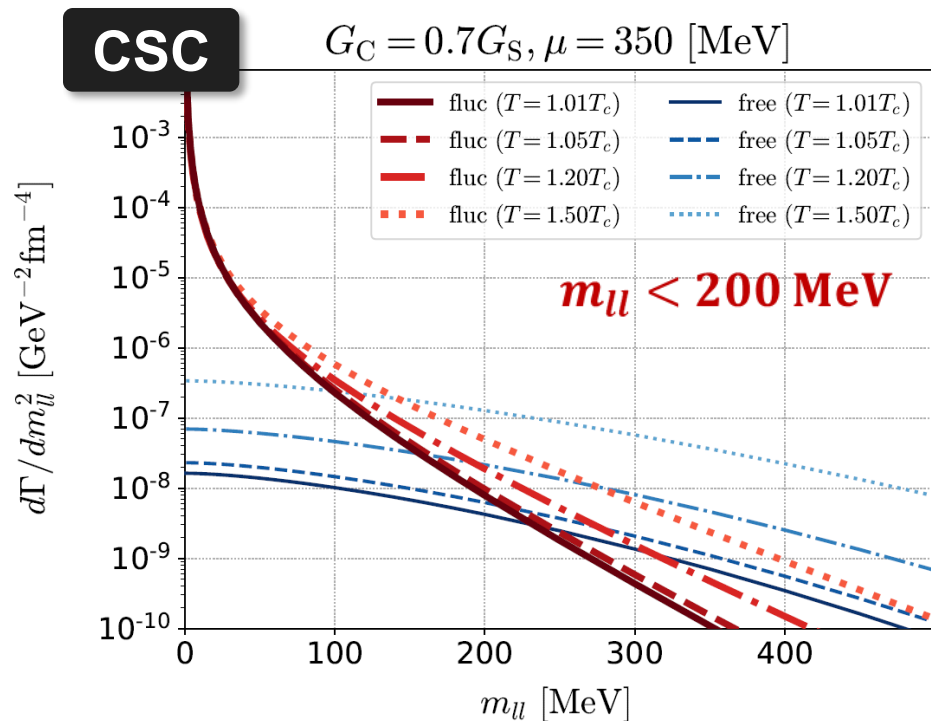
$$\frac{d^4\Gamma}{dk^4}(\mathbf{k}, \omega) = -\frac{\alpha}{12\pi^4} \frac{1}{\omega^2 - \mathbf{k}^2} \frac{1}{e\beta\omega - 1} \rho(\mathbf{k}, \omega),$$

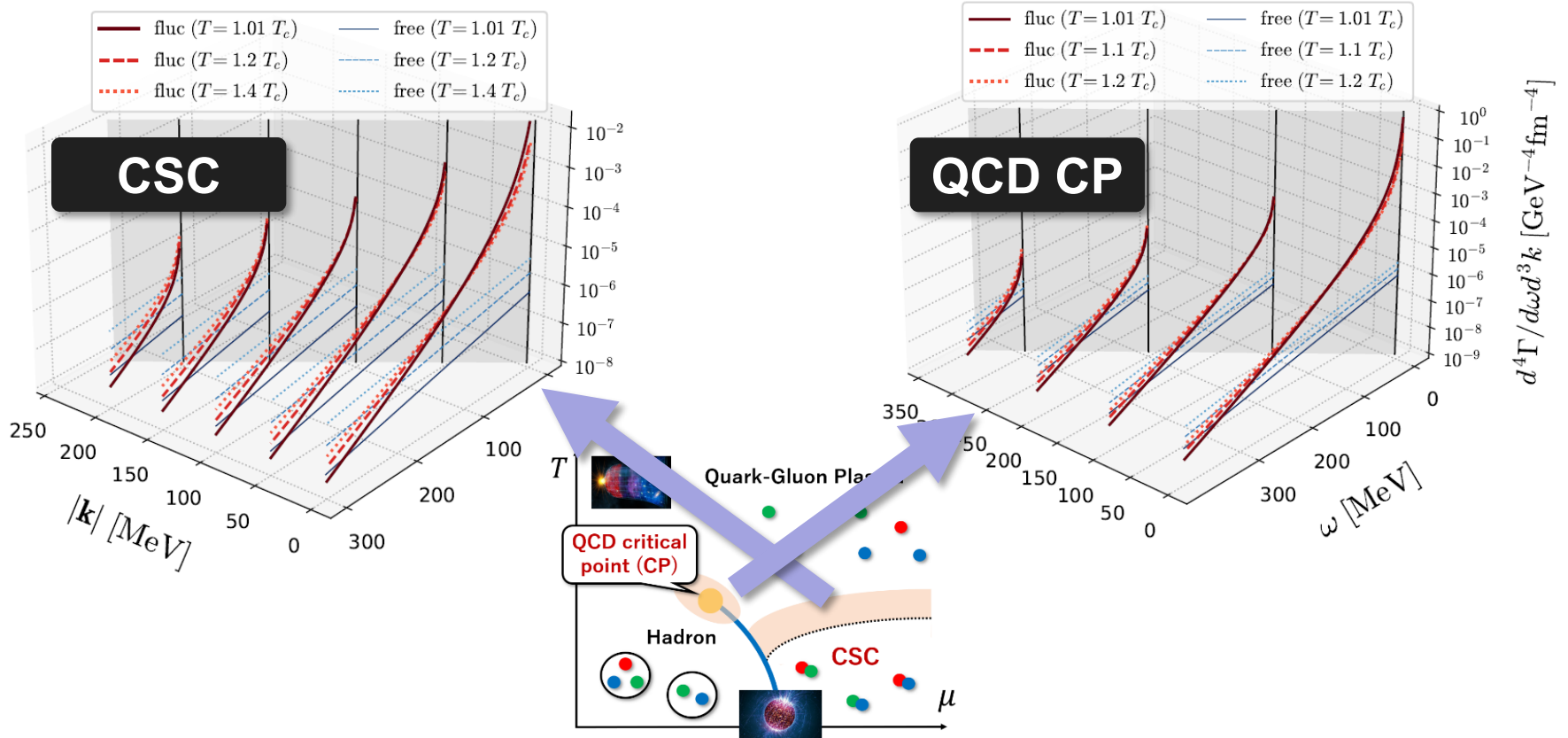
$$\rho(\mathbf{k}, \omega) = g_{\mu\nu} \text{Im} \Pi_{AL}^{R\mu\nu}(\mathbf{k}, \omega)$$

Invariant mass spectra

$$\frac{d\Gamma}{dM^2} = \int d^3k \frac{1}{2\omega} \frac{d^4\Gamma}{d^4k} \Big|_{\omega=\sqrt{k^2+M^2}}$$

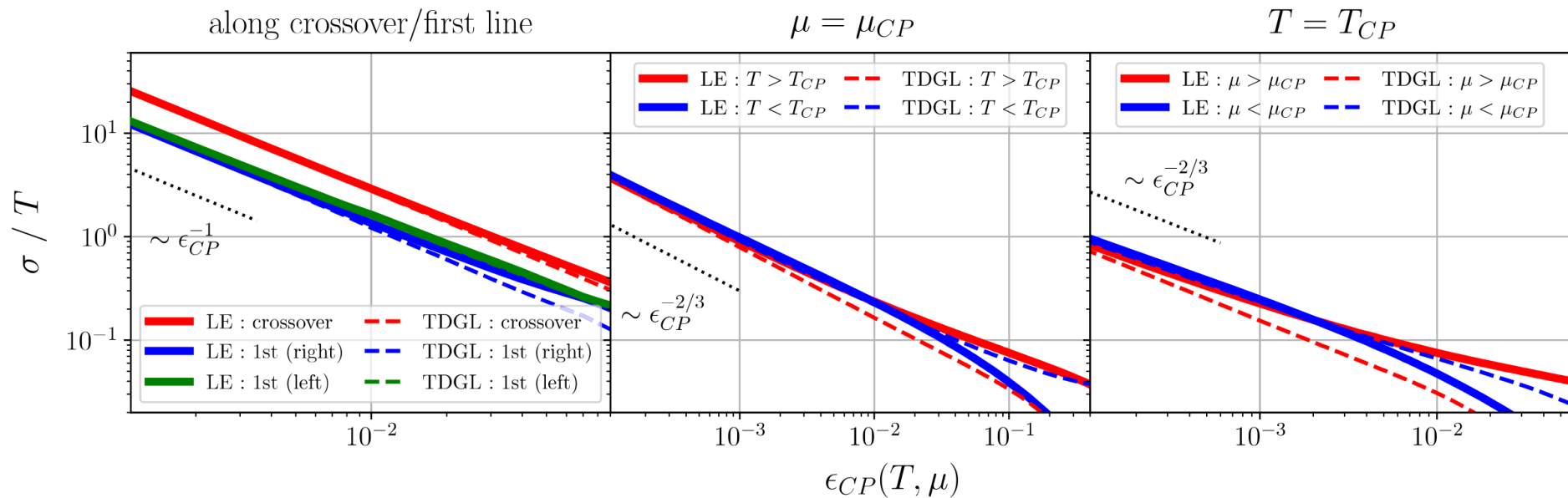
If the enhancement is confirmed,
it may possibly give an experimental evidence
of the phase transition to CSC & QCD CP!



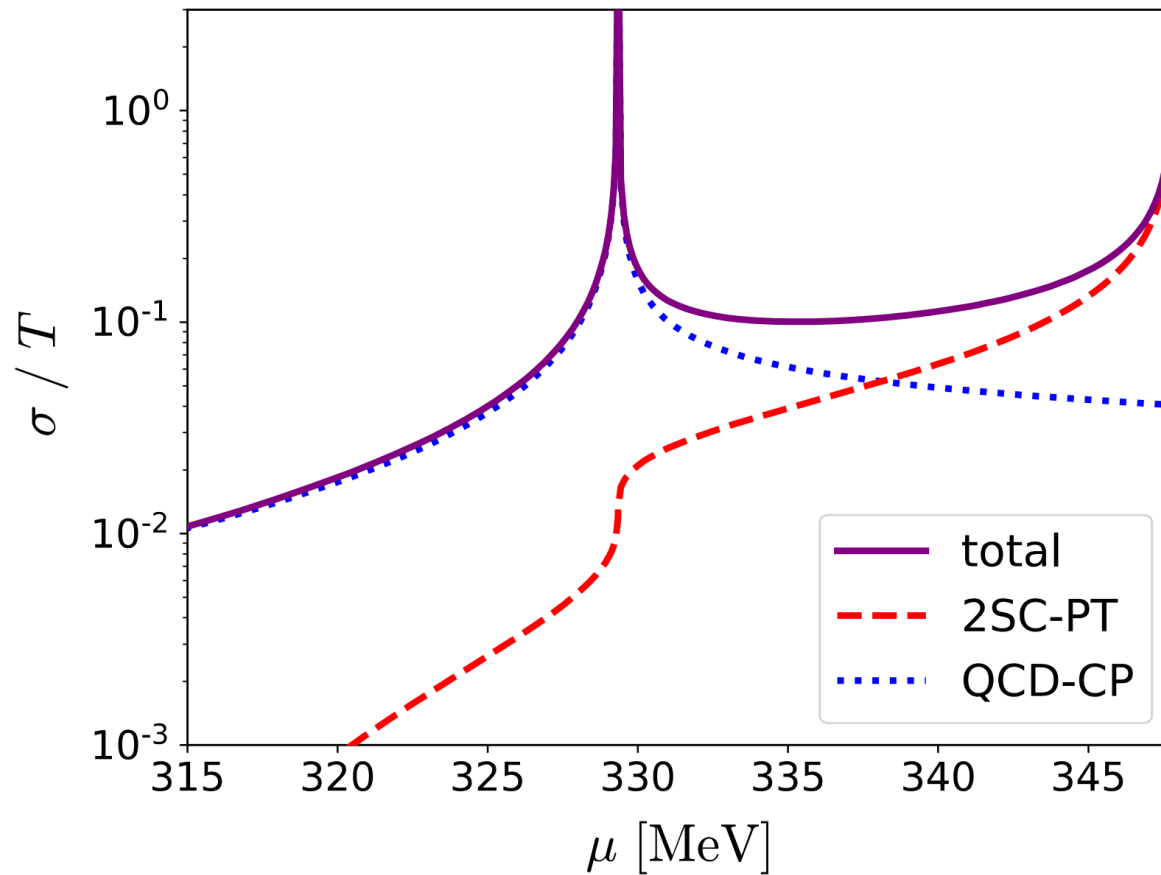


What is the fundamental difference between the two cases?
 → See the critical behavior of transport coefficients!

Electric conductivity around QCD-CP



Chemical potential dependence of conductivity around QCD-CP



Brief summary

1. We have shown how the dilepton production rate is affected by the soft modes of the QCD CP as well as 2SC, which is very specific to QCD,
2. and the soft modes also cause an enhancement of electric conductivity around the critical region, as the para-conductivity in metal superconductors.
3. The anomalous enhancement of the di-electron production in the very low energy region. We admit that it would be an experimental challenge to detect them, but it should be a quite useful to identify the existence of these phase transitions in the QCD phase diagram.

Remark:

Such an experiment may provide a part of a **multi-messenger HI science**, analogously to **multi-messenger astronomy** consisting of gravitational Waves, photons, neutrinos, and other cosmic rays. >> Dynamical model cal.

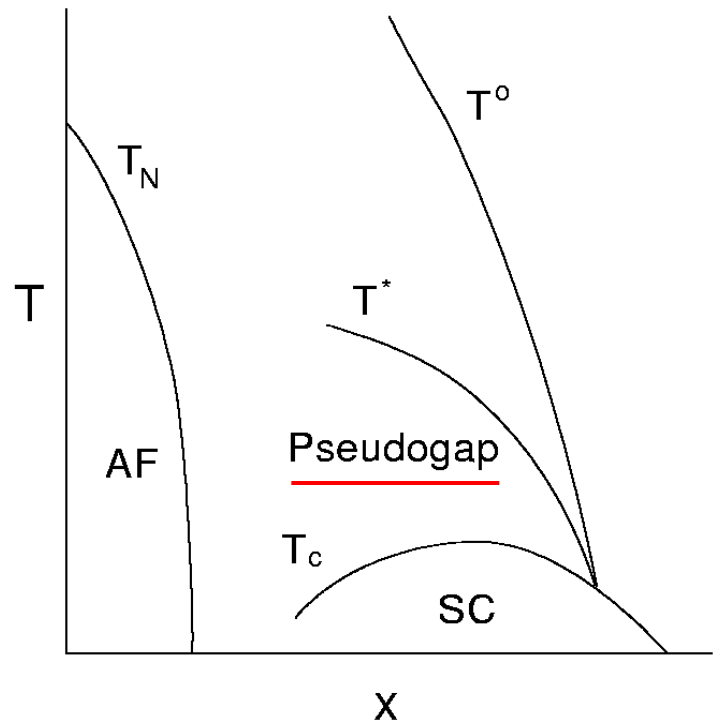
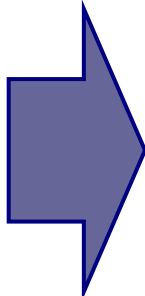
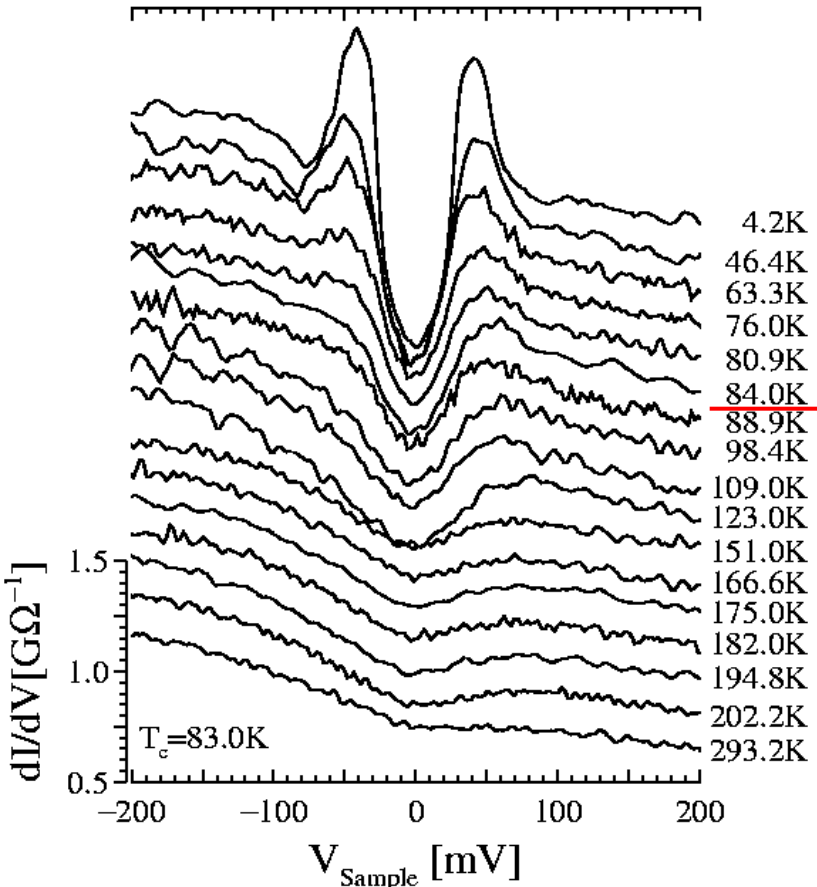
Further development: an enhancement of relaxation times,
T. Nishimura, M. Kitazawa and TK,
Annals Phys. 469 (2024) 169768

BACK UPS

● Pseudogap in High T_c Superconductors

: Anomalous depression of the density of state near the Fermi surface in the normal phase.

Conceptual phase diagram of HTSC cuprates



Renner et al.('96)

The origin of the pseudogap in HTSC is **still controversial.**

The pseudo gap in the quark spectra in the quark matter prior to CSC

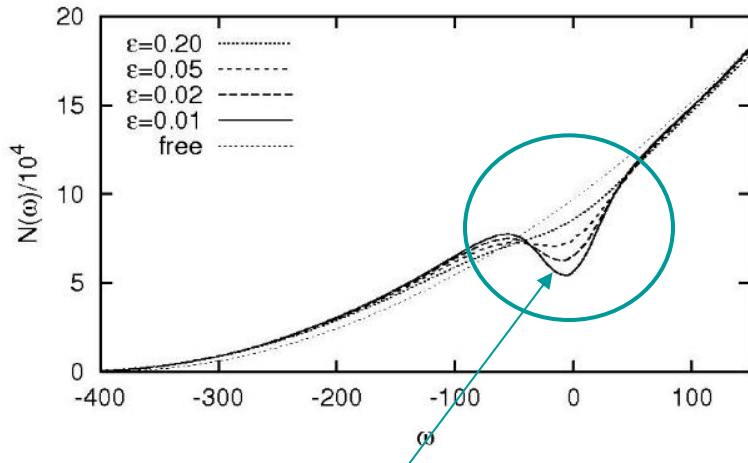
M. Kitazawa, T. Koide, Y. Nemoto and T.K., Phys. Rev. D70 (2004)

$$G(\mathbf{k}, \omega_n) = \frac{1}{G^0(\mathbf{k}, i\omega_n) - \Sigma(\mathbf{k}, i\omega_n)} \quad G^0(\mathbf{k}, i\omega_n) = \left[(i\omega_n + \mu)\gamma^0 - \mathbf{k} \cdot \vec{\gamma} \right]^{-1} = \rightarrow$$

$$\Sigma(\mathbf{k}, \omega_n) = \text{[diagram: shaded circle with diagonal lines]} = \text{[diagram: loop with arrow]} + \text{[diagram: loop with two arrows]} + \text{[diagram: loop with four arrows]} + \dots$$

$$\equiv \text{[diagram: wavy line with arrow labeled } \mathbf{q}, i\omega_m \text{ and } \mathbf{k} + \mathbf{q}, i\omega_n + i\omega_m \text{]} = T \sum_m \int \frac{d^3 \mathbf{q}}{(2\pi)^3} \Xi(\mathbf{k} + \mathbf{q}, \omega_n + \omega_m) G^0(\mathbf{q}, \omega_m)$$

Soft mode



A considerable depression in the quark level density; an incomplete gap, i.e., the pseudogap!

Quite interesting!, but may be difficult to experimentally confirm it, unfortunately.

Pseudo gap !

Is there any chance to probe the enhanced diquark soft modes of CSC in an experiment?

Functional RG approach with meson-quark model

Takeru Yokota, Kenji Morita and TK, PTEP (2016) 073D01

$$S_\Lambda [\bar{\psi}, \psi, \phi = (\sigma, \vec{\pi})] = \int_0^{\frac{1}{T}} d\tau \int d^3\vec{x} \left\{ \bar{\psi} (\not{\partial} + g_s(\sigma + i\vec{\tau} \cdot \vec{\pi}\gamma_5) - \mu\gamma_0) \psi + \frac{1}{2}(\partial_\mu\phi)^2 + V_\Lambda(\phi^2) - c\sigma \right\}$$

$$V_\Lambda(\phi^2) = \frac{1}{2}m_\Lambda^2\phi^2 + \frac{1}{4!}\lambda(\phi^2)^2$$

Λ	m_Λ/Λ	λ_Λ	c/Λ^3	g_s
1000M eV	0.79 4	2.00	0.0017 5	3.2



Vacuum
value

M_q	M_π	M_σ	σ_0
286Me V	137MeV	496MeV	93Me V

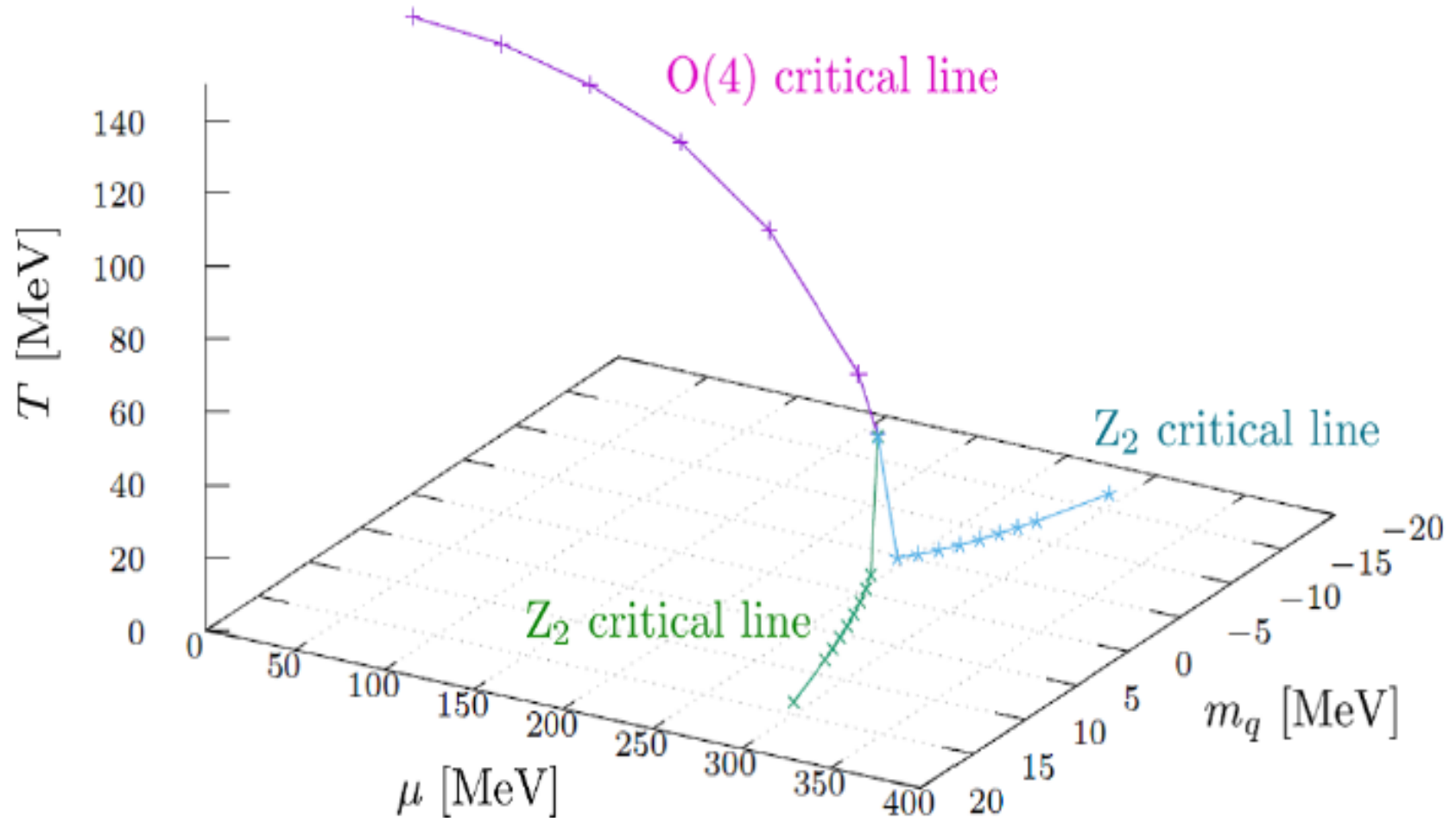
R. Tripolt, L. Smekal, J. Wambach,
PRD90 (2014)

M_q is the constituent
quark mass.
 $M_q = g_s\sigma_0$

M_π and M_σ are the screening
masses.
 $M_\pi = \frac{1}{\sigma_0} \frac{\partial U_0}{\partial \sigma} (\sigma_0) \frac{M_\sigma}{\partial^2 U_0} (\sigma_0)$
 $= \frac{M_\sigma}{\partial \sigma^2} (\sigma_0)$

Quark-mass dependence of the phase transition by FRG analysis with meson-quark model

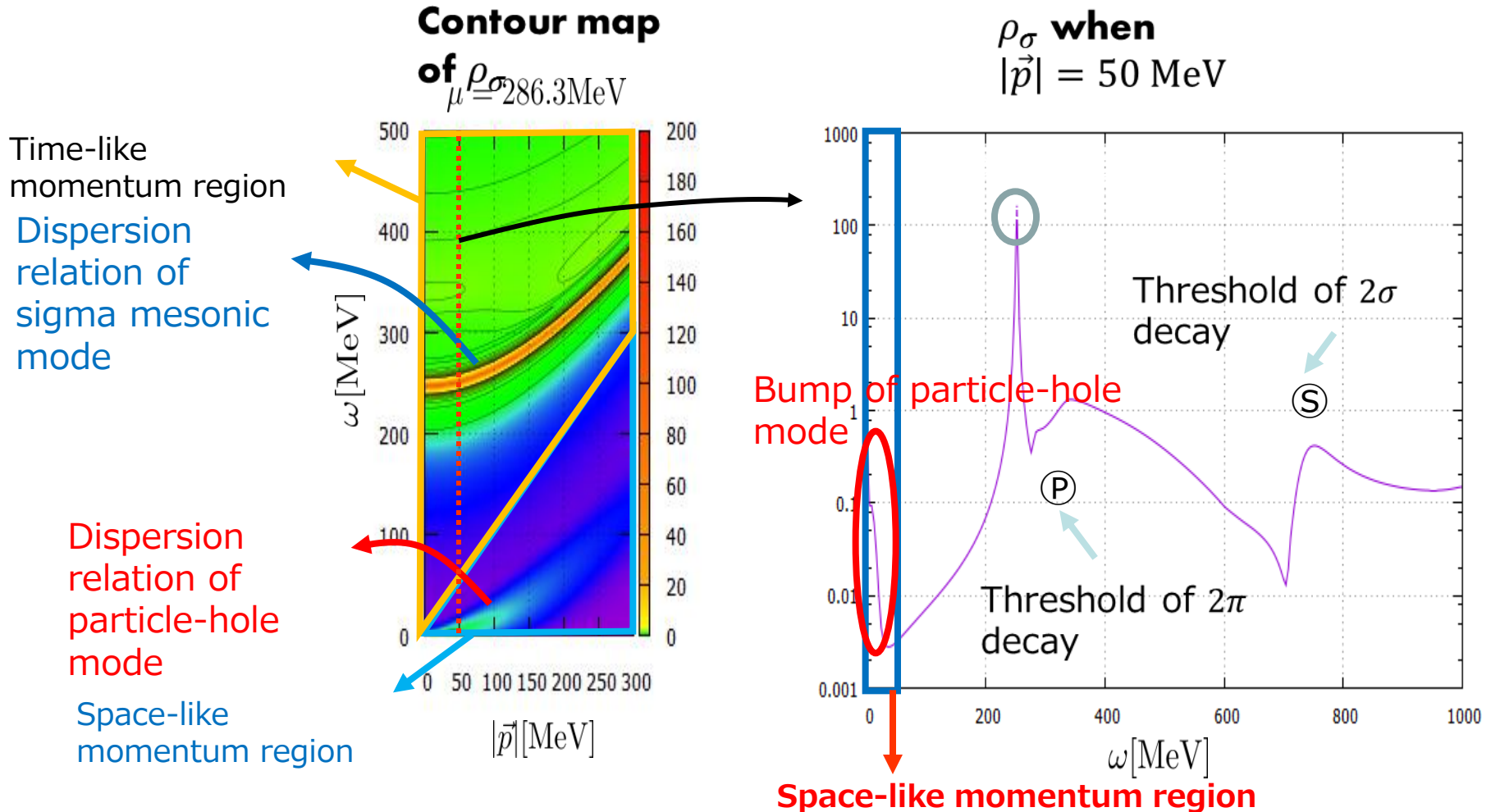
Takeru Yokota, Kenji Morita and TK, PTEP (2016) 073D01



Spectral function of scalar-density collective mode around QCDCP

Takeru Yokota, Kenji Morita and TK, PTEP (2016) 073D01

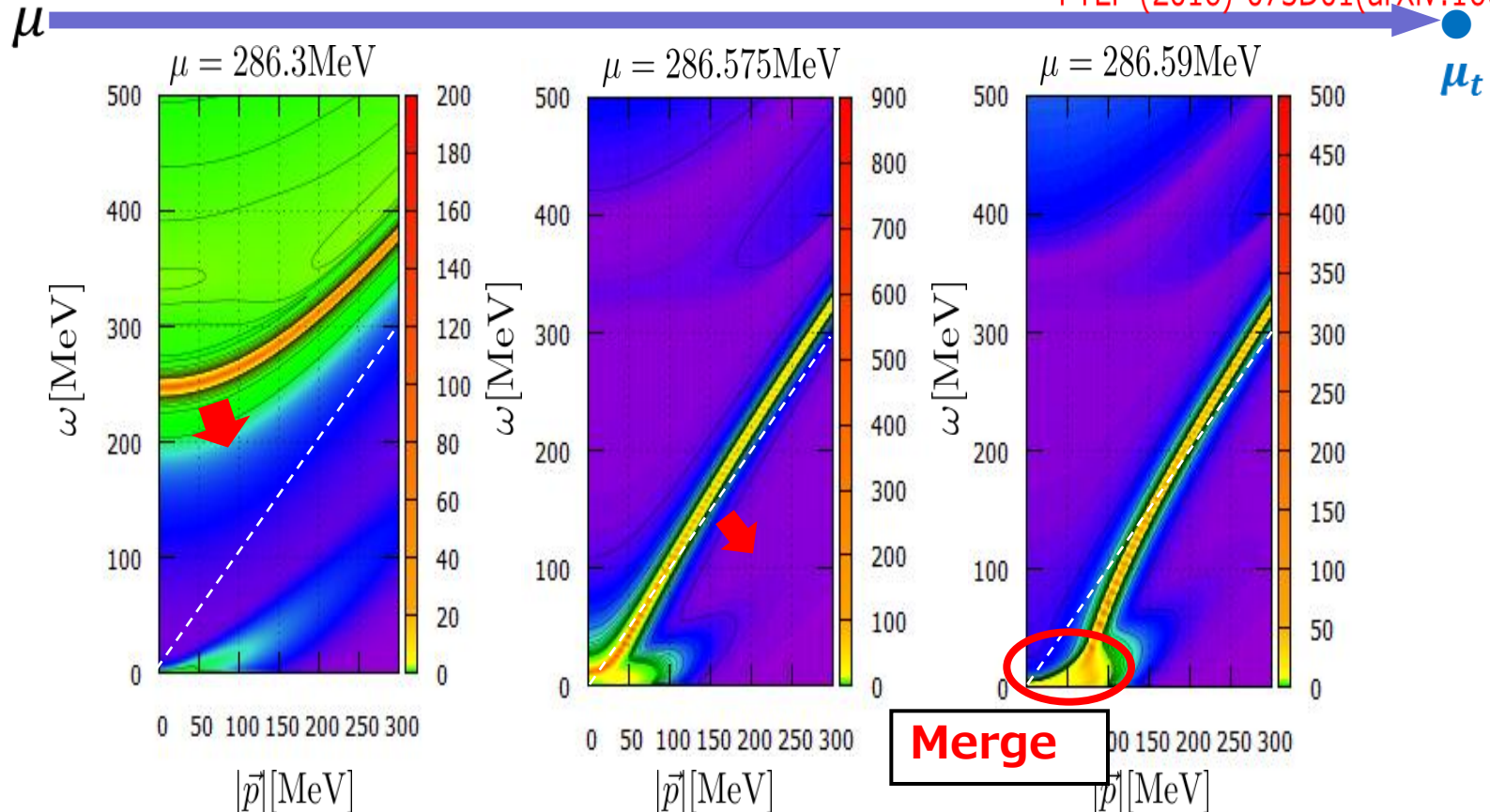
We fix $T = 5.1$ MeV below. (At this temperature, the transition chemical potential is $\mu_t = 286.69$ MeV)



Anomalous softening of the sigma once located in the time-like region to merge into the phonon mode

We fix $T = 5.1$ MeV. (At this temperature, the transition chemical potential is $\mu_t = 286.69$ MeV)

Takeru Yokota, Kenji Morita and TK,
PTEP (2016) 073D01(arXiv:1603.02147)



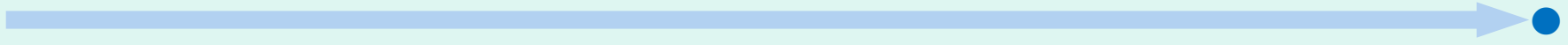
As the system approaches the QCD CP, the dispersion relation of sigma mesonic mode shifts

to low-energy region and merges with the bump of the particle-hole modes in the space-like momentum region.

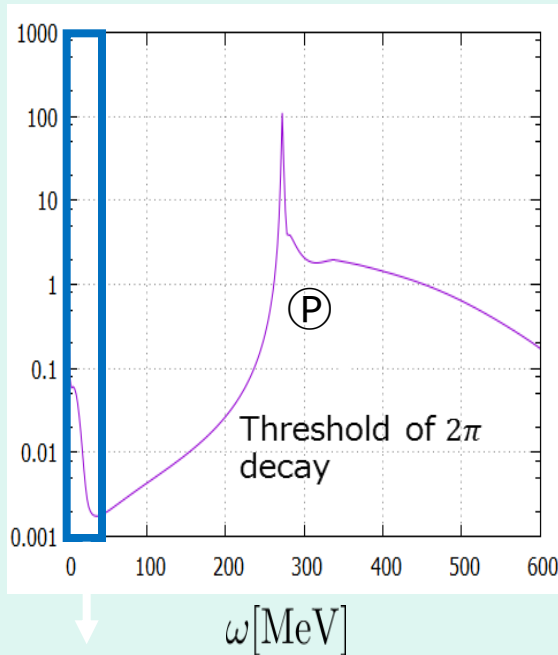
Development of the phonon mode near the QCD CP

Takeru Yokota, Kenji Morita and TK,
PTEP (2016) 073D01(arXiv:1603.02147)

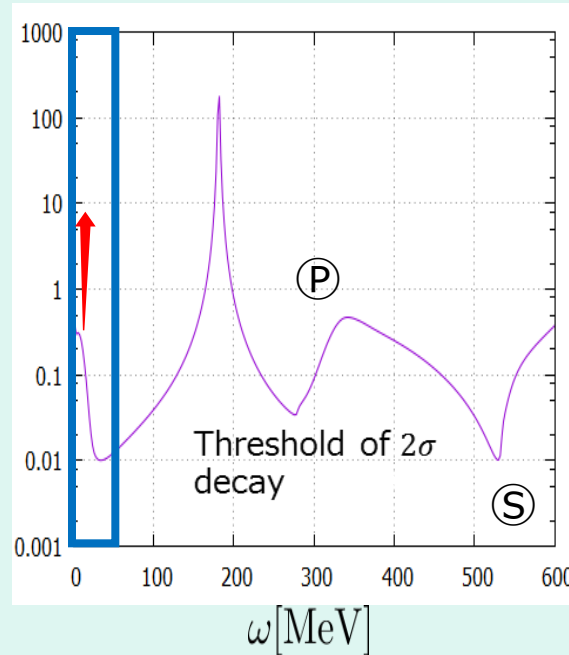
- We fix $T = 5.1$ MeV and $|\vec{p}| = 50$ MeV. (At this temperature, the transition chemical potential is $\mu_t = 286.69$ MeV)

μ  μ_t

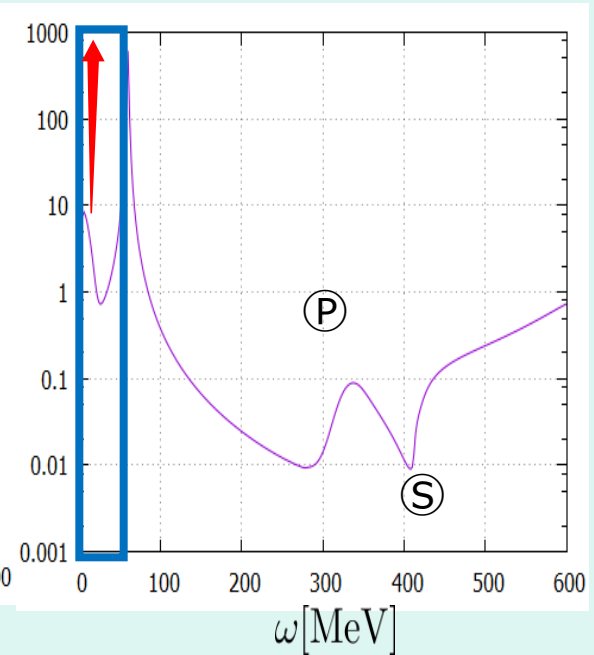
$\mu = 286.0$ MeV



$\mu = 286.5$ MeV



$\mu = 286.57$ MeV



Space-like momentum
region

- The particle-hole mode becomes soft.

A remark to be omitted.

The merge of the sigma and phonon and appearance of the **tachyonic mode** @ $m_q=8.88$ MeV, which implies a **larger scalar-density coupling**.

Smaller current quark mass with smaller scalar-density coupling shows no such anomalous behavior.

Appearance of an acausal tachyonic mode at finite momenta!



Suggesting an instability of the system, to an inhomogeneous phase, with modulated sigma field as well as the density, no pion field, i.e., of the real kink type.

**For the detailed analysis, see
T. Yokota, K. Morita and T.K., Phys. Rev. D96, 074028 (2017).**

c.f. Real Kink type: D.Nickel, M. Buballa (2010).
DCDW: Nakano, Tatsumi (2005); chiral limit.


★ Importance of Vector-type Interaction in CSC

- Vector interaction naturally appears in the effective theories.

- Instanton-anti-instanton molecule model Shaefer, Shuryak ('98)

$$L = G \left\{ \frac{2}{N_C^2} [(\bar{\psi} \tau^a \psi)^2 + (\bar{\psi} \tau^a i \gamma_5 \psi)^2] - \frac{1}{2N_C^2} [(\bar{\psi} \tau^a \gamma^\mu \psi)^2 + (\bar{\psi} \tau^a \gamma^\mu \gamma_5 \psi)^2] \right\} + L_8$$
- Renormalization-group analysis N. Evans et al. ('99)

$$L_{LL}^0 = G_{II} \left\{ (\bar{\psi}_L \gamma^0 \psi_L)^2 - (\bar{\psi}_L \gamma^\mu \psi_L)^2 \right\}$$


 $G_V / G_S = 1/4$

$$-G_V (\bar{\psi} \gamma^\mu \psi)^2 \iff \text{density-density correlation}$$

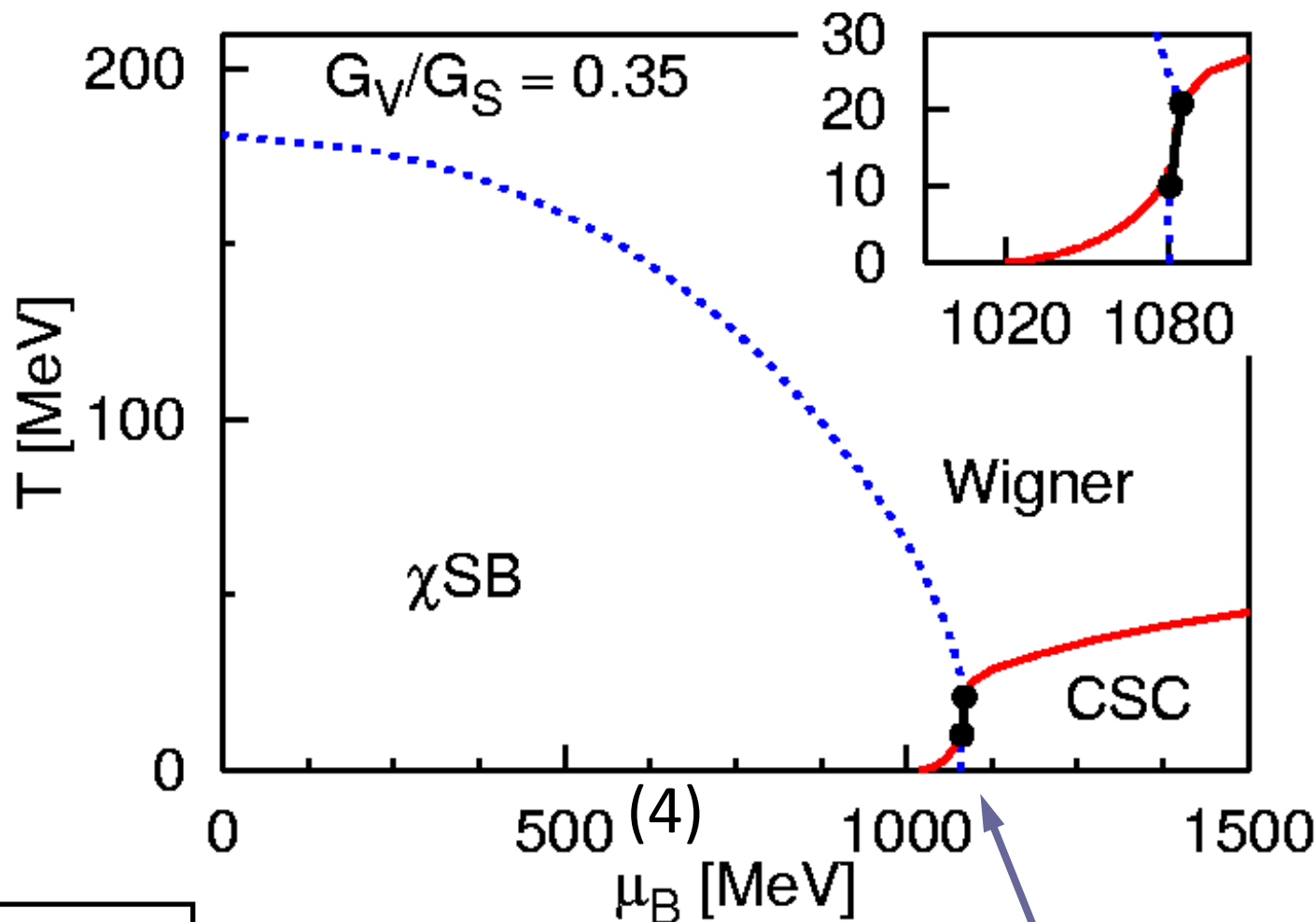
- $-G_V (\bar{\psi} \gamma^0 \psi)^2 \rightarrow -G_V \langle \bar{\psi} \gamma^0 \psi \rangle^2 = -G_V \rho^2 \quad \rho = \langle \bar{\psi} \gamma^0 \psi \rangle$

Chiral restoration is punished by the vector interaction!

With color superconductivity transition incorporated:

Two critical end point!

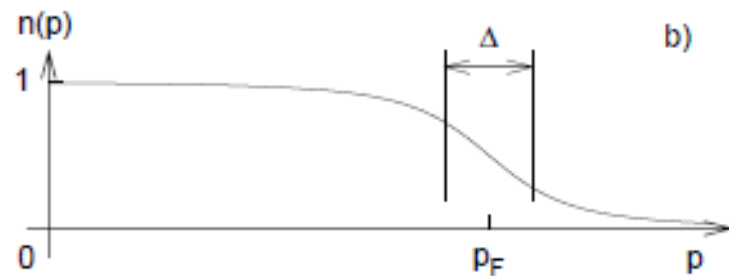
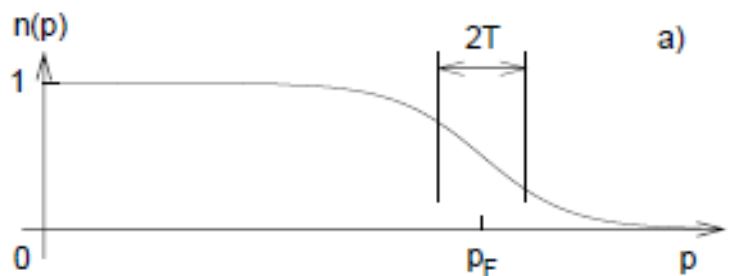
M. Kitazawa, T. Koide, Y. Nemoto and T.K., PTP ('02)



$$G_V / G_S = 0.35$$

Another end point appears from lower temperature, and hence **there can exist two end points** in some range of G_V !

Similarity of the effect of temperature and pairing gap on the chiral condensate.



M. Kitazawa et al.
PTP, 110 (2003), 185:
arXiv:hep-ph/0307278

T



Δ

An enhancement of dilepton-pair production

$$\frac{dR_{ee}}{d^4q} = -\frac{\alpha}{12\pi^4 Q^2} \text{Im} \Pi^{R\mu}_{\mu} \frac{1}{e^{q^0/T} - 1}$$

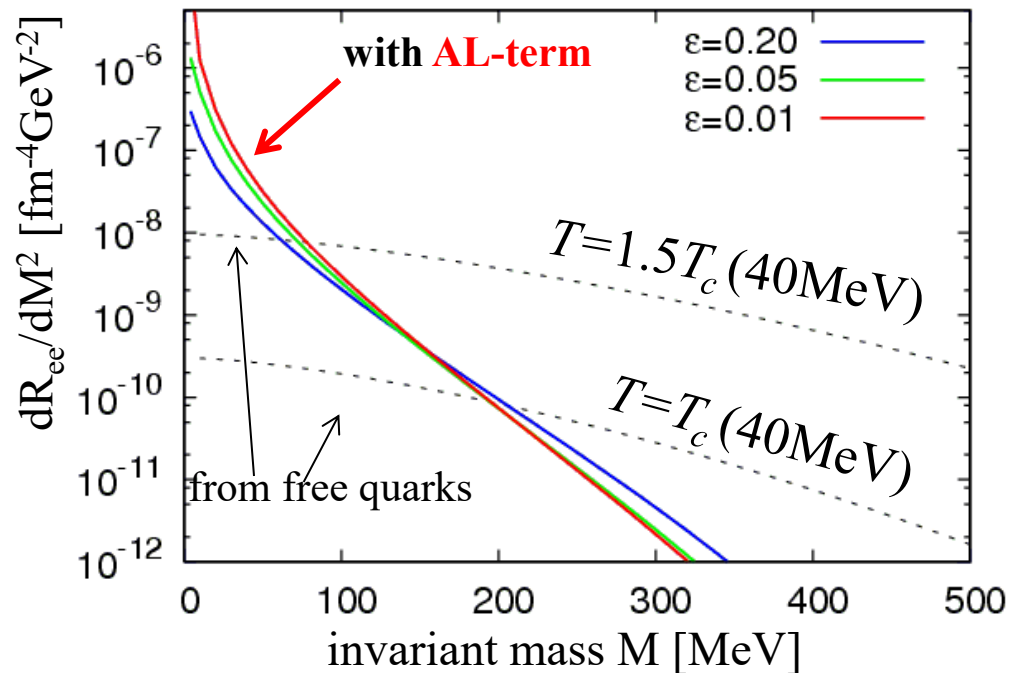
-per invariant mass

$$\frac{dR_{ee}}{dM^2} = \int \frac{d^3q}{2q^0} \frac{dR_{ee}}{d^4q}$$

Prominent enhancement at $M < 150 \text{ MeV}$.

The peak becomes sharp as $\varepsilon \rightarrow 0$.

M. Kitazawa, Y. Nemoto and T.K.,
PoS CPOD07 (2007) 041; arXiv.0711.4429



c.f. other channel: **anomalous absorption of sound;**

B.O. Kerbikov and M.S. Lukashov, arXiv:1607.00125

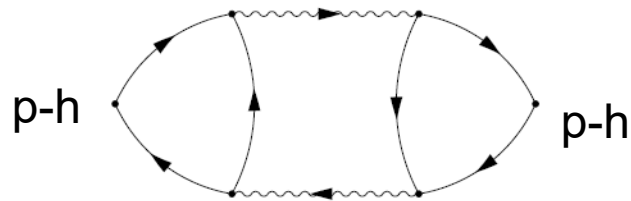


FIG. 2. Feynman diagram for the AL polarization operator for the sound absorption.

$$\text{Im} \Pi = -\omega g^2 \frac{m^4}{2^5 p_0^4 \pi^3} (v_0^2 + 1)^2 \ln^2 \frac{\omega_D}{2\pi T_c} \left(\frac{T_c}{T - T_c} \right)^{3/2}$$

➔ enhanced near CP!

Unfortunately,
a weak coupling calculation a la BCS theory.

Spectral function of density fluctuations

The density fluctuation depends on the transport as well as thermodynamic quantities which show an anomalous behavior around the critical point.

Especially, **the existence of the density-temperature coupling.**
Missing in the previous analyses.

For non-relativistic case with use of Navier-Stokes eq.
L.D. Landau and G.Placzek(1934),
L. P. Kadanoff and P.C. Martin(1963),
R. D. Mountain, Rev. Mod. Phys. 38 (1966), 38
H.E. Stanley, 'Intro. To Phase transitions and critical phenomena'
(Clarendon, 1971)

We apply for the first time relativistic hydrodynamic equations to analyze the spectral properties of density fluctuations, and examine possible critical phenomena.

Notice: The 1st-order can be valid for describing the hydrodynamic modes with a long wave-length without encountering the causality problem.

Relativistic Hydrodynamics

Y.Minami, TK, Prog.Theor.Phys. 122 (2010) 881-910 • e-Print: 0904.2270 [hep-th]

$$\left\{ \begin{array}{l} \partial_\mu T^{\mu\nu} = 0 \\ \partial_\mu N^\mu = 0 \end{array} \right. \quad \begin{array}{l} T^{\mu\nu} = (\epsilon + P)u^\mu u^\nu - P g^{\mu\nu} + \tau^{\mu\nu} \\ N^\mu = n u^\mu + \nu^\mu \end{array}$$

$\tau^{\mu\nu}, \nu^\mu$ dissipative terms

(1) Energy-frame

$$\begin{aligned} \tau^{\mu\nu} &= \eta \left[\partial_\perp^\mu u^\nu + \partial_\perp^\nu u^\mu - \frac{2}{3} \Delta^{\mu\nu} (\partial_\perp \cdot u) \right] + \zeta \Delta^{\mu\nu} (\partial_\perp \cdot u) \\ \nu^\mu &= \kappa \left(\frac{nT}{w} \right)^2 \partial_\perp^\mu \left(\frac{\mu}{T} \right) \quad u_\mu \nu^\mu = 0 \quad u_\mu \tau^{\mu\nu} = 0 \end{aligned}$$

(2) Particle frame ; Eckart(1940), unstable $u_\mu \tau^{\mu\nu} u_\nu = 0$

Tsumura-Kunihiro-Ohnishi, Phys.Lett.**B646**(2007);
Tsumura & T.K., PL**B668**(2008).

$$\nu^\mu = 0$$

$$\begin{aligned} \tau^{\mu\nu} &= \eta \left[\partial_\perp^\mu u^\nu + \partial_\perp^\nu u^\mu - \frac{2}{3} \Delta^{\mu\nu} (\partial_\perp \cdot u) \right] - \zeta' (3u^\mu u^\nu - \Delta^{\mu\nu}) (\partial_\perp \cdot u) \\ &\quad + \kappa (u^\mu \partial_\perp^\nu + u^\nu \partial_\perp^\mu) T \quad \tau^\mu_\mu = 0 \end{aligned}$$

(3) Israel-Stewart

Linear approximation around the thermal equilibrium;

$$u^\mu(\vec{r}, t) = u_0^\mu + \delta u^\mu(\vec{r}, t) \quad n(\vec{r}, t) = n_0 + \delta n(\vec{r}, t) \quad \text{etc}$$

In the rest frame of the fluid,

$$u_0^\mu = (1, \mathbf{0}) \quad \delta u^\mu(\vec{r}, t) = (0, \vec{v}(\vec{r}, t))$$

Inserting them into $T^{\mu\nu}$, N^μ , and taking the linear approx.

- Linearized Landau equation (Lin. Hydro in the energy frame);

$$\frac{\partial \delta n}{\partial t} + n_0 \nabla \cdot \vec{v} - \kappa \frac{n_0}{w_0} \left[\frac{T_0}{w_0} \nabla^2 (\delta P) - \nabla^2 (\delta T) \right] = 0$$

$$w_0 \frac{\partial \vec{v}}{\partial t} - \eta \nabla^2 \vec{v} - \left(\frac{1}{3} \eta + \zeta \right) \nabla (\nabla \cdot \vec{v}) + \nabla (\delta P) = 0$$

$$n_0 \frac{\partial \delta s}{\partial t} - \frac{\kappa}{T_0} \nabla^2 (\delta T) + \frac{\kappa}{w_0} \nabla^2 (\delta P) = 0$$

Rel. effects

with

$$\delta P(x) = \frac{w_0 c_s^2}{n_0 \gamma} \delta n(x) + \frac{w_0 c_s^2 \alpha_P}{\gamma} \delta T(x)$$

$$\delta s(x) = -\frac{w_0 c_s^2 \alpha_P}{n_0^2 \gamma} \delta n(x) + \frac{\tilde{c}_n}{T_0} \delta T(x)$$

Solving δn as an initial value problem using Laplace transformation, we obtain

$$S_{nn}(\vec{k}, \omega) = \langle \delta n(\vec{k}, \omega) \delta n(\vec{k}, t=0) \rangle, \text{ in terms of the initial correlation.}$$

Spectral function of density fluctuations in the Landau frame

Y.Minami, TK, Prog.Theor.Phys. 122 (2010) 881-910 • e-Print: 0904.2270 [hep-th]

In the long-wave length limit, $k \rightarrow 0$

$$\frac{S_{nn}(\vec{k}, \omega)}{\langle (\delta n(\vec{k}, t=0))^2 \rangle} = \left(1 - \frac{1}{\gamma}\right) \frac{2\Gamma_R k^2}{\omega^2 + \Gamma_R^2 k^4} + \frac{1}{\gamma} \left(\frac{\Gamma_B k^2}{(\omega + c_s k)^2 + \Gamma_B^2 k^4} + \frac{\Gamma_B k^2}{(\omega - c_s k)^2 + \Gamma_B^2 k^4} \right)$$

thermal mode sound modes

Rel. effects appear only in the width of the peaks.

$$\Gamma_R = \chi \quad \Gamma_B = \Gamma + \frac{1}{2} c_s^2 T_0 (\kappa / w_0 - 2\chi\alpha_p)$$

$$\Gamma = \frac{1}{2} [\chi(\gamma - 1) + \nu_l]$$

α_p rate of isothermal exp.

thermal expansion rate: $\chi = \frac{\kappa}{n_0 C_P}$

c_s : sound velocity γ : specific heat ratio Long. Dynamical :

$$\nu_l = \left(\zeta + \frac{4}{3}\eta \right) / w_0$$

enthalpy

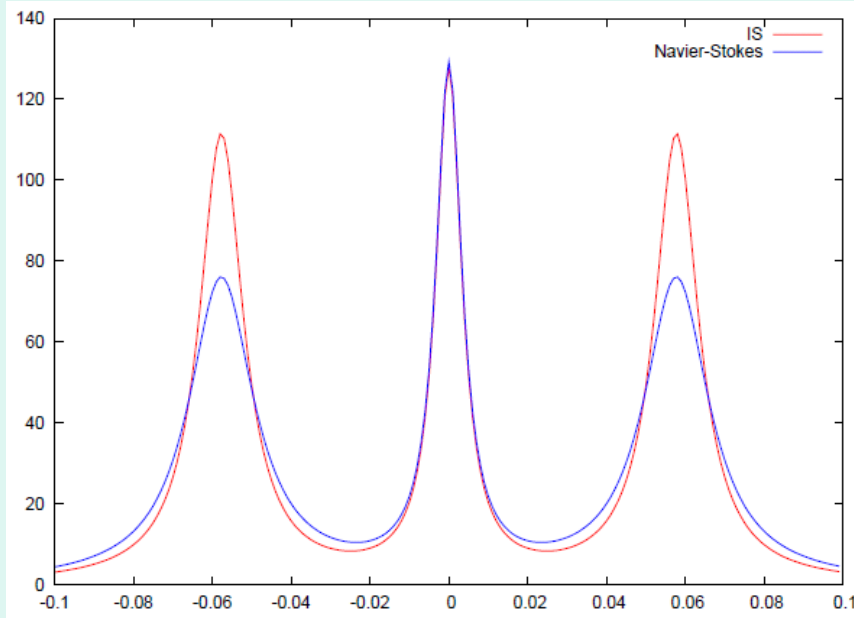
Notice: $\gamma = c_p / c_n = t^{-\gamma+\alpha} \rightarrow \infty$

As approaching the critical point, the ratio of specific heats diverges!

The strength of the sound modes vanishes out at the critical point.

Spectral function from I-S eq.

$$S_{nn}(\vec{k}, \omega) / \langle (\delta n(\vec{k}, t = 0))^2 \rangle$$



— I-S
— “Non-relativistic”

ω [1/fm]

$$\Gamma_B = \Gamma - c_s^2 T_0 (2\chi\alpha_P - \kappa/\omega_0)$$

For $\tau_\kappa > \frac{\kappa T_0}{\omega_0}$

$$\frac{S_{nn}(\vec{k}, \omega)}{\langle (\delta n(\vec{k}, t = 0))^2 \rangle} = \left(1 - \frac{1}{\gamma}\right) \frac{2\chi k^2}{\omega^2 + \chi^2 k^4} + \frac{1}{\gamma} \left(\frac{\Gamma_B k^2}{(\omega + c_s k)^2 + \Gamma_B^2 k^4} + \frac{\Gamma_B k^2}{(\omega - c_s k)^2 + \Gamma_B^2 k^4} \right)$$

$$+ O(k^2) \times \left[\frac{1/\tau_\zeta}{\omega^2 + (1/\tau_\zeta)^2} + \frac{1/\tau_\eta}{\omega^2 + (1/\tau_\eta)^2} + \frac{\omega_0 / (\tau_\kappa \omega_0 - \kappa T_0)}{\omega^2 + [\omega_0 / (\tau_\kappa \omega_0 - \kappa T_0)]^2} \right]$$



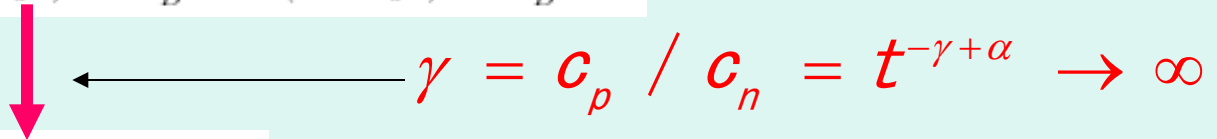
No contribution in the long-wave length limit $k \rightarrow 0$.

Conversely speaking, the first-order hydro. Equations have no problem to describe the hydrodynamic modes with long wave length, as it should.

Critical behavior of the density-fluctuation spectral functions

Y.Minami, TK, Prog.Theor.Phys. 122 (2010) 881-910 • e-Print: 0904.2270 [hep-th]

$$S_{nn}(\vec{k}, \omega) = \langle (n(\vec{k}, t=0))^2 \rangle \left[\left(1 - \frac{1}{\gamma}\right) \frac{2\Gamma_R k^2}{\omega^2 + \Gamma_R^2 k^4} + \frac{1}{\gamma} \left\{ \frac{\Gamma_B k^2}{(\omega - c_s k)^2 + \Gamma_B^2 k^4} + \frac{\Gamma_B k^2}{(\omega + c_s k)^2 + \Gamma_B^2 k^4} \right\} \right]$$


 $\gamma = c_p / c_n = t^{-\gamma+\alpha} \rightarrow \infty$

$$\langle (n(\vec{k}, t=0))^2 \rangle \frac{2\Gamma_R k^2}{\omega^2 + \Gamma_R^2 k^4}$$

$$\langle (n(\vec{k}, t=0))^2 \rangle \sim (|T - T_c| + k^2)^{-1}$$

Critical opalescence

In the vicinity of CP, only the Rayleigh peak stay out, while the sound modes (Brillouin peaks) die out.

c.f. The dynamical critical exponent.

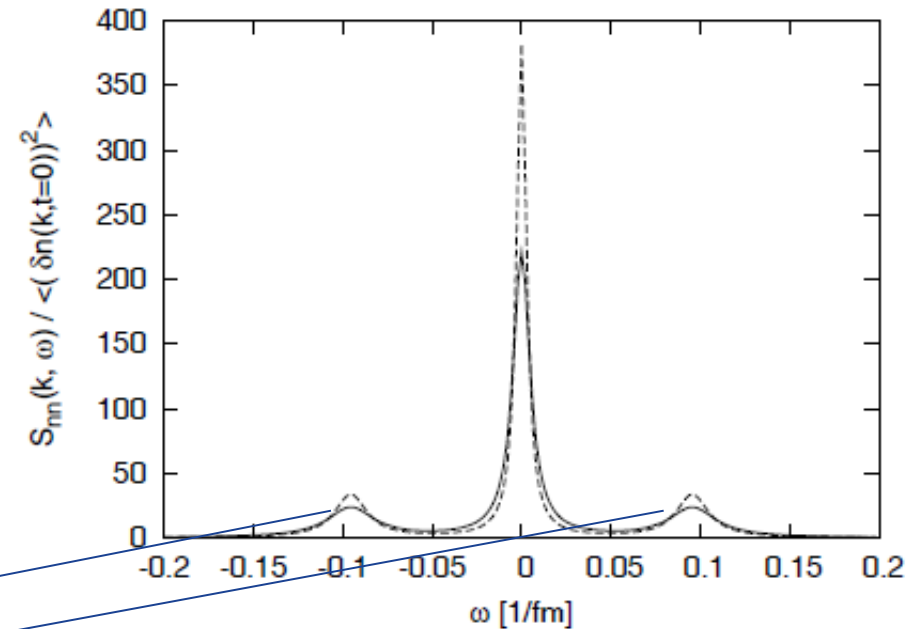
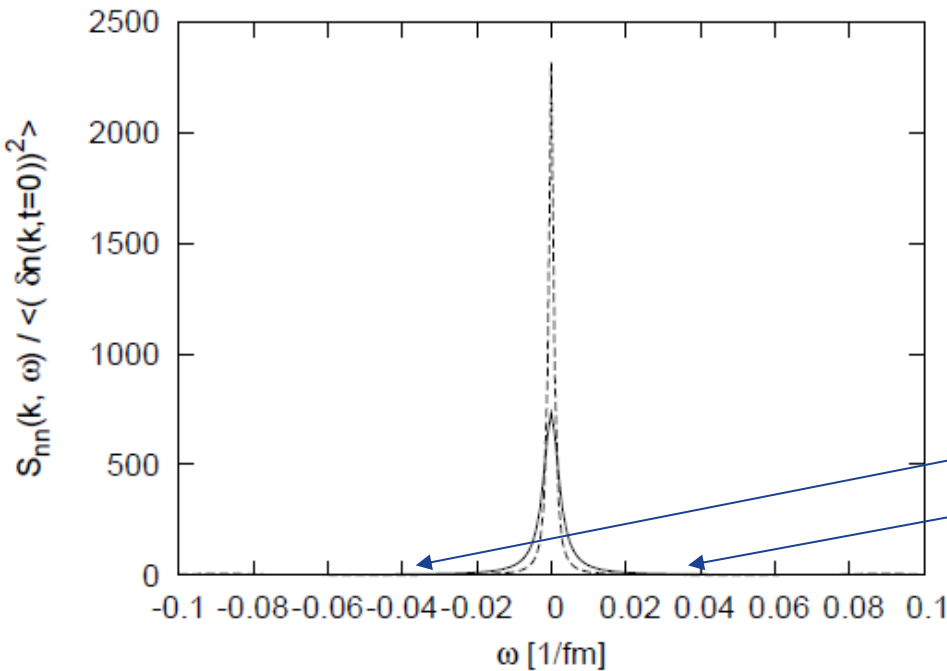
$$z = 2 + (\gamma - a) / \nu \square 3$$

So, the divergence of Γ_B and the viscosities therein can not be observed, unfortunately.

Spectral function of density fluctuation at CP

0.4

$t = 0.1$

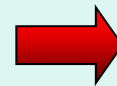


$$\tau \equiv (v_s - v_c) / v_c =$$

**The sound mode (Brillouin) disappears
Only an enhanced thermal mode remains.
Furthermore, the Rayleigh peak is
enhanced, meaning the large energy
dissipation.**

Spectral function at CP

The soft mode around QCD CP is thermally induced density fluctuations, but not the usual sound mode.



Suggesting interesting critical phenomena related to sound mode.

Why at all do sound modes die out at the Critical Point ?

The correlation length

$$\xi = \xi_0 t^{-\nu}$$

The wave length of sound mode

$$\lambda_s$$

The hydrodynamic regime!

$$\xi \ll \lambda_s$$

However, around the critical point

$$\xi \longrightarrow \infty \quad \text{as} \quad t \longrightarrow 0$$

$$\lambda_s \ll \xi$$

The system behaves as an aggregate of clumps of matter with a diameter ξ .

So the hydrodynamic sound modes can not be developed around CP!